ASTEROSEISMOLOGY: THE ROYAL ROAD TOWARDS THE INTERNAL PHYSICS OF STARS

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MAMSIE WANSIE





Take-home message



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Seismic waves offer in-situ measurements of internal stellar physics: new look upon stars



KULEUVEN Stellar interiors: poorly known

Rotation? Convection? Mixing? Magnetism?



KULEUVEN Starquakes as Modern Tool

Stellar evolution is dictated by stellar interior

Asteroseismology to the rescue: requires long-term uninterrupted high-precision data



KULEUVEN Stellar oscillations probe stellar interiors





KULEUVEN Space Asteroseismology Revolution



KU LEUVEN Ingredients: temporal/spatial

- Periodic perturbations of equations of physics:
 oscillations are the eigenmodes of the star
- Each mode described by spherical harmonic & frequency:
- Dominance of acting forces?
 - 1. pressure (acoustic waves)
 - 2. buoyancy (gravity waves)
 - 3. Coriolis (inertial waves)
 - 4. Lorentz (Alfvén waves)
 - 5. tidal (tidal waves)



Data-driven modelling





Theoretical predictions



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KU LEUVEN Aims of Asteroseismology



SOME APPLICATIONS:

1) WEIGHING, SIZING, AGEING OF STARS

2) INTERNAL ROTATION OF STARS

KULEUVEN Helioseismology paved the way



(Christensen-Dalsgaard, 2002, Reviews of Modern Physics)

KULEUVEN Red Giants: successors of Suns

Courtesy: NASA Kepler



(Hekker & Christensen-Dalsgaard, 2017, Astronomy & Astrophysics Review)

Sizing stars to 1% precision



Weighing/ageing stars to 2/10% precision



KU LEUVEN R, M, age for Exoplanet Research

Asteroseismology of Host Star: factor ~2 improvement for exoplanet radius + age delivery!



Huber et al. (2013) Van Eylen et al. (2014, 2018), Campante et al. (2016), Chontos et al. (2019)

KULEUVEN Ages for Galactic Archaeology

Seismic mass, radius, age of red giants from scaling relations



(Silva Aguirre et al. 2012, Miglio et al. 2013, Stello et al. 2015, Huber et al. 2017, Hon et al. 2019, Bellinger et al. 2019, Sharma et al. 2019, Jie Yu et al. 2020,...)

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KU LEUVEN Chemical evolution?



nuclear burning

macroscopic physics: element transport, e.g. rotation, waves, magnetism,...

KULEUVEN Rotational splitting of modes



Star more massive than Sun: Kurtz et al. (2014)

KU LEUVEN Internal Stellar Rotation



Beck et al. (2012, Nature)

KU LEUVEN "Watching" stars grow old



Successors of sun-like stars: Deheuvels et al. (2014)

KULEUVEN Asteroseismic estimates of Ω_{core}



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KU LEUVEN Measuring Ω_{core} versus Ω_{env}



"Standard SSE" needs fixes... (figure from Aerts, 2021, RMP, 111 stars)



Stars with convective core rotate quasi-rigidly

AM transport to keep ~rigid rotation & agree with AM of WDs

Magnetism/Tayler Instability:

Fuller et al. (2019), Takahashi & Langer (2020) Loi et al. (2019, 2020), Bugnet et al. (2021)

and/or

IGWs: Rogers (2015); Edelmann et al. (2019); Horst et al. (2020)

KULEUVEN Ongoing TESS/Gaia/Spectroscopic Surveys



KULEUVEN Onward to PLATO (2026+)

8% Data Rate is Guest Observer program via open ESA calls, incl. ToO option: welcome!





Much more to it: tidal, magneto-, pre-MS, nonlinear, GIW, ... asteroseismology Aerts, 2021, RMP: https://arxiv.org/abs/1912.12300 general introduction & update for non-expert