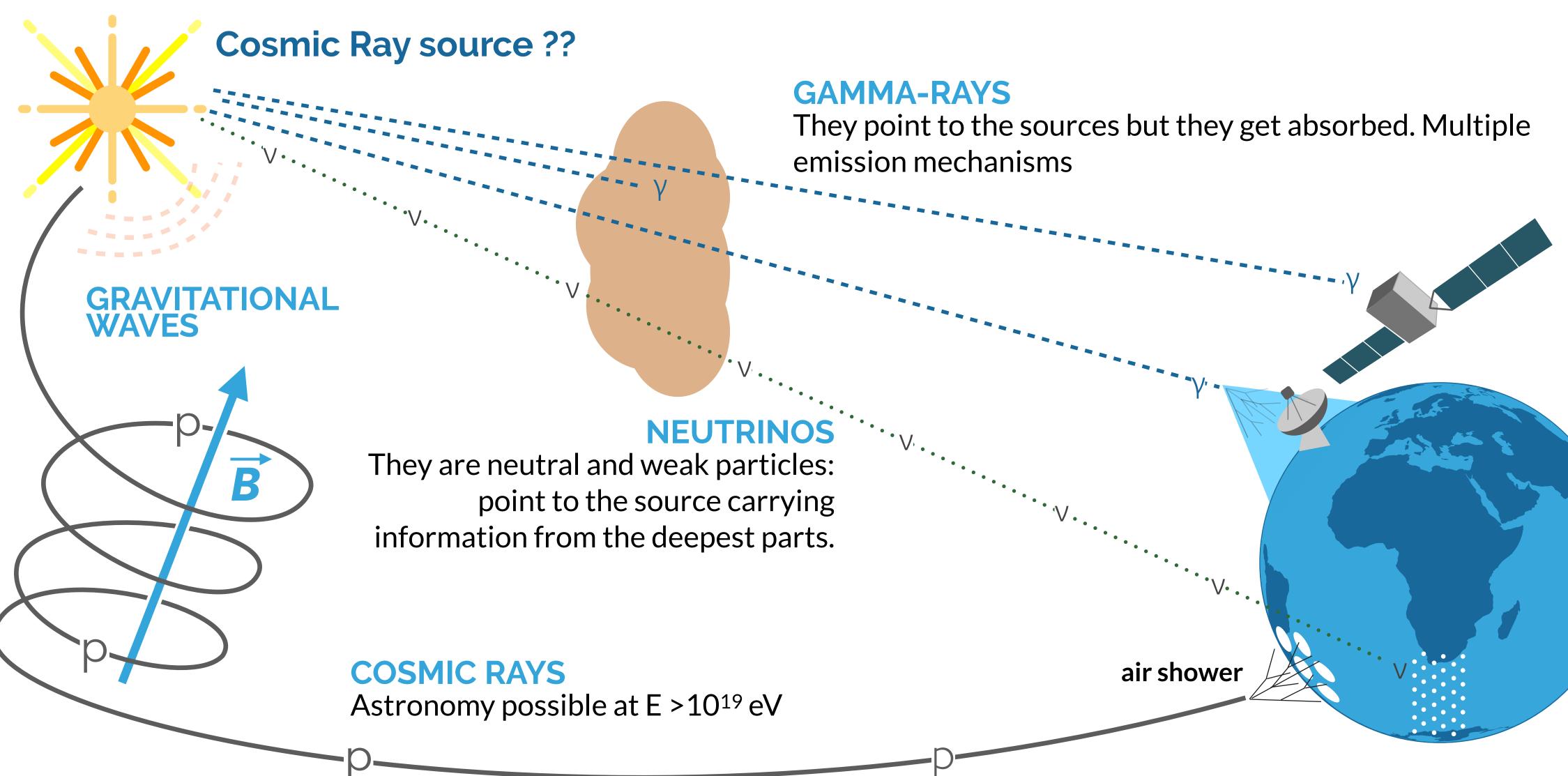
# IceCube and Beyond **IIHE Annual Meeting**

J. A. Aguilar on behalf of the IIHE IceCube group



# **Multimessenger Astronomy**



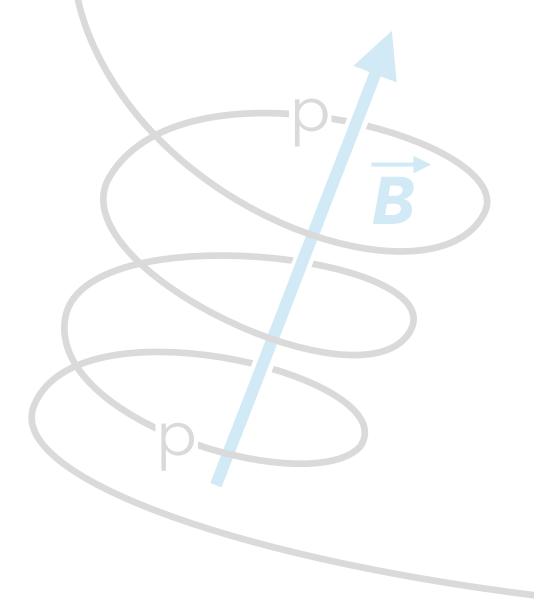




# **Neutrino Astronomy**

#### **Cosmic Ray source ??**

### ATIONAL

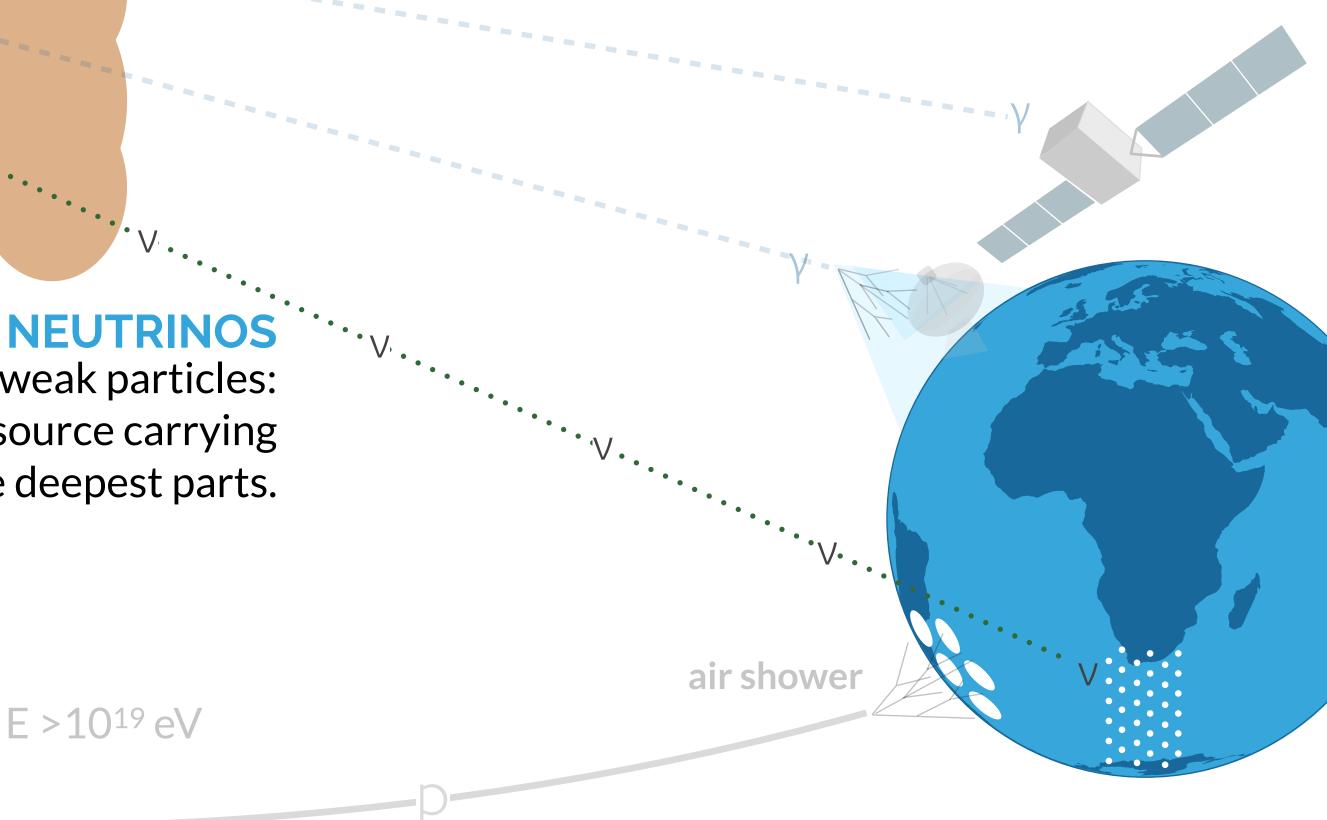


They are neutral and weak particles: point to the source carrying information from the deepest parts.

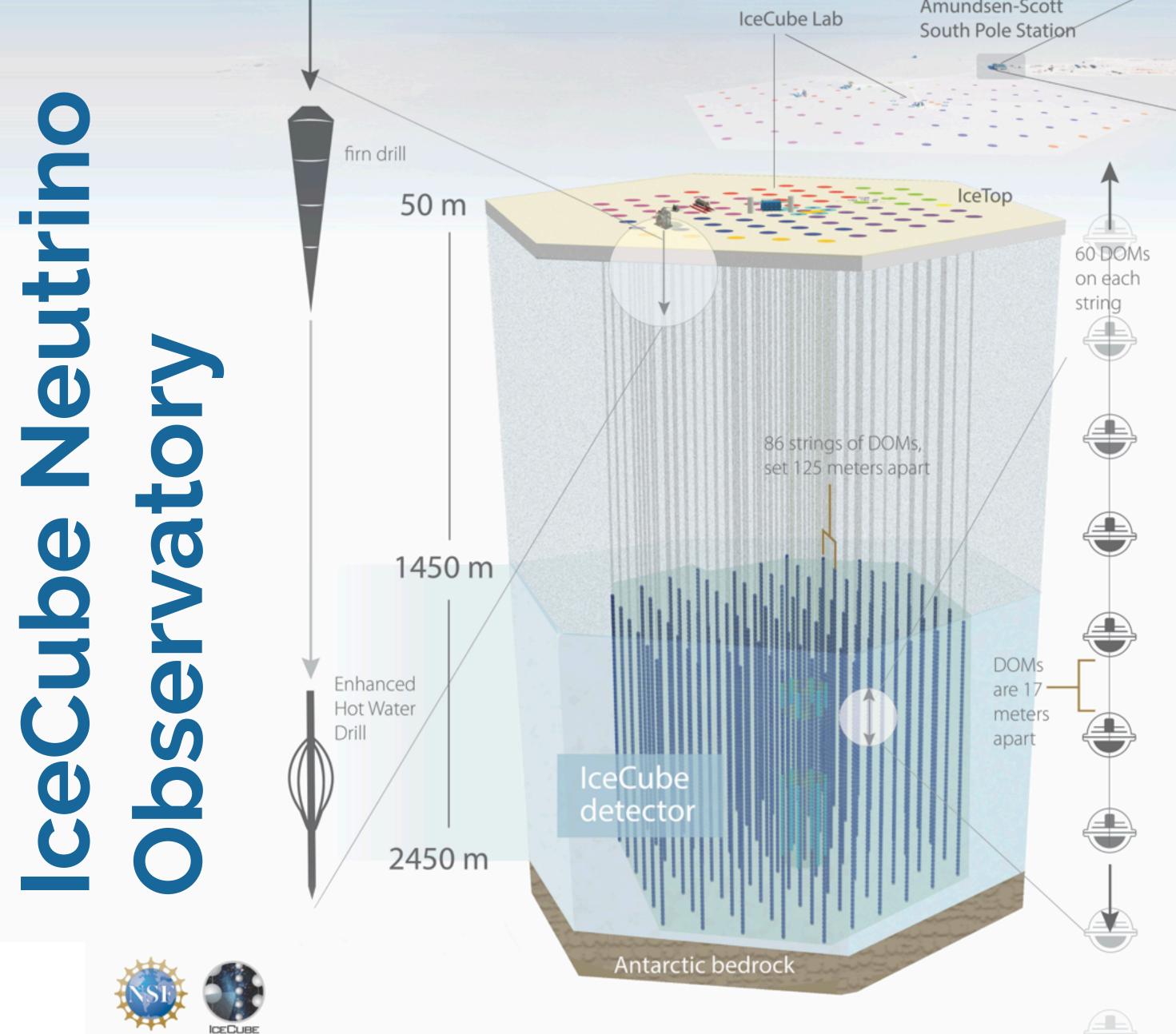
#### **COSMIC RAYS** Astronomy possible at E >10<sup>19</sup> eV

#### GAMMA-RAYS

They point to the sources but they get absorbed. Multiple emission mechanisms





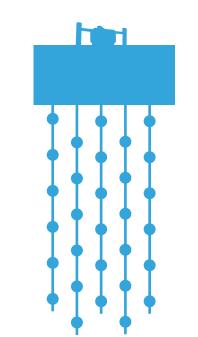


Amundsen-Scott



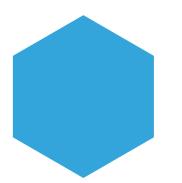


5,160 Digital Optical Modules (DOMs)



**86 string with 60 DOMs** each

**6** denser strings called DeepCore



**1** km<sup>2</sup> surface array with 324 DOMs: IceTop



**Completion in December** 2010

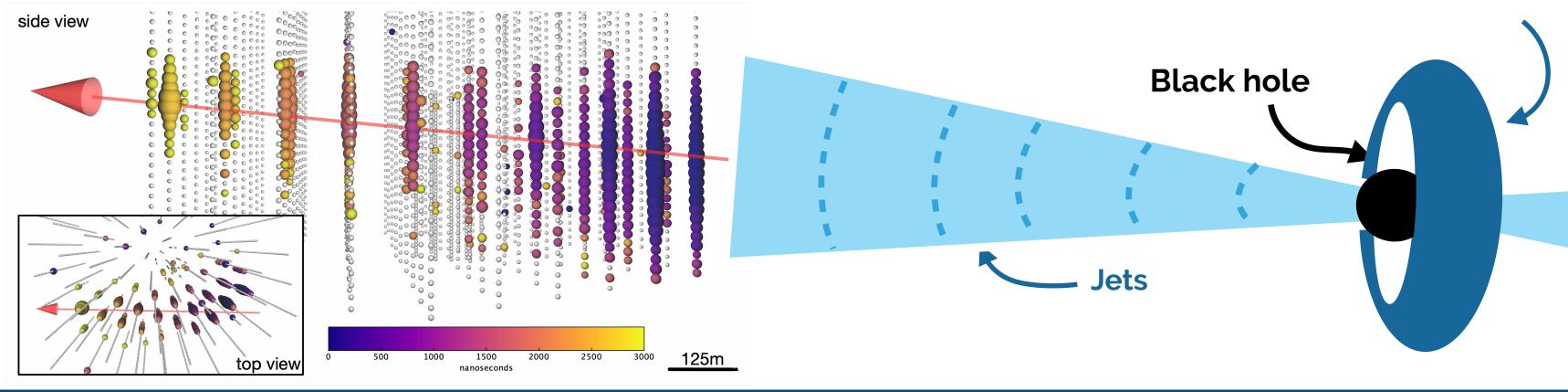


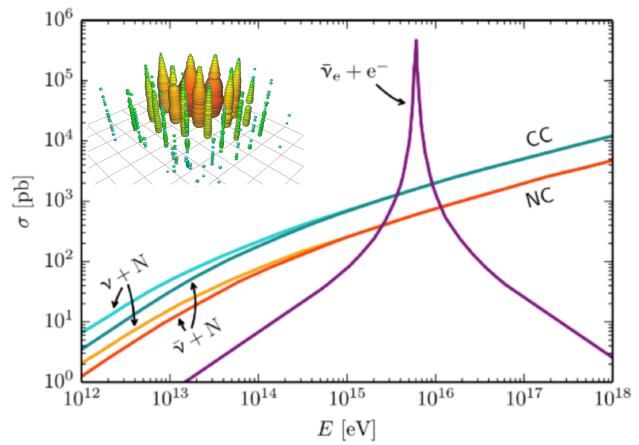




### 2017 TXS 0506+056

Coincidence of a neutrino event with a flare of gamma-rays from a Blazar



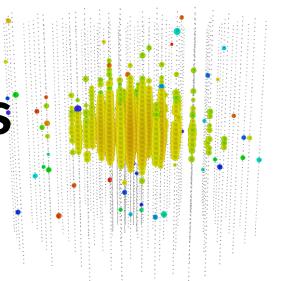


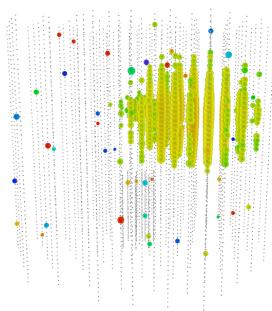
### 2021 Glashow **Resonance Event**

Observation of neutrino event at the Glashow energies (first  $\overline{\nu}_{\rho}$ )

### 2013 **Astrophysical Neutrinos**

Detection of the first flux of astrophysical neutrinos

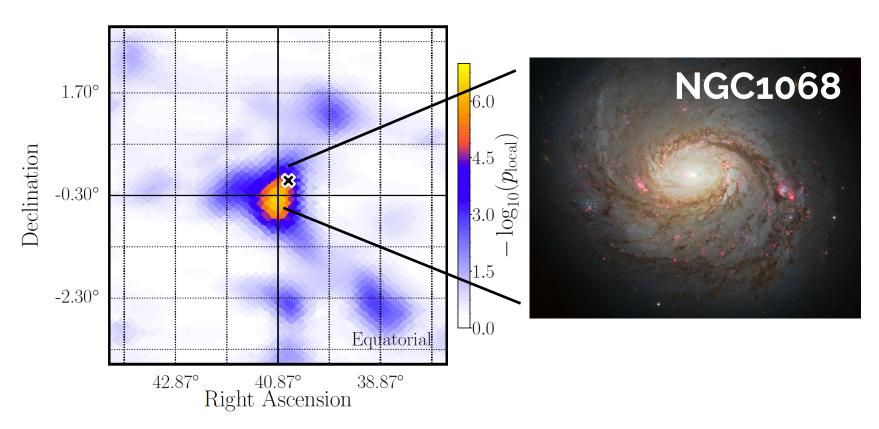




**Accretion Disk** 

### 2021 **M77**

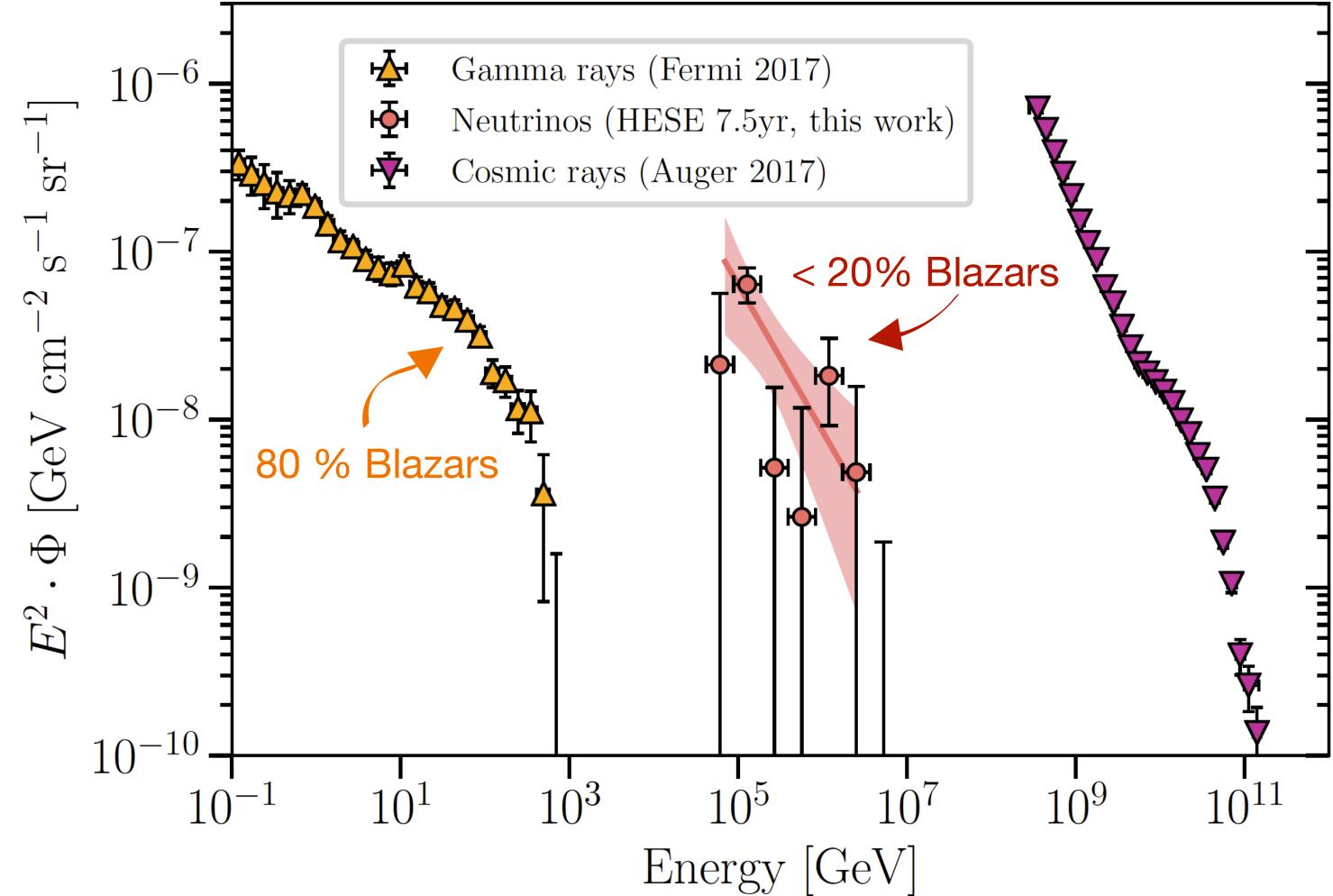
First hint of a point source in 10 years of data ( $2.9\sigma$ ). Improved point source analysis on the way







### **Astrophysical Neutrinos** State-of-art



Annual N.

- Spectral index of astro. flux:  $\gamma = 2.3 - 2.9$ depends on analysis / energy range
- Similar energies among messengers ... but also evidence for different origin!
- Gamma-obscured sources?

### **Astrophysical Neutrinos @ IIHE** Search for High-Energy Neutrinos from Ultra-Luminous Infrared Galaxies

### • ULIRGS: $L_{IR} \ge 10^{12} L_{\odot} (8 - 1000 \ \mu \text{m})$

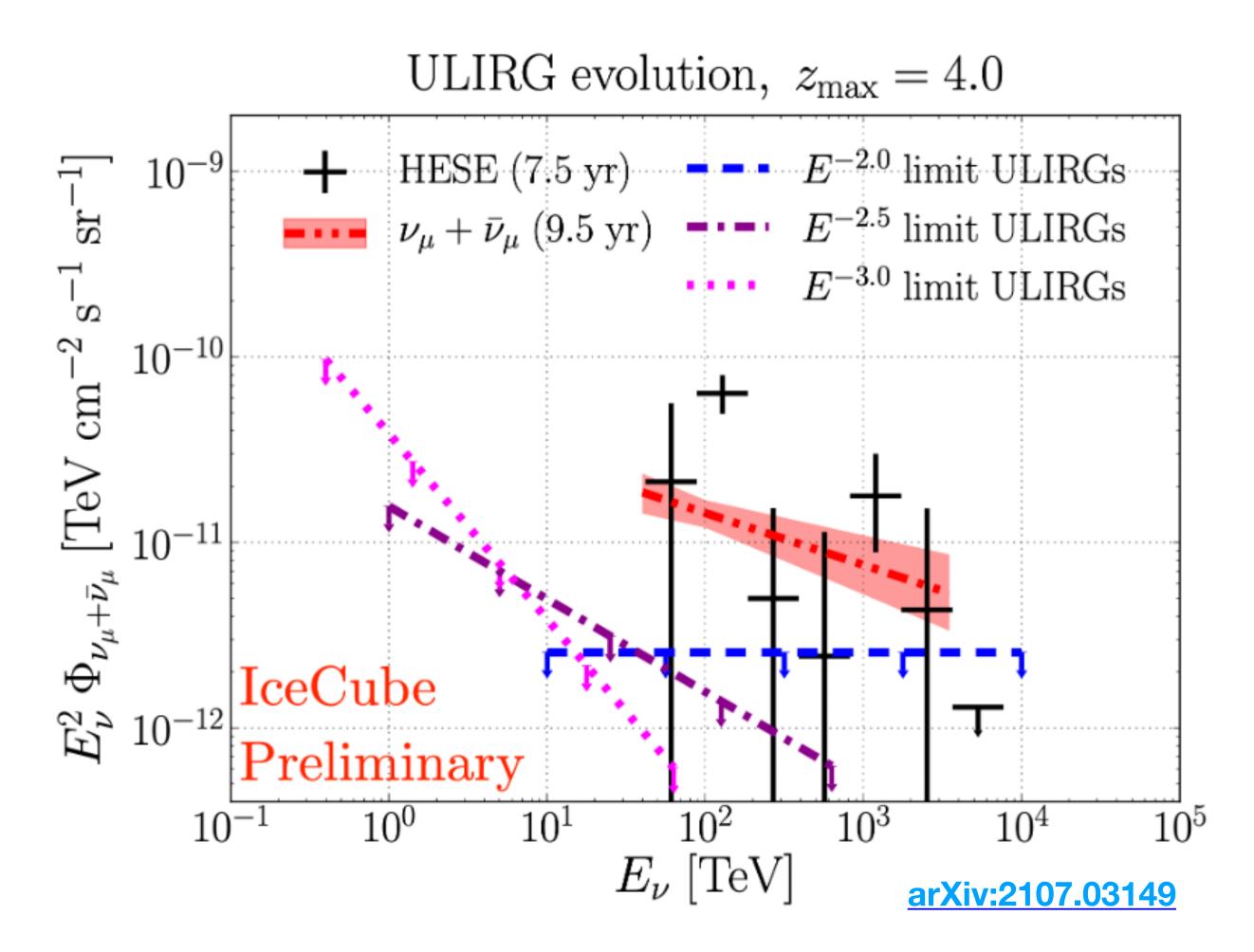
- Powered by starburst/AGN
- Candidate neutrino sources

#### • Stacking analysis

- 75 local ULIRGs (z  $\leq$  0.13)
- 7.5 years of data

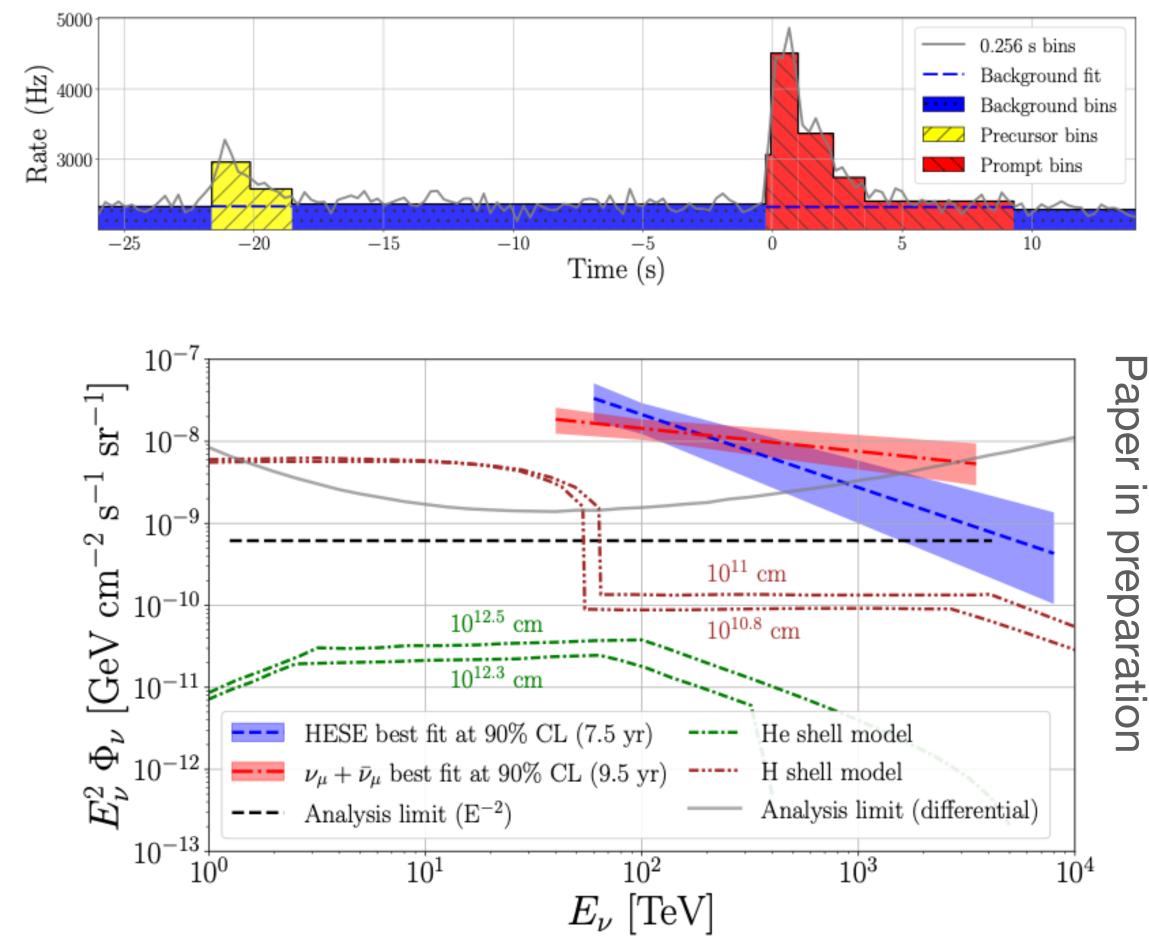
#### • No neutrinos found

- Set upper limits
- Constrained diffuse contribution of ULIRGs
- Constrained model predictions

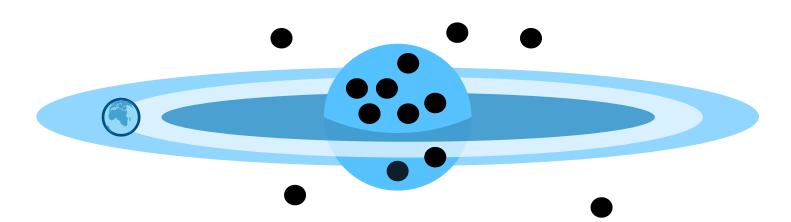


### **Astrophysical Neutrinos @ IIHE** Neutrinos from gamma-ray burst precursors

- Transient events ofter a unique discovery opportunity.
- Analysed light curves of 2684 bursts from Fermi-GBM
  - 10% shows signs of precursor emission
  - New temporal features identified!
  - Published in PRD: arXiv:2004.03246
- Performed 2 searches with IceCube to look for coincident neutrinos
  - No significant coincidences observed
  - Able to limit model predictions!



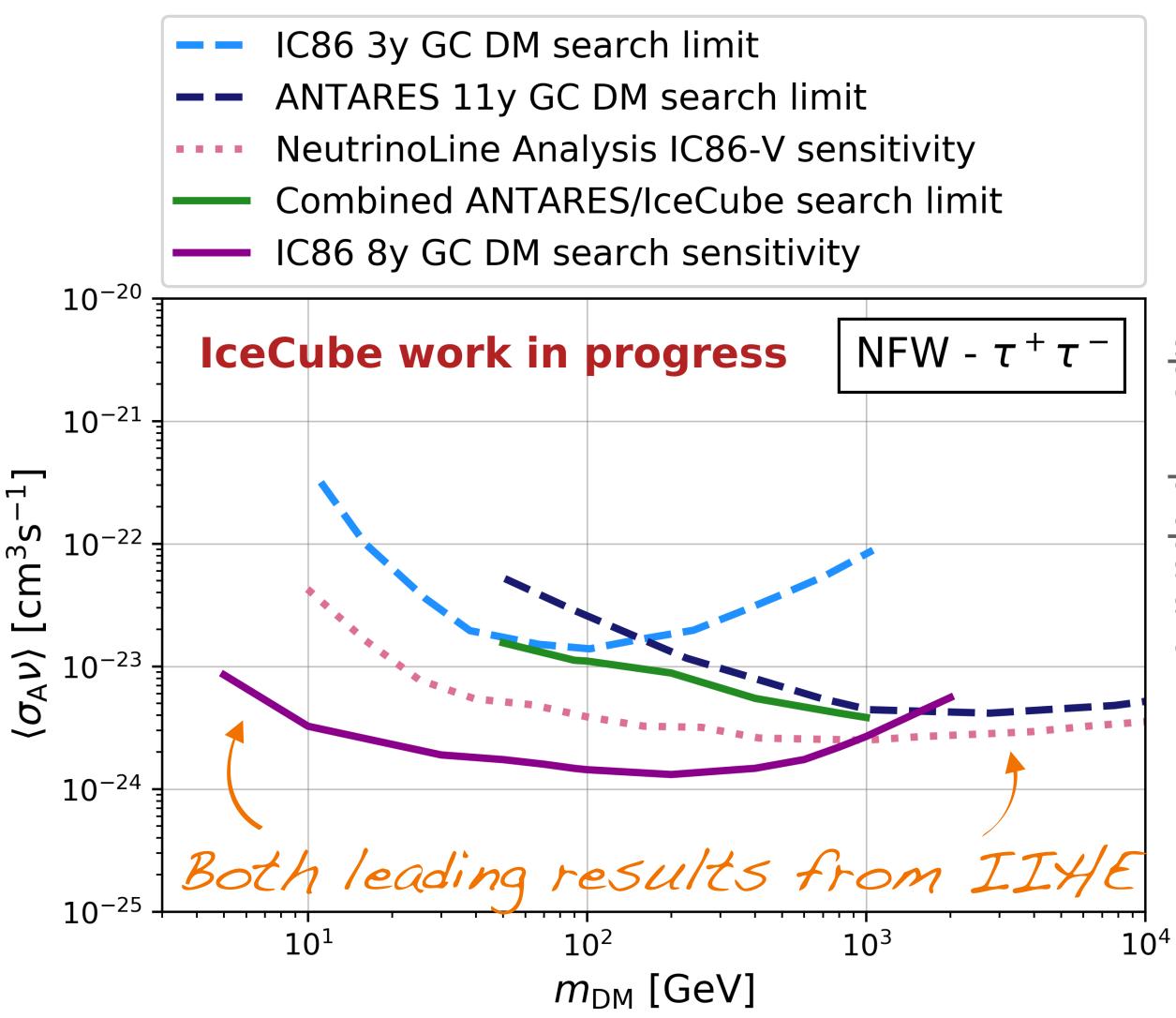
### **Particle Physics @ IIHE Dark Matter Searches**



- Use neutrinos to search for annihilation/decay signatures of Dark Matter
  - Combined analysis with ANTARES using published data. [PhysRevD.102.082002]
- Performed 2 additional analyses:
  - 1. First analysis using energy and the neutrino spectra (Neutrino Lines) with the Service de Theory.

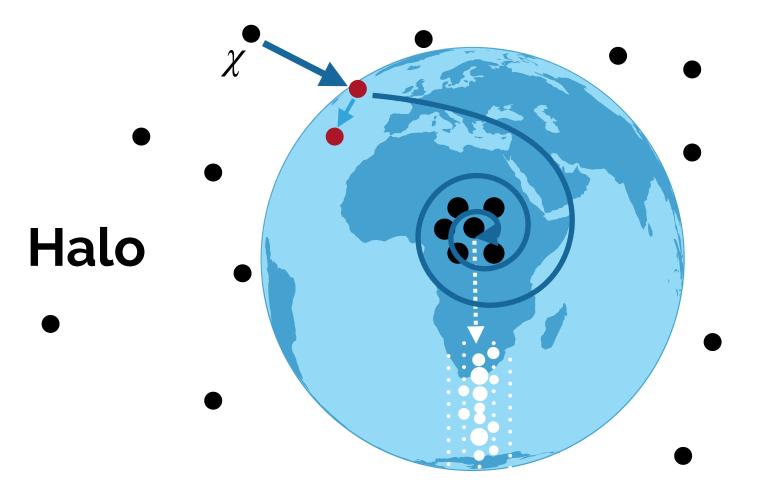
2. Extending towards lower masses with Deep Core. Best limits in neutrino channel in the whole mass range.

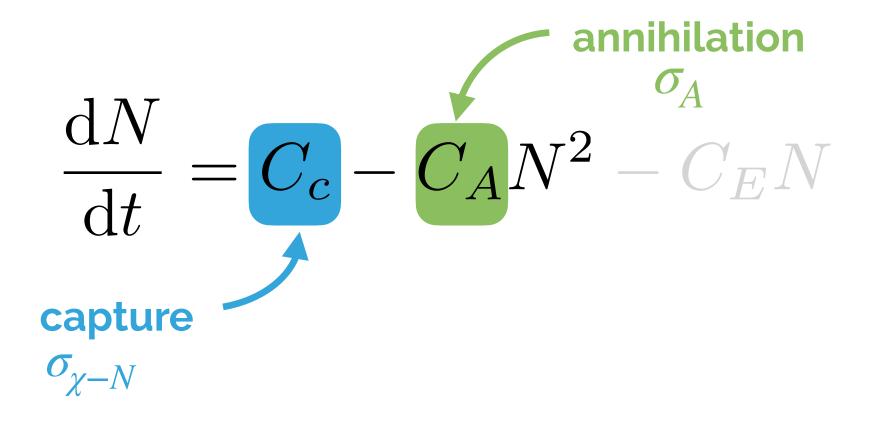




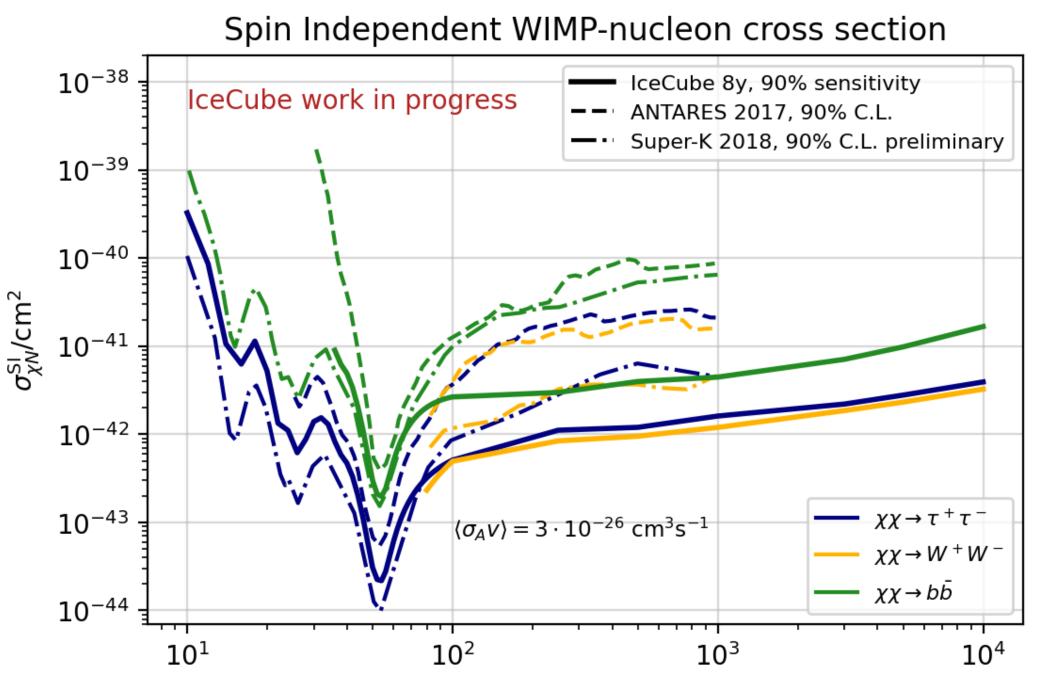


### **Particle Physics @ IIHE Dark Matter from the Center of the Earth**



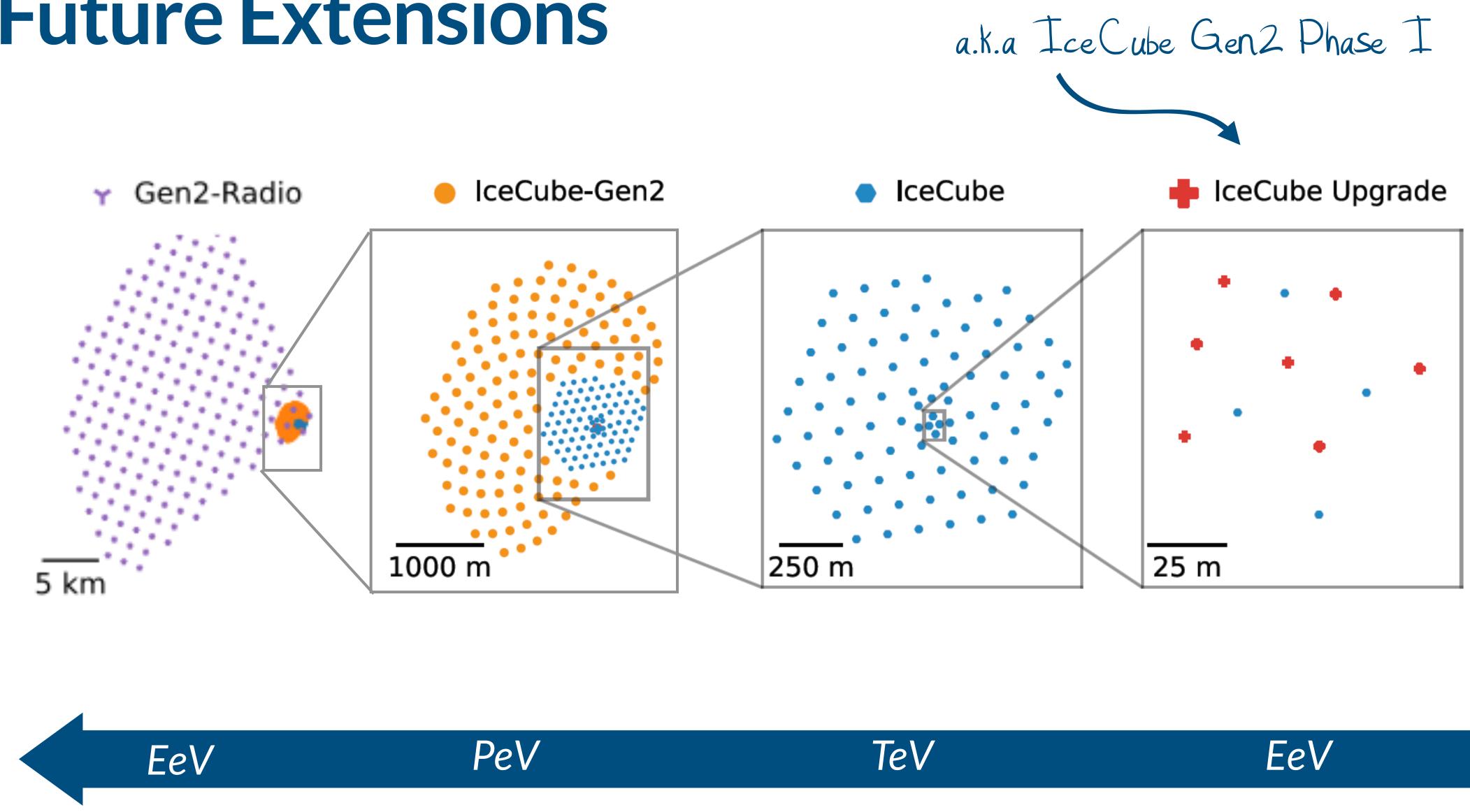


- source.
- We can relate the  $\sigma_{\!A}$  and  $\sigma_{\!\chi-N}$
- IceCube has the best sensitivity above 100 GeV
- Analysis recently unblinded (no results public yet).



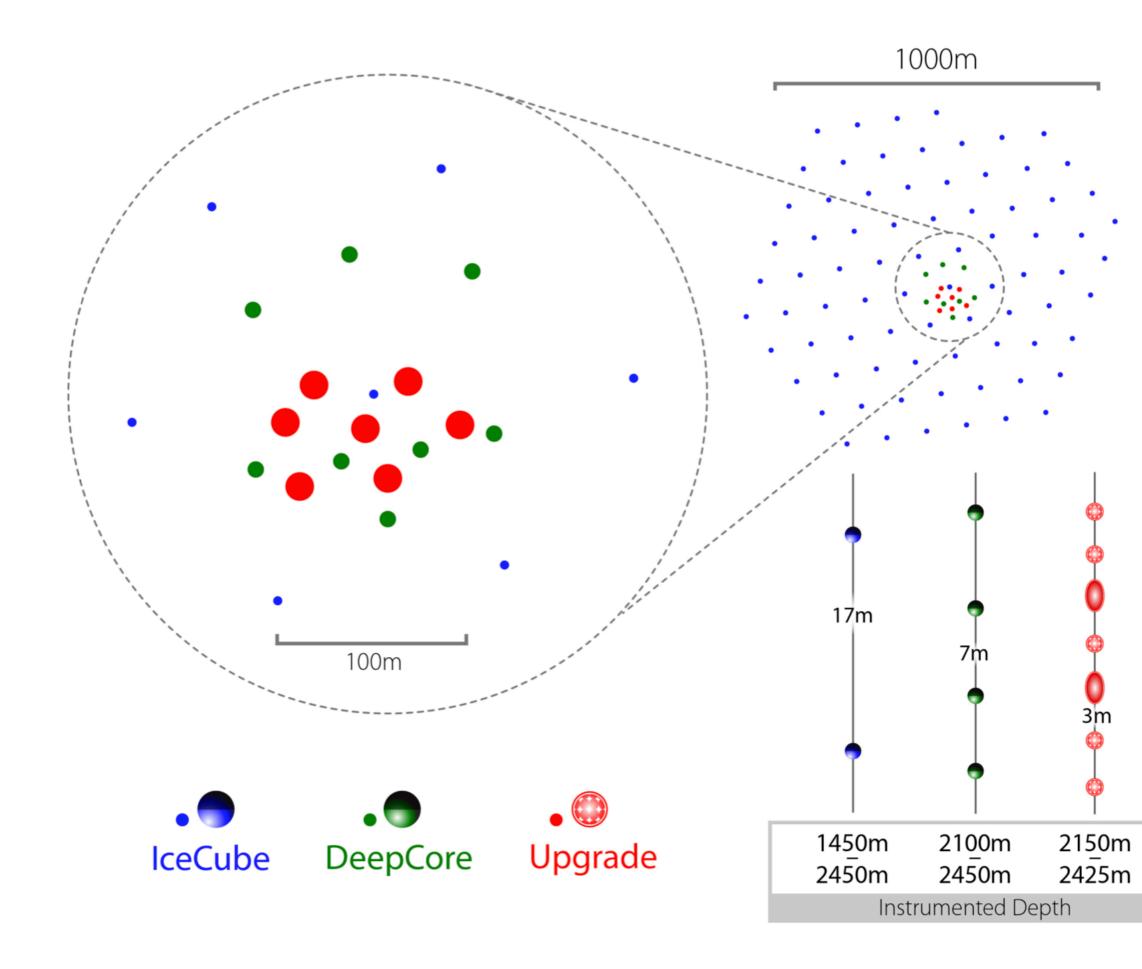
• Signal from the Earth cannot be  $m_{i}^{M/GeV}$  interpreted as an astrophysical

# **Future Extensions**





### IceCube-Upgrade IceCube-Gen2 Phase I



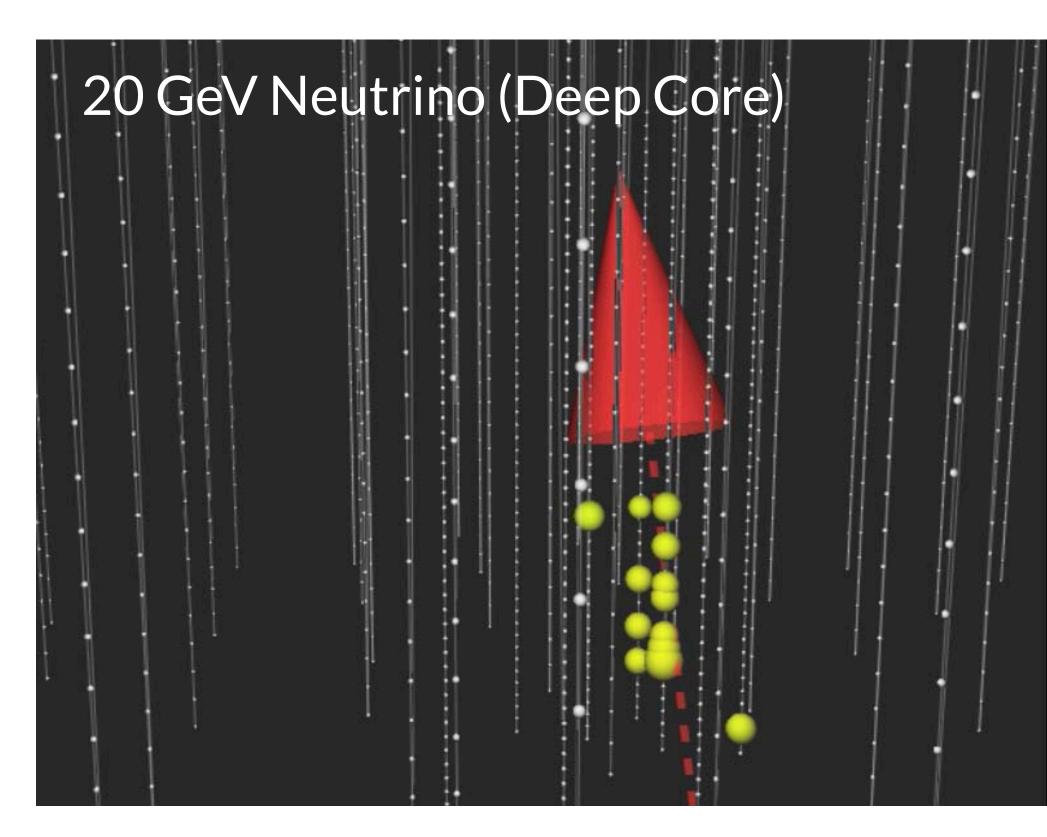


- Improved angular and energy reconstructions.
- Precision measurement of atmospheric neutrino oscillations.
- Construction scheduled for 2022, delayed because COVID-19, rescheduling undergoing (1-2 years)

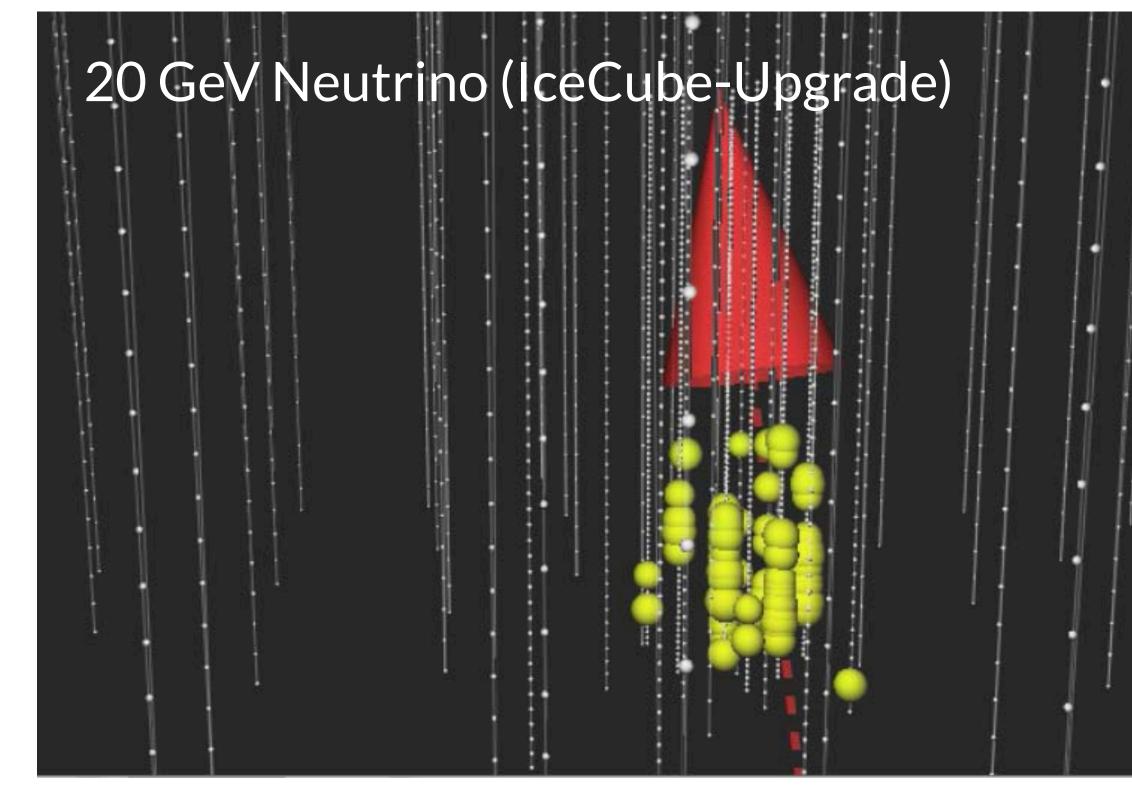




### **IceCube-Upgrade** Performance



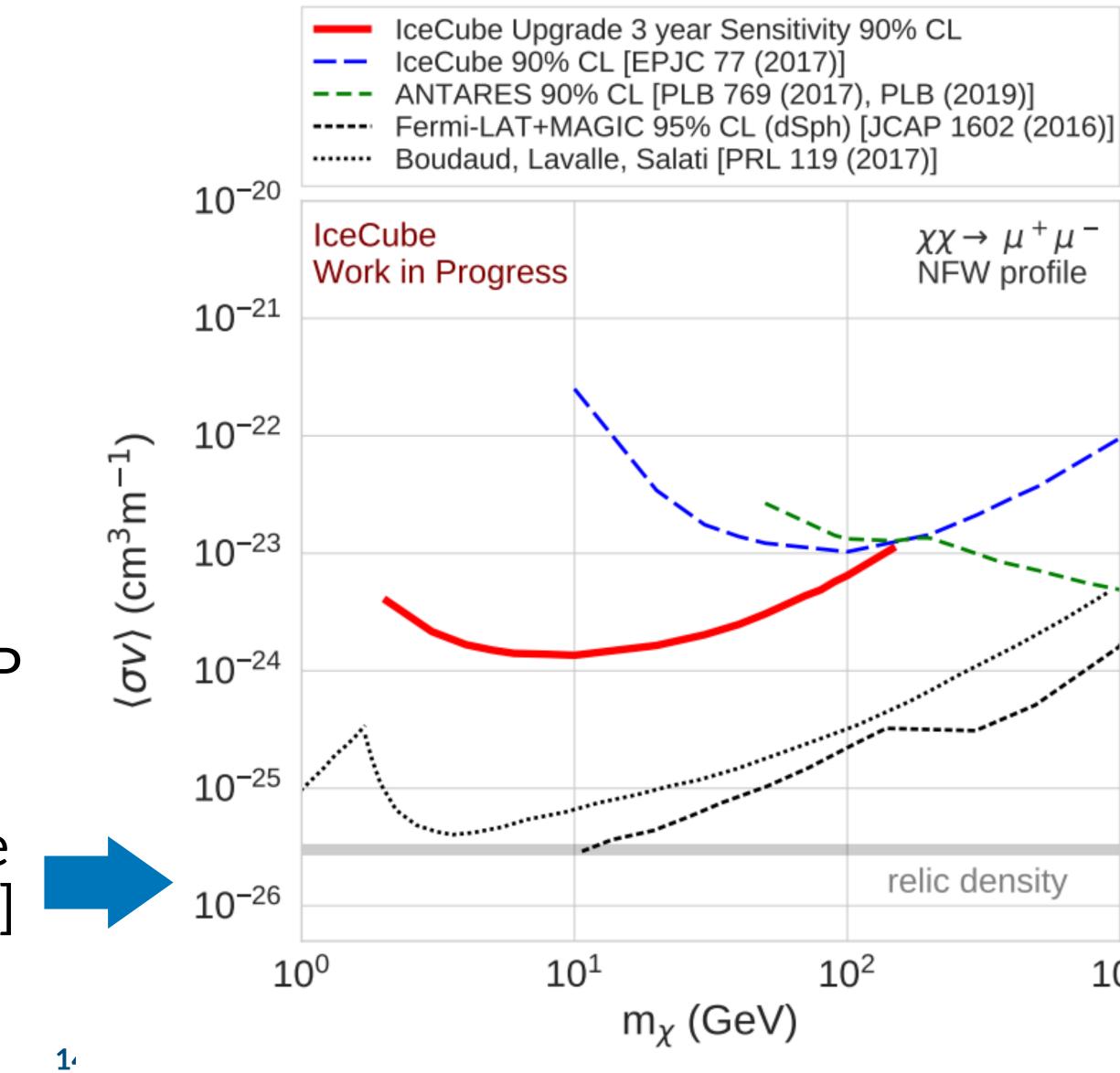






## **IceCube-Upgrade** Science Case

- Unprecedented sensitivity to atmospheric neutrino mixing parameters and neutrino mass ordering
- Detailed calibration of ice properties.
- Expanding beyond the TeV-WIMP paradigm.
- Preliminary studies for DM made at the IIHE [PoS (ICRC2019) 506]

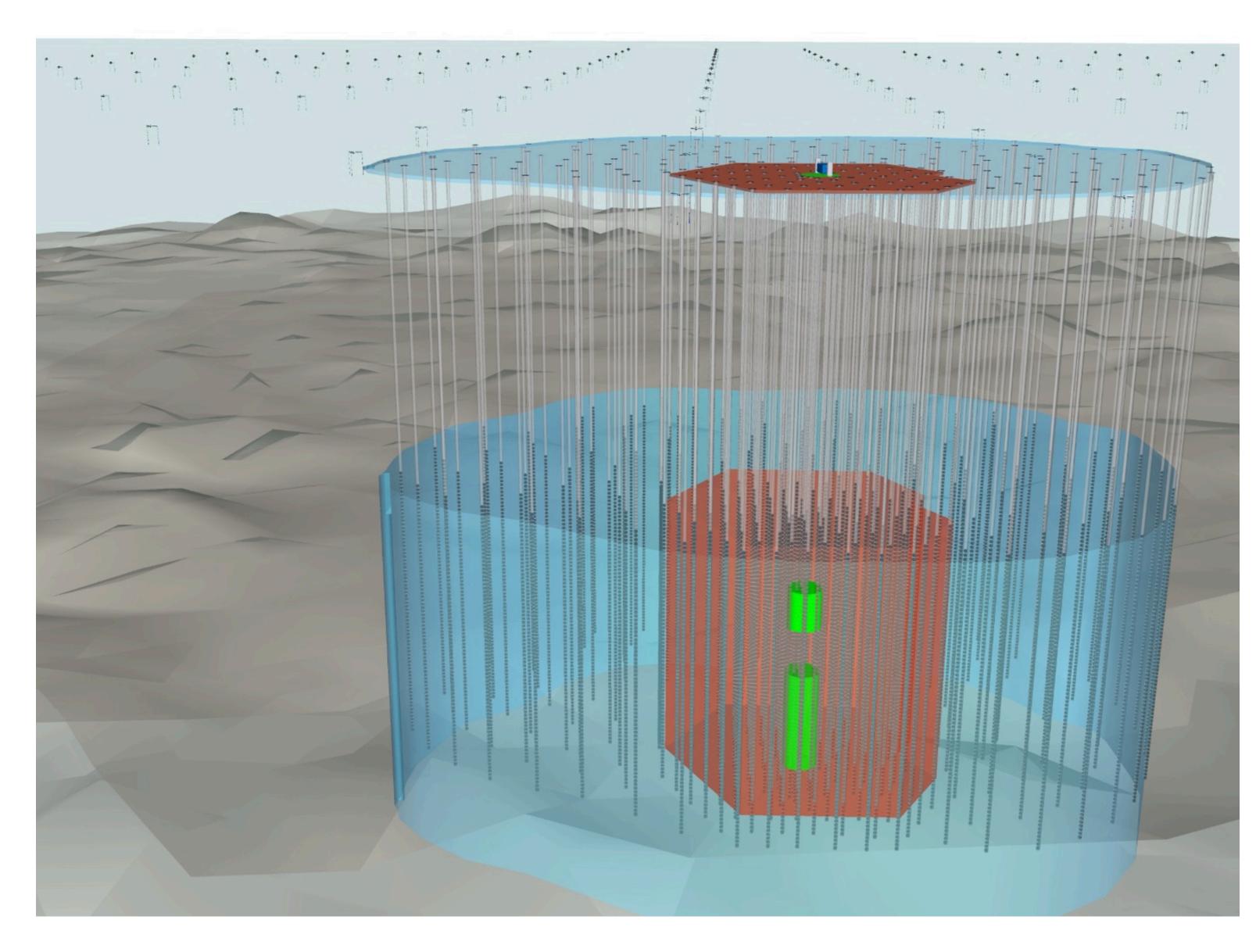






# IceCube-Gen2

- Three new elements, leveraging complimentary technologies, to achieve sensitivity to MeV-EeV neutrinos:
  - Enlarge deep optical array
  - Surface Array extension
  - Shallow Radio Array



## IceCube-Gen2 Science

- 5x improvement in effective area
- 2x improvement in angular resolution

### Multimessenger spectroscopy

Is there a change in the spectrum? Is there a cut-off? Are there cosmogenic neutrinos there?

10<sup>-10</sup>

10<sup>-5</sup> '

10

10

10

10<sup>-9</sup> '

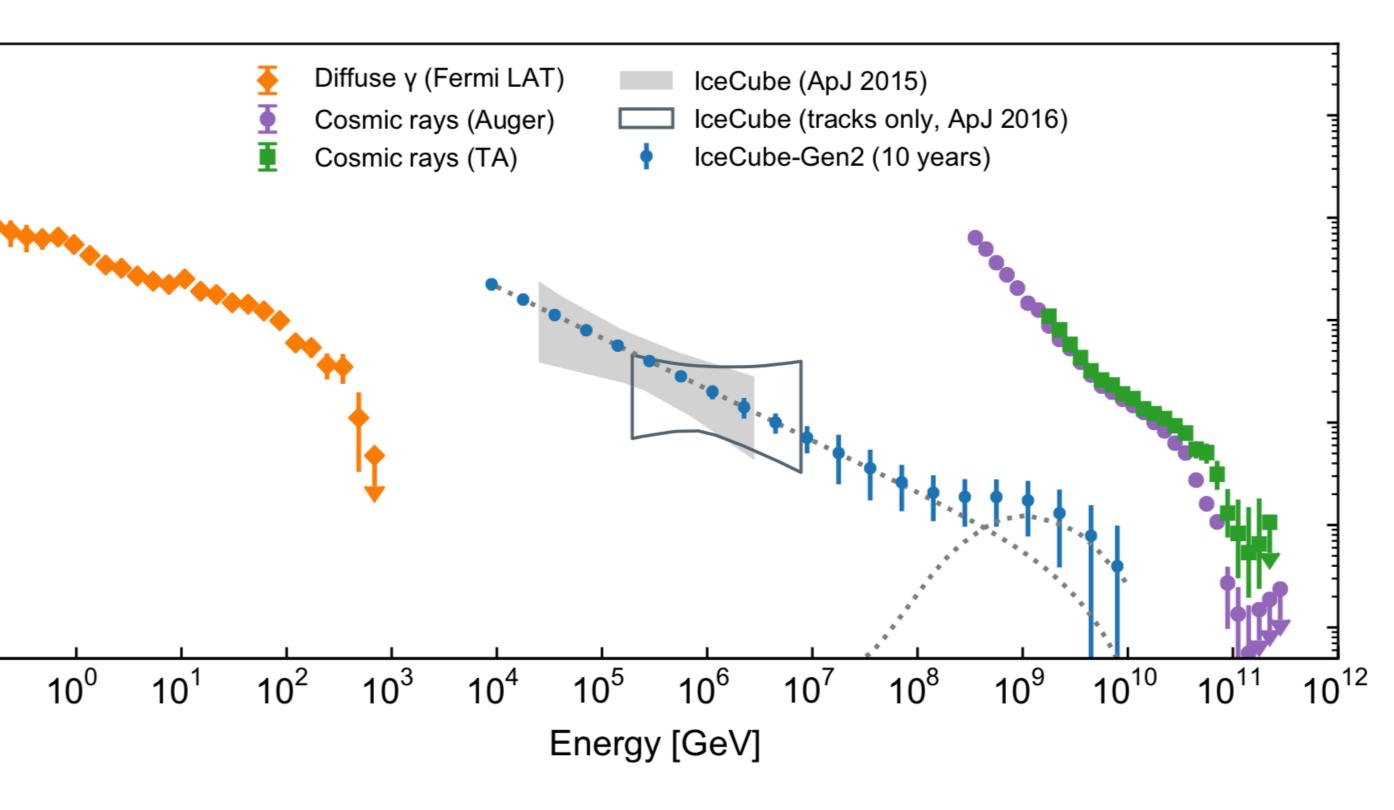
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cm<sup>-2</sup>1

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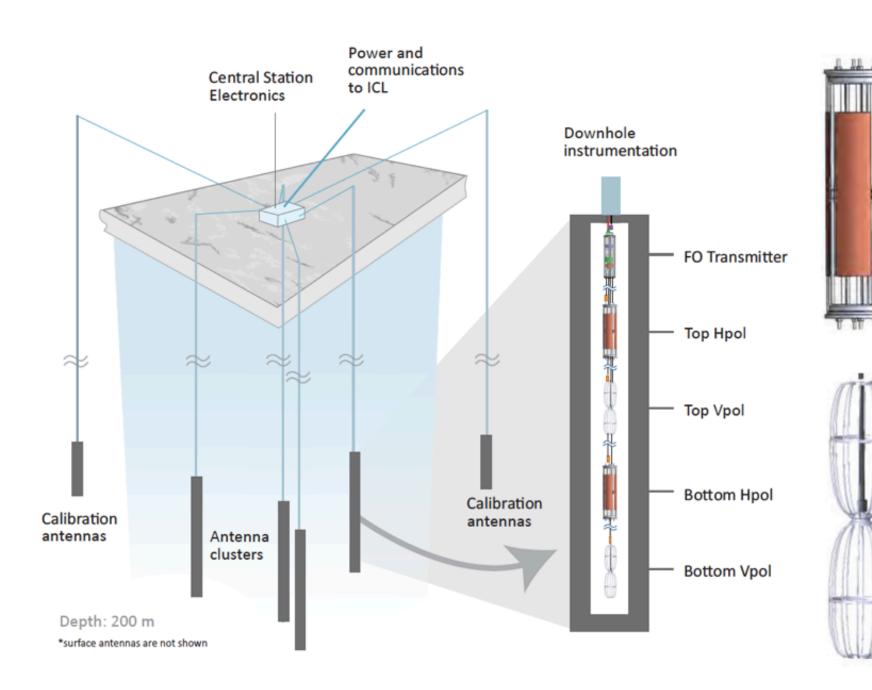
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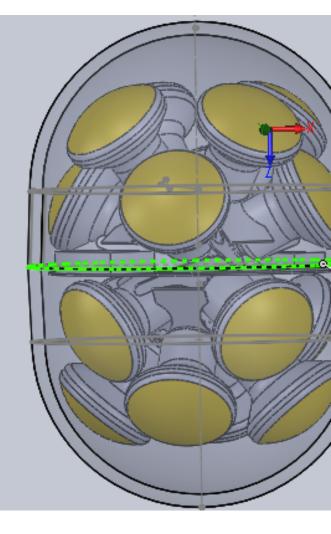
### IceCube-Gen2 Instrumentation

# Pixelated optical modules, surface technology, radio technology, ...









Hpol antenna

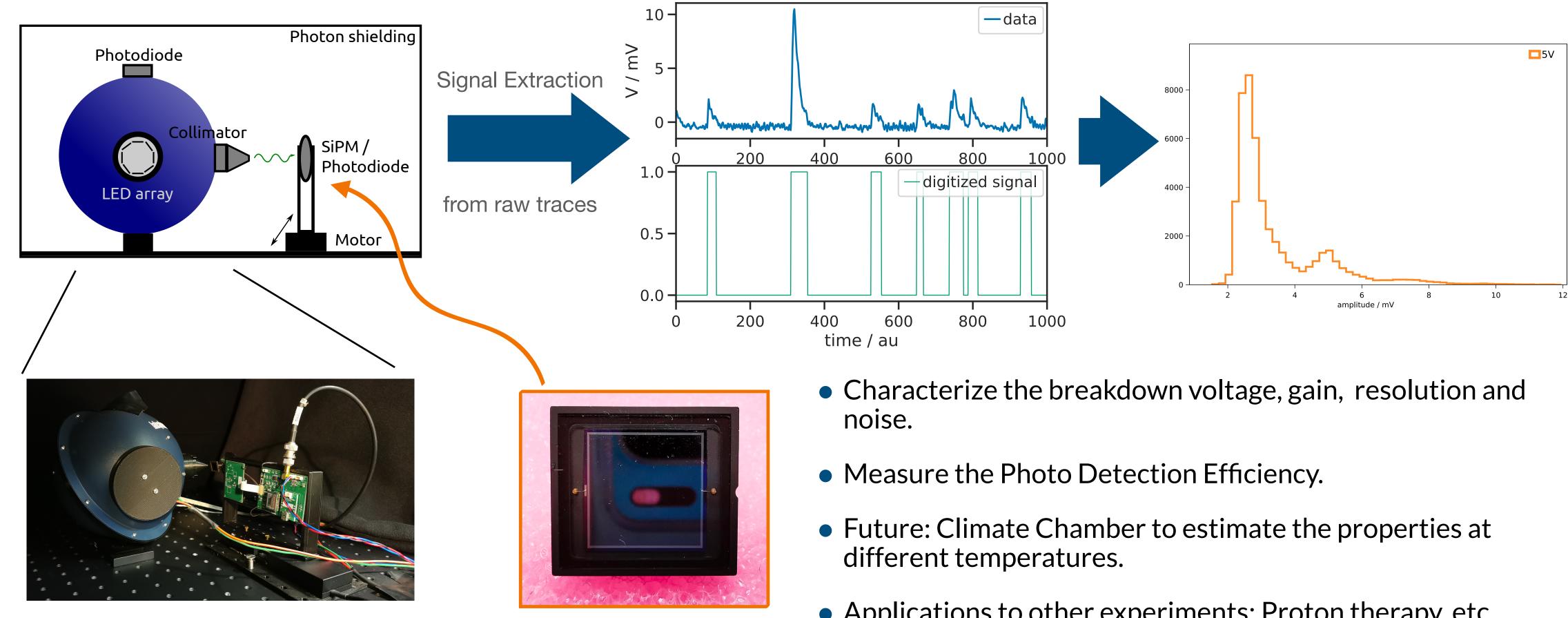
Vpol antenna



17



## IceCube-Gen2@IIHE **Calibration and Characterization of SiPMs**



- Applications to other experiments: Proton therapy, etc.

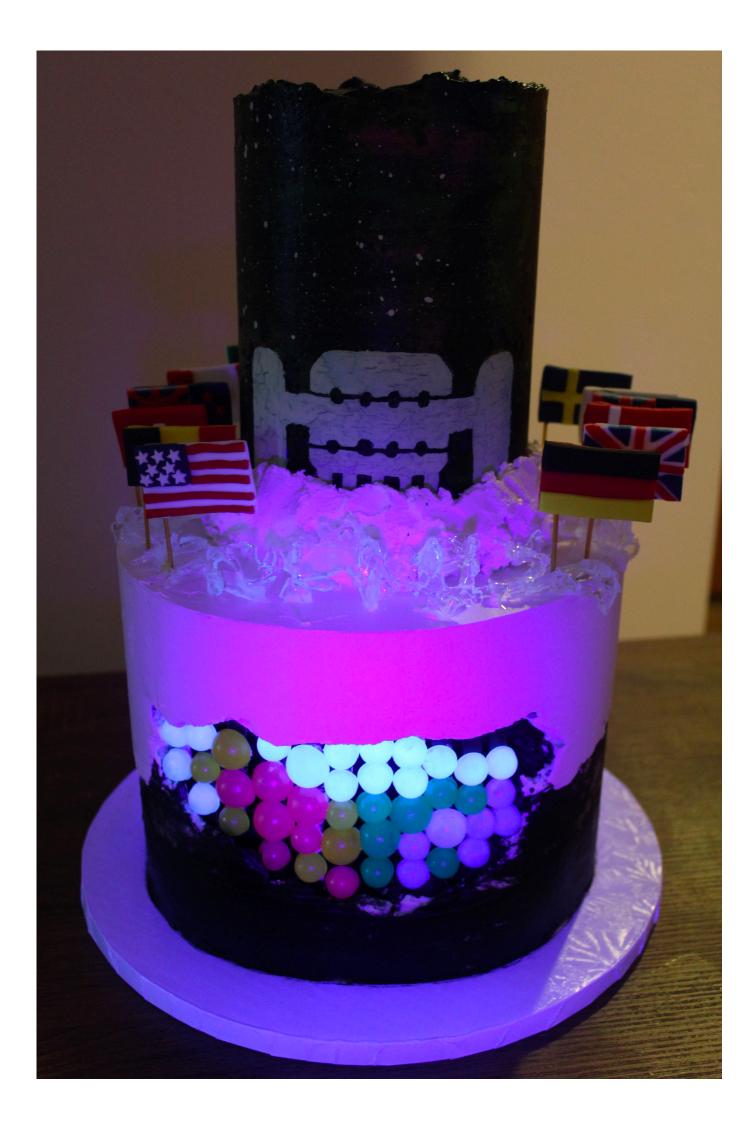
# Conclusions

- IceCube just had its 10 years birthday!
- We witnessed several important results in this past decade, and the IIHE was very much involved but...
- ...we need a bigger detector!
- IceCube-Gen2 is designed and optimized to harvest the enormous scientific opportunities.



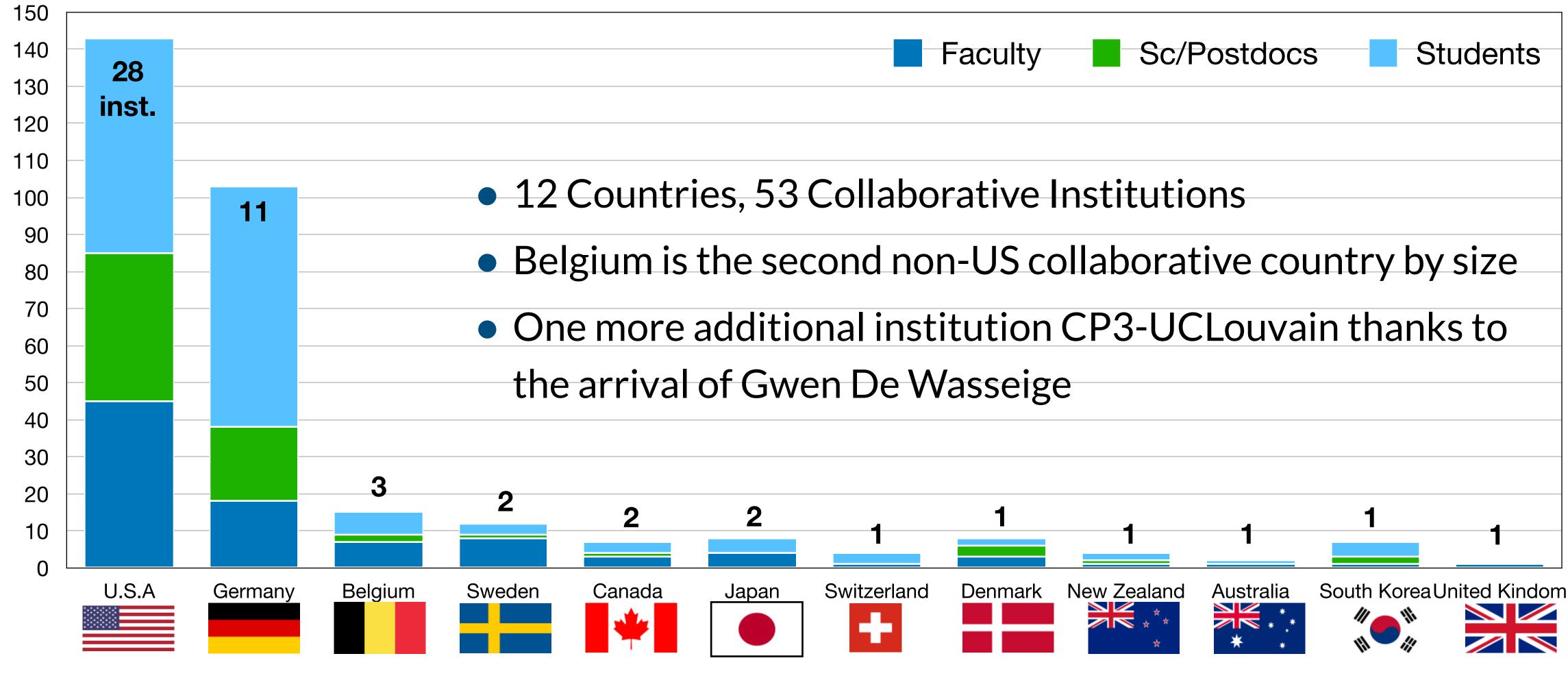
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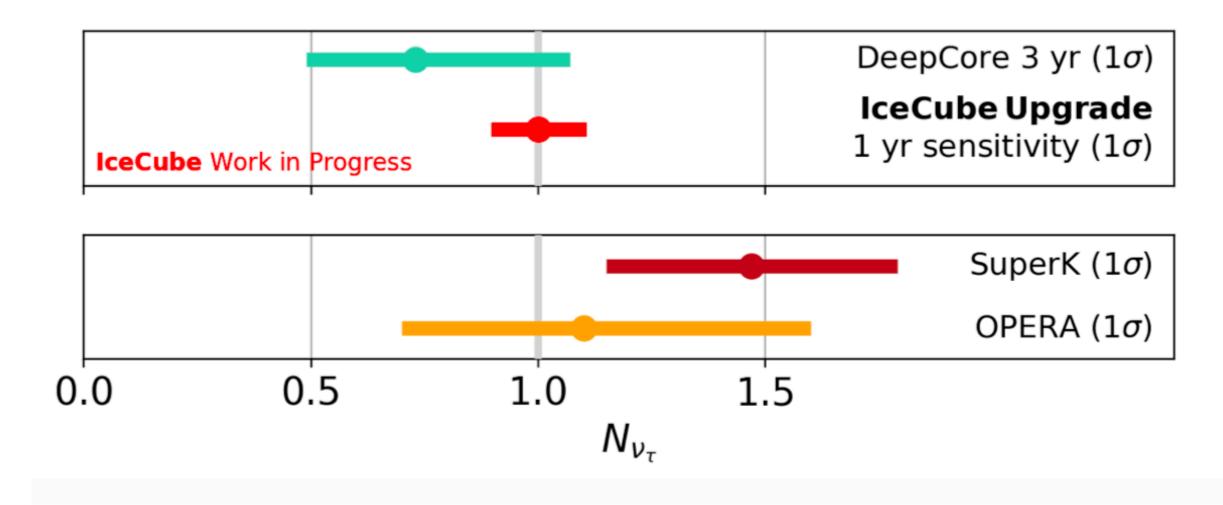
# **Belgium in IceCube**



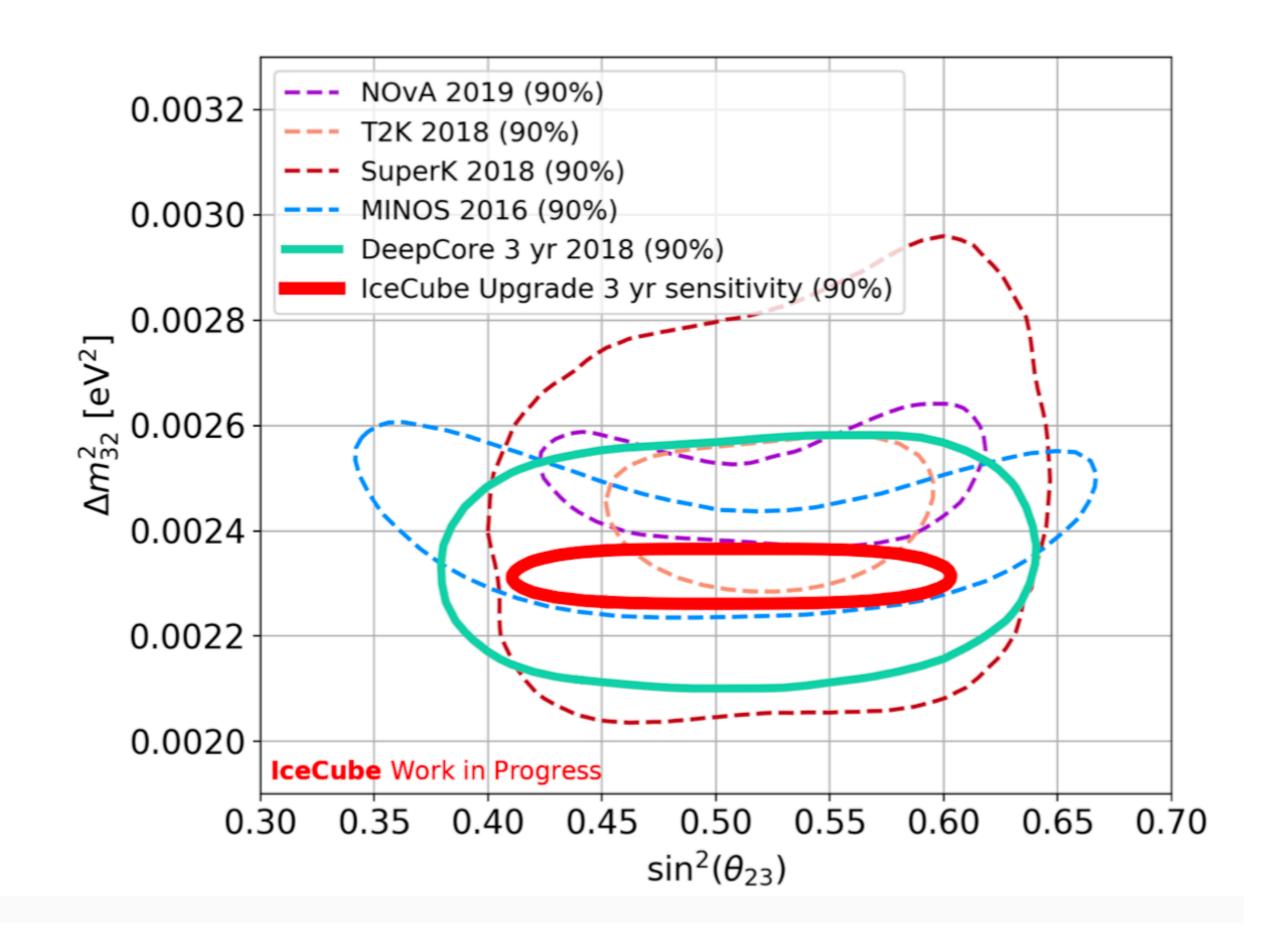


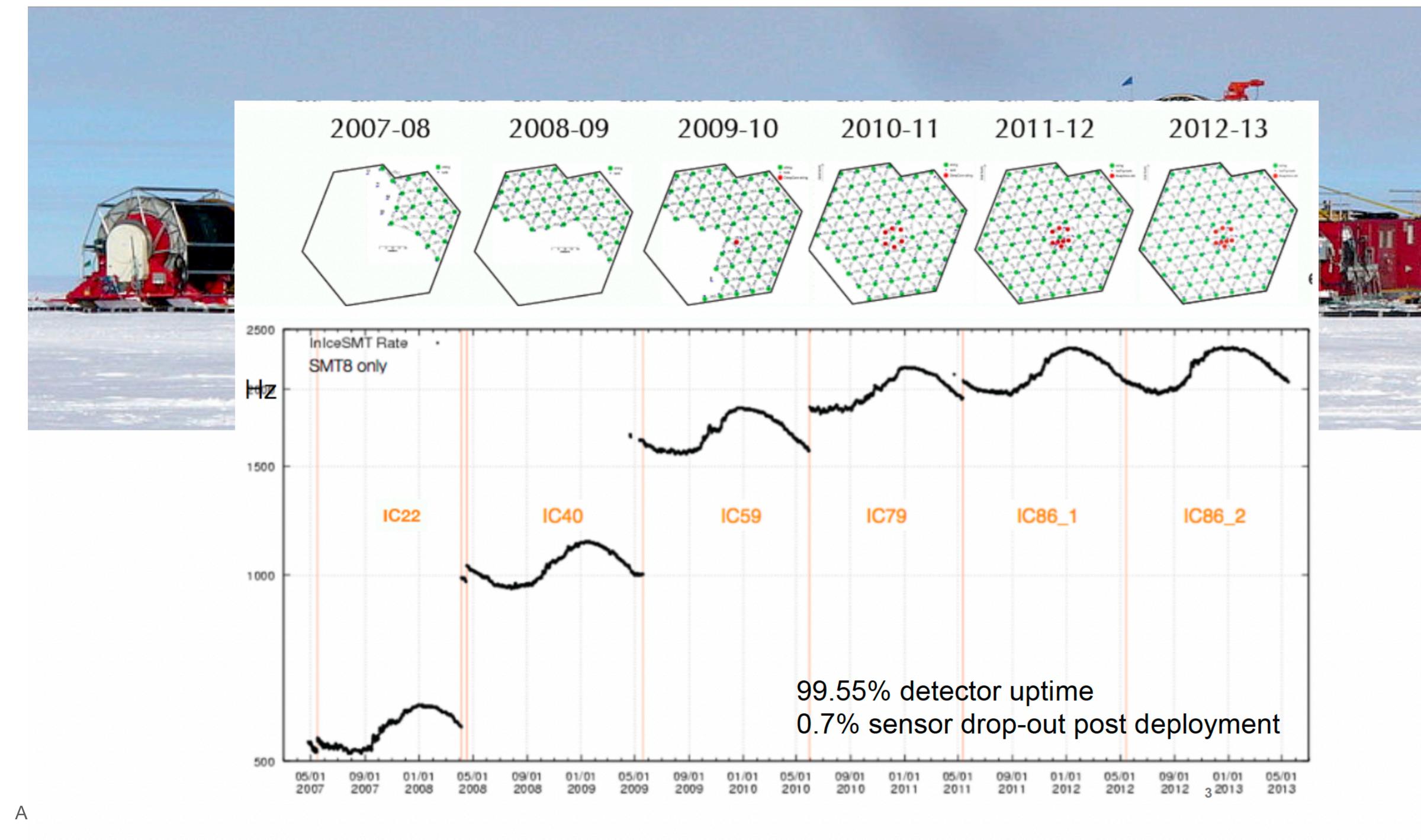


# IceCube-Upgrade



Annual Meeting 2021







#### **IceCube Installation**



Operating sensors in the ice since 2006, with no evidence for aging

#### New surface technology



Scintillator / radio station deployed at South Pole (2019) (PoS ID 314)

#### IceCube Upgrade / Gen2 Phase 1



Deployment of next generation sensors (see next slide)

#### **Radio-Tests in Greenland**



Radio technology deployed in Greenland (2021, see S. Wissel et al., PoS ID 001)





### **IceCube-Gen2 Plan** Simplified Plan

IceCube Upgrade (Phase 1)

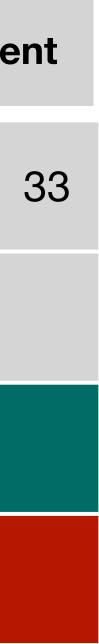
IceCube-Gen2 (Phase 2) MREFC

International (non-US) funding

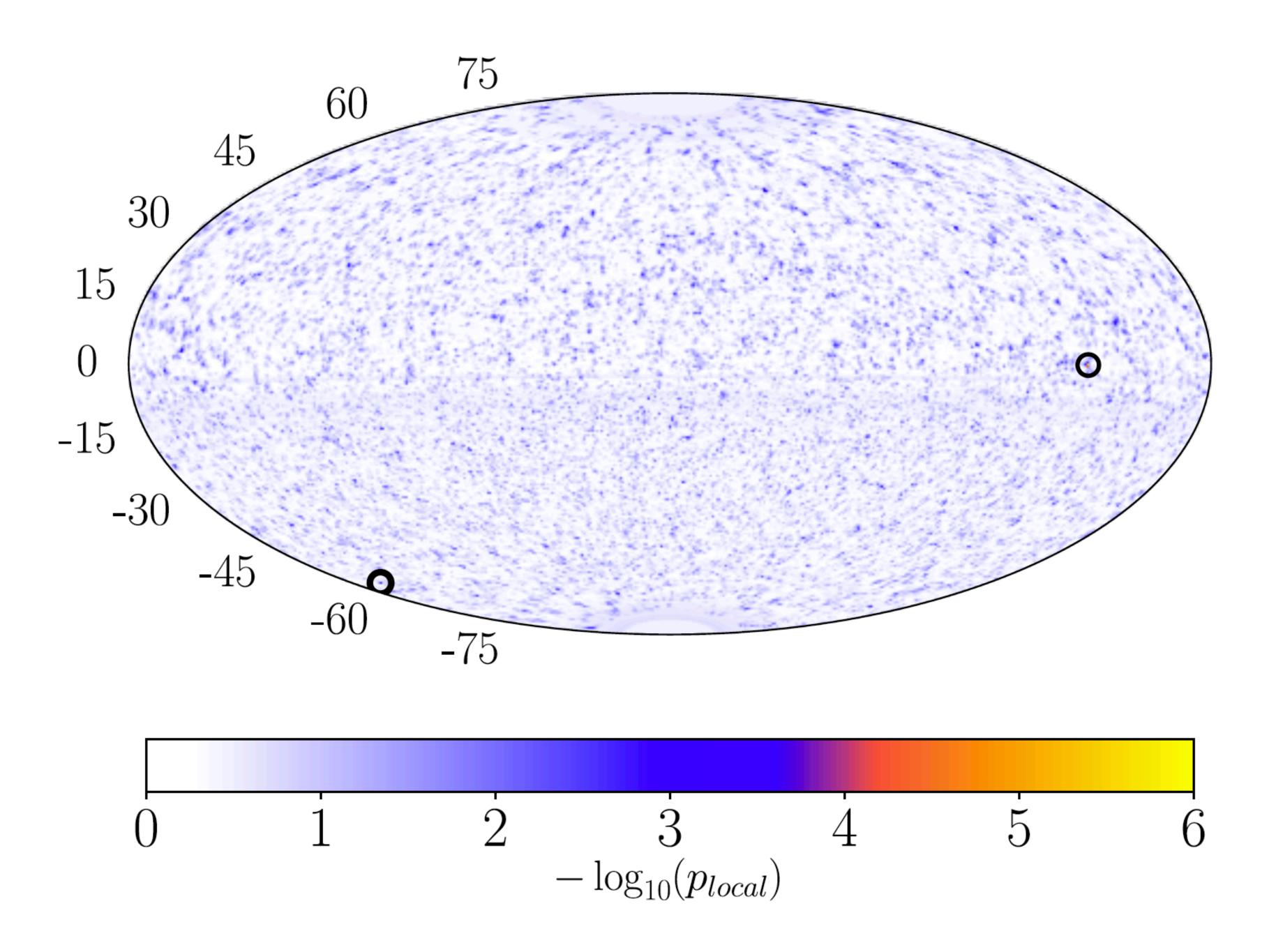
19	20	21	22	23	24	25	26	27	28	29	30	31	32	32



#### Major field activity / Deployment







Annual Meeting 20