

FEEBLY INTERACTING DARK MATTER

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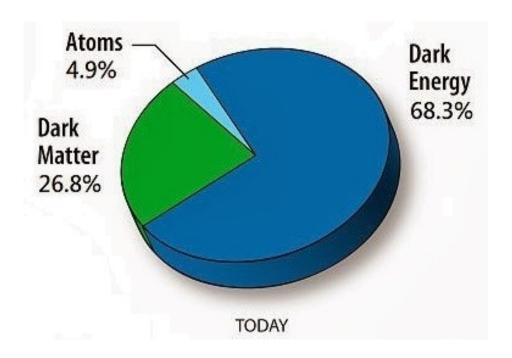
HEP@VUB PhD Event - 19/11/2020





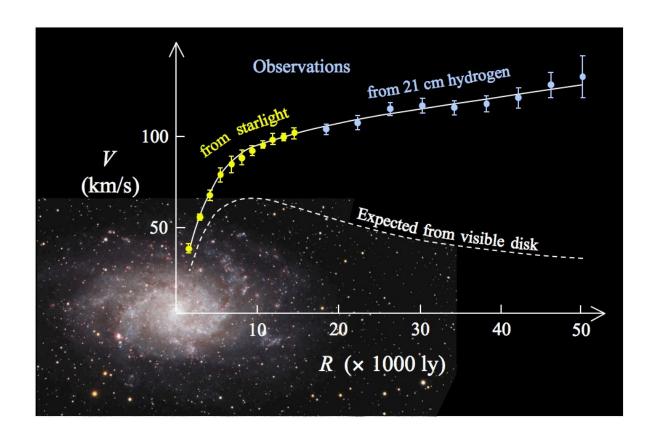


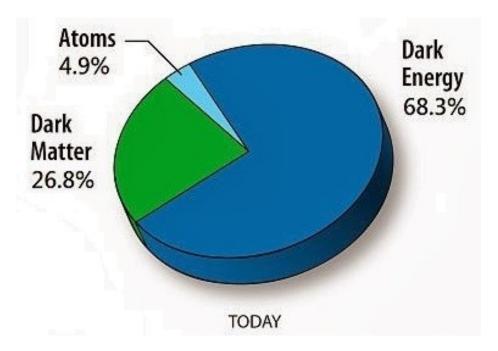
- Five times more dark than visible matter in our universe
- Compelling amount of evidence





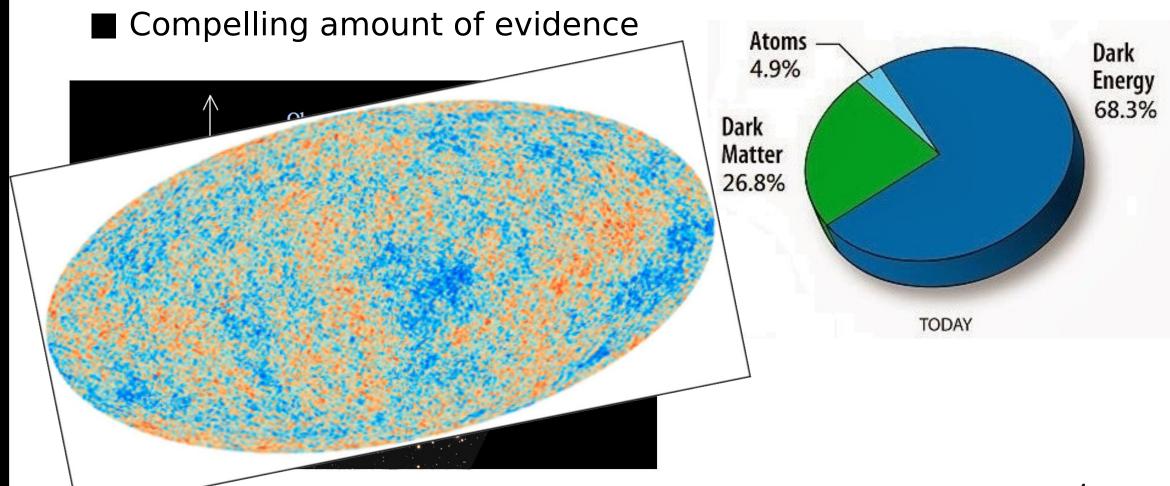
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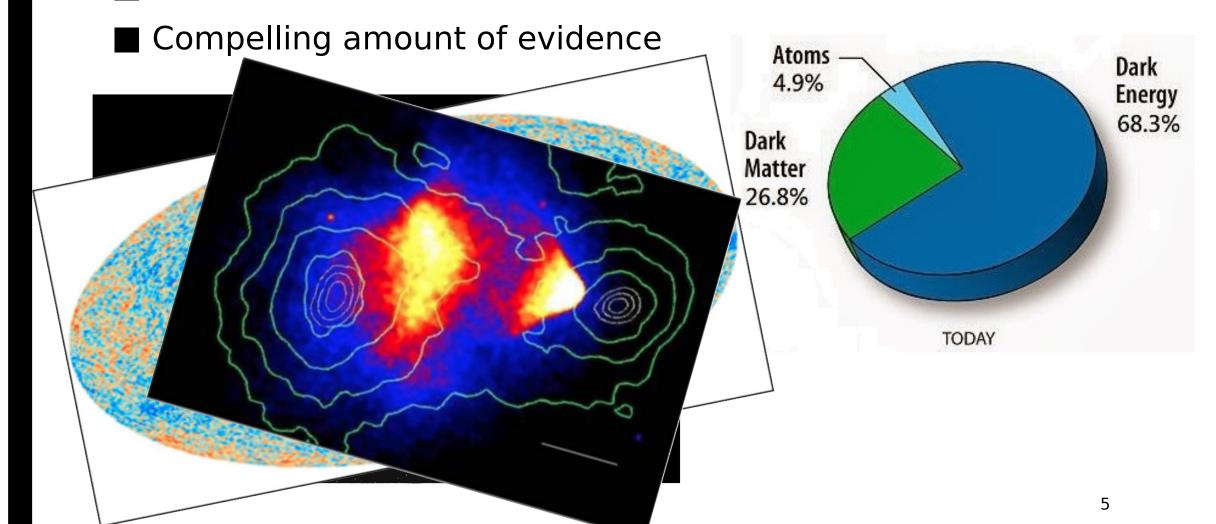


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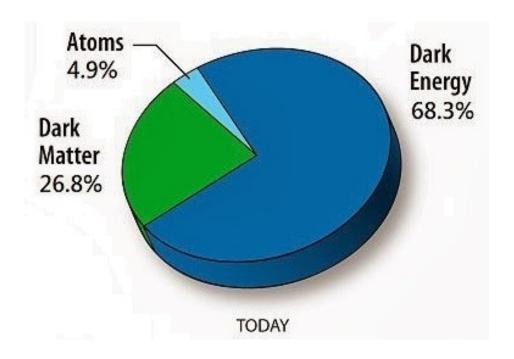




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Only Gravitational evidence

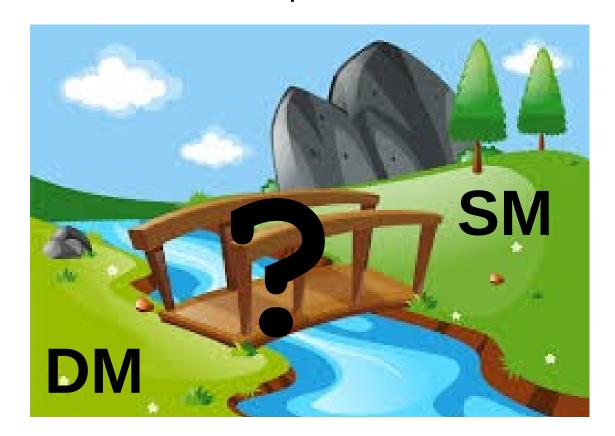
No info about DM nature





Open questions

■ How does dark matter couples to the SM?





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- How is it produced in the early universe?





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- How does dark matter couples to the SM?
- How is it produced in the early universe?
- How can we detect it?

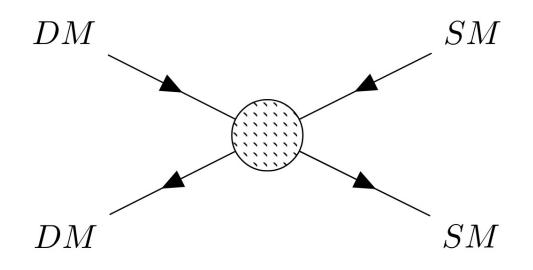


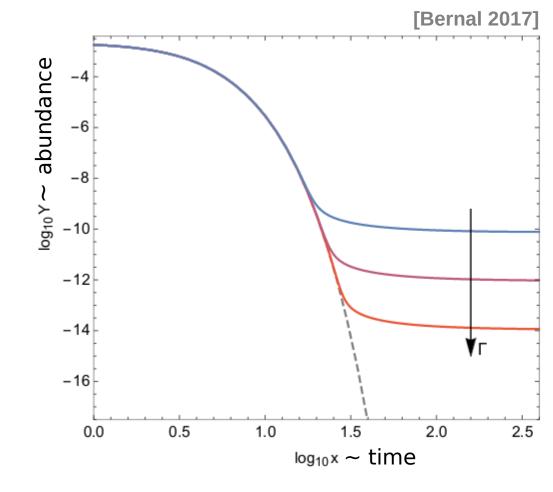


Weakly Interacting Massive Particle

[Lee & Weinberg 1977]

- DM annihilations happening fast
- DM in chemical equilibrium
- Freeze-out when expansion rate of the universe catches up







Direct detection constraints

Same process setting relic abundance responsible for DD signal

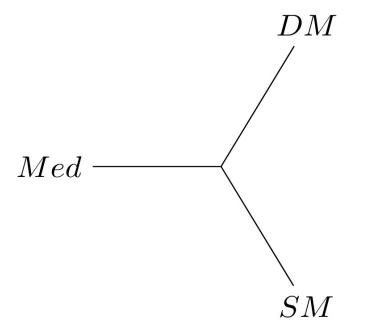
Strong constraints from experiments [Xenon Coll. 2018] DMDMWIMP-nucleon σ_{SI} [cm²] SMSMWIMP mass [GeV/c²]

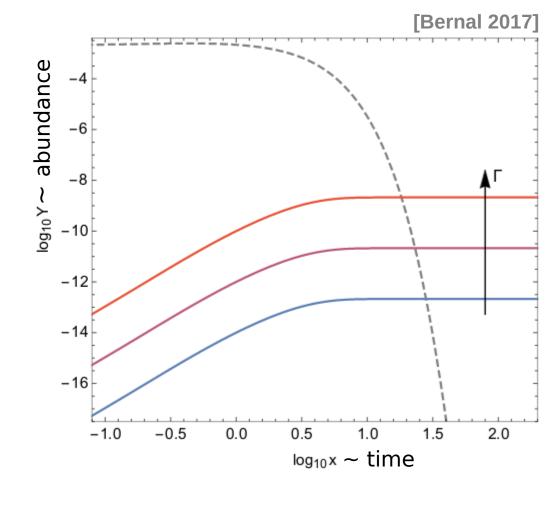


Feeble Interacting Massive Particle

[Hall 2009]

- Dark matter interacting feeble with the SM
- Annihilations are never fast
- Never in equilibrium
- Gradual production → freeze-in





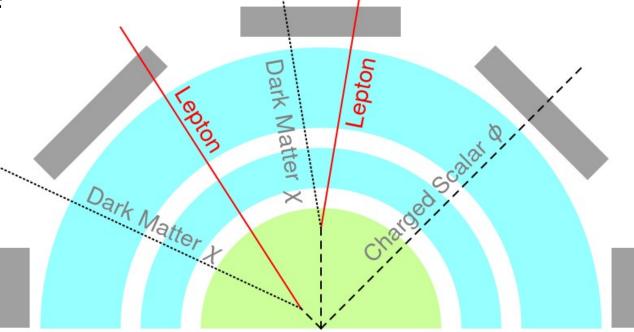


Long Lived Particles

- Mediator is charged under the SM
- Can be produced in pp collision
- Feeble interactions gives rise to LLPs

Strong link between lifetime and production rate

→ Very predictive





Summary

- Nature of DM still unknown
- WIMP produced through freeze-out mechanism
- Strong constraints from direct detection experiments
- Alternative: FIMP produced through freeze-in
- Gives rise to Long Lived Particles signatures at the LHC