

Lunar Detection of UHE Cosmic Particles with LOFAR

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What is it about ?

1. Detecting UHECRs and neutrinos with Radio Antennas (LOFAR HBA) from the near surface Moon(Lunar Askaryan Technique)

Why the Moon ?

- 1. Moon lack atmosphere (Askaryan '62 and Dagkesamanski et al '89)
- 2. Good radio properties
- 3. Large detector area/volume (~1700km radius)

Results from previous studies and what we are doing differently

Zero/Null detection so far (>1GHz observation except WRST)
Low Frequency (110-190Mz bandwidth with 200MHz sampling rate)
Beamforming for improved sensitivity (5/24 core stations due to limited bandwidth)





What is the Motivation Here ?

1. Understand the cut-off

2. Put an upper limit on UHECRs or neutrino flux (no detection) to constrain theoretical (top-down) models.

NB:

- 1. At low frequency, the whole lunar surface forms part of detector volume
- 2. Each station beam is larger the Moon (0.5deg)





Lunar detection of UHE cosmic particle





Simulated sensitivity of LOFAR

1. Null result requires upper limit.

2. Effective aperture (area * solid angle through which neutrinos is detectable)

3. MC-code for detailed calculation of isotropic aperture for neutrinos (Crs is still in progress)









Assumed a smooth lunar surface

* Surface roughness included, 5 stations and 100 hours of livetime

* work in progress for cosmic rays.