Extracting the meter-scale structure of lightning plasma with the LOFAR telescope





Brian Hare & LOFAR Lightning collaboration

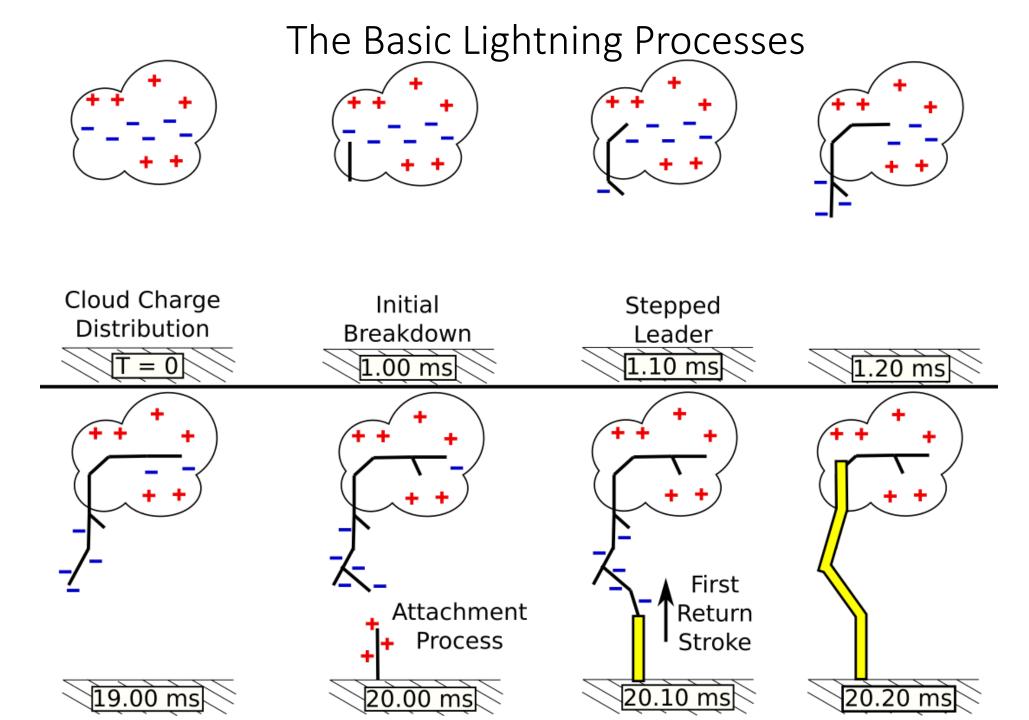


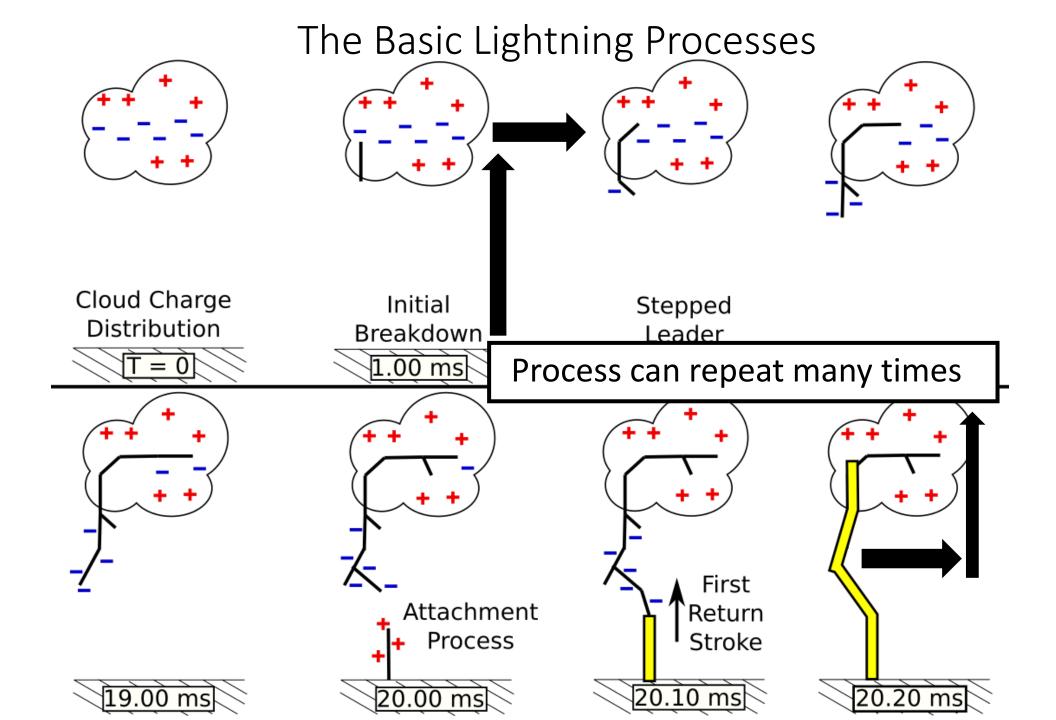


UB VRIJE UNIVERSITEIT BRUSSEL

AST RON

Netherlands Institute for Radio Astronomy

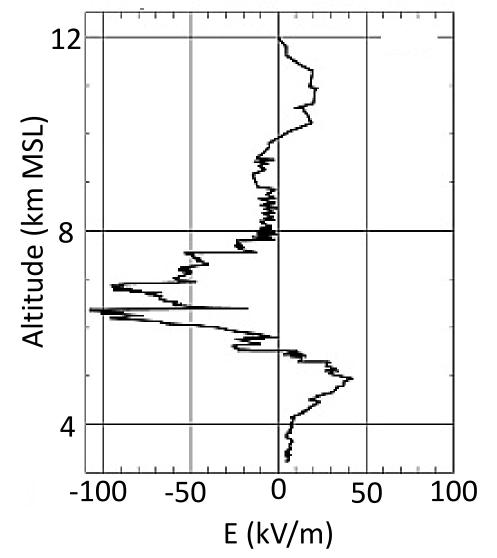






Lightning Initiation

balloon-borne electric field measurement



Dielectric strength of air is 3000 kV/m

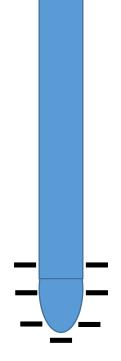
Order-of-magnitude higher then measured thunderstorm electric fields

Two current major hypothesis:

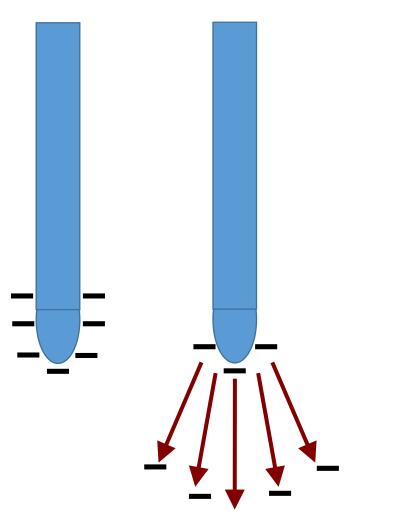
-Cosmic Ray Air Showers

-Electric field amplification by hydrometeors

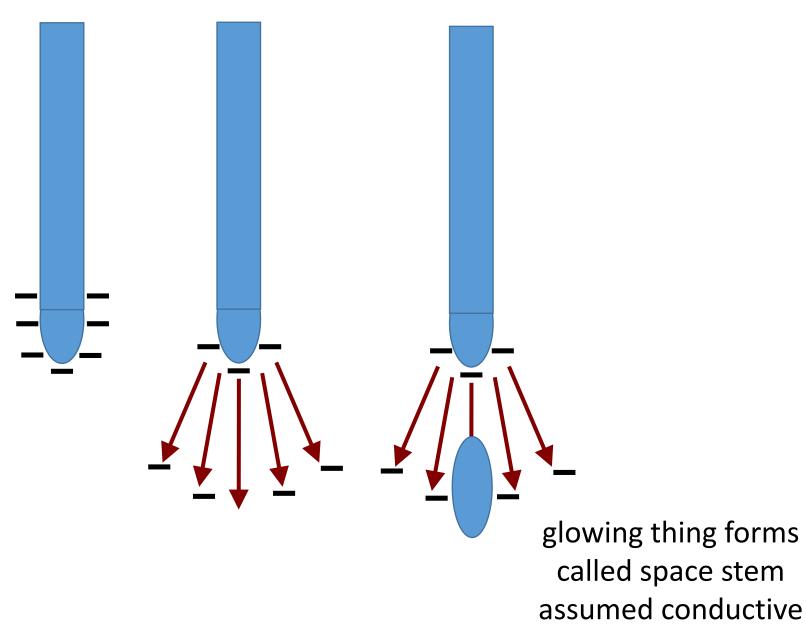
-There are others

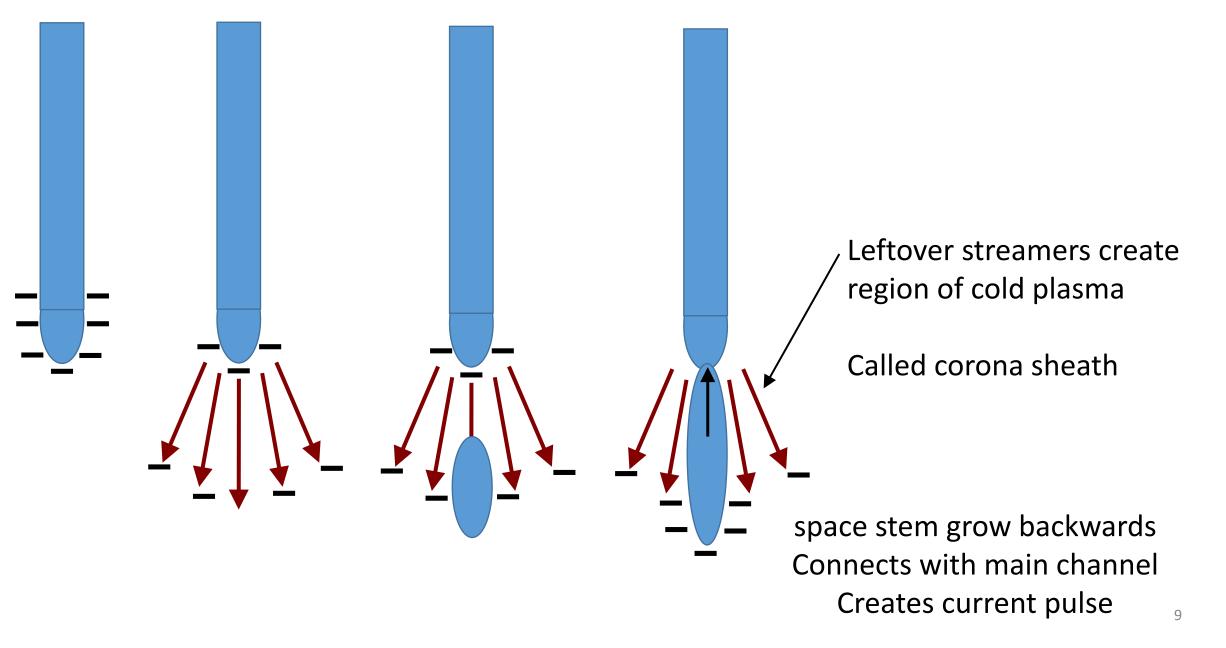


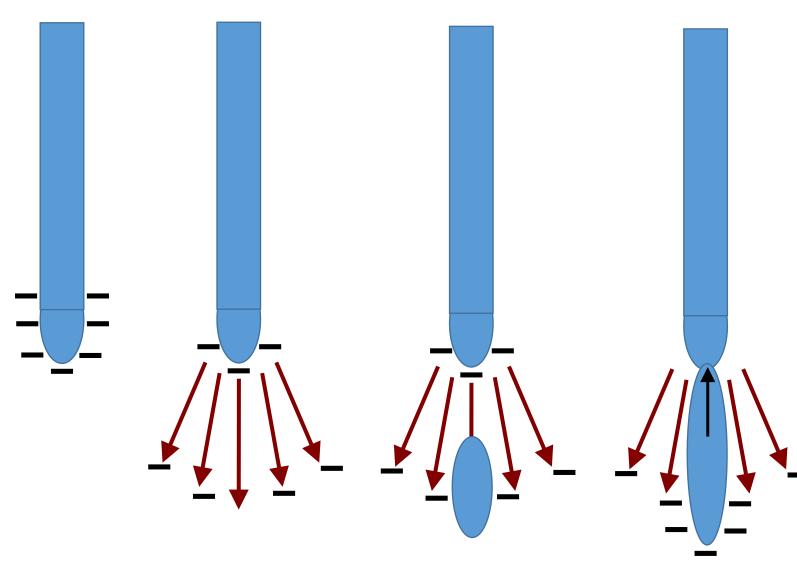
Start with existing channel



Corona Burst cold plasma surges forward







- Pieced together from camera observations and long lab sparks
- Cannot be modeled
- What part emits VHF
- What is a space stem?

Terrestrial Gamma Ray Flashes



Credit: NASA/Goddard Space Flight Center/J. Dwyer/Florida Inst. of Technology

- Intense bursts of gamma radiation from lightning
 - Energies up to 20 MeV
 - Produced around 10 km altitude
 - Intense flux saturates orbiting gamma-ray observatories

The Big Questions in Lightning

- How does lightning get started?
 - Measured electric fields are too small to make a spark via typical mechanisms

- How does lightning grow?
 - The plasma physics is extremely complex, and too complicated for current computers to model

- How does lightning emit gamma rays?
 - This strongly depends on how the lightning grows

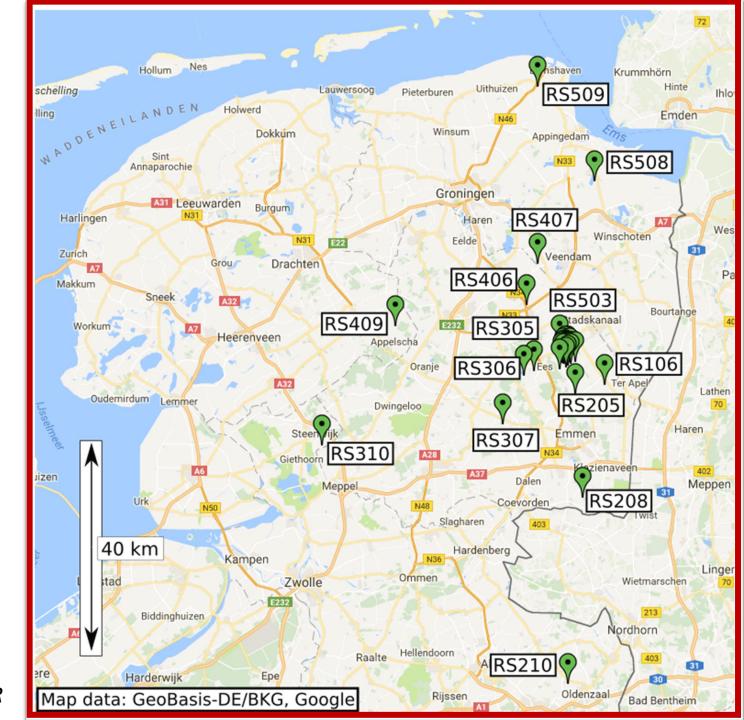
Dutch LOFAR

- 24 "core" stations
- 14 "remote" stations
- 3200 km² enclosed area

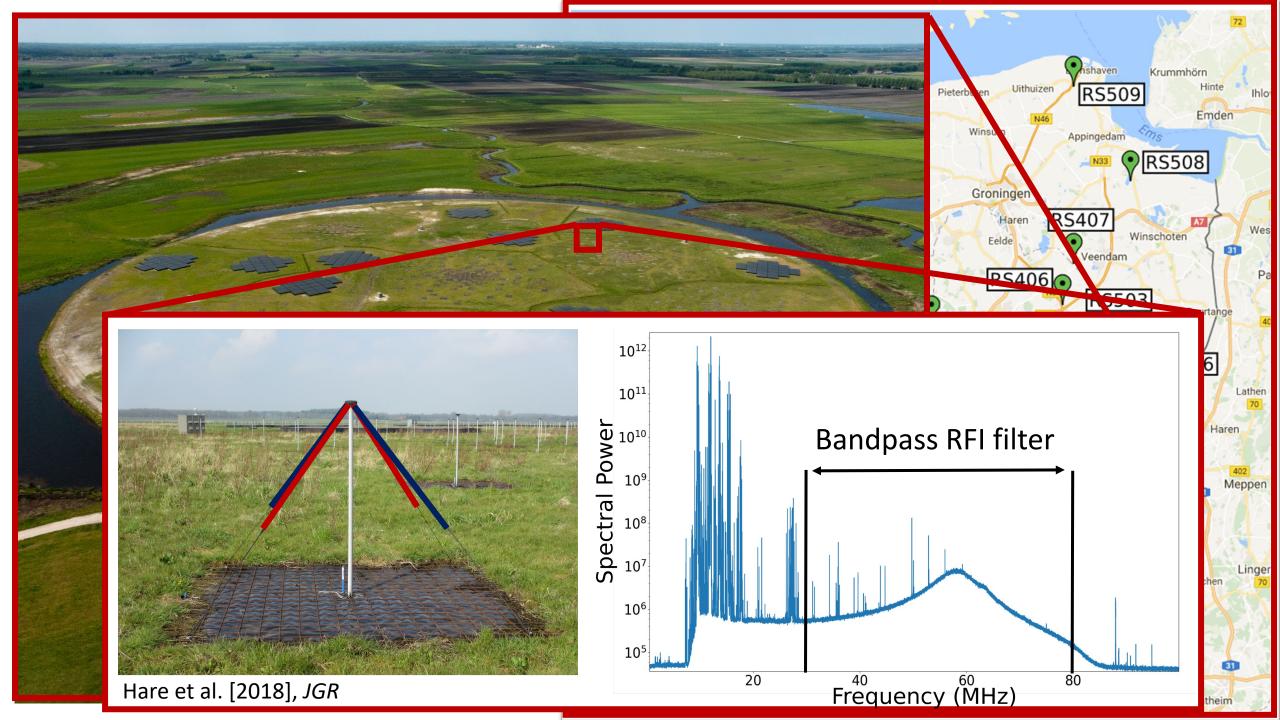
Per Station:

- 96 low-band antennas
 - 10 90 MHz
 - 48 dual-polarized pairs
 - We use 6 dual-polarized pairs out of the 48 pairs
- 20 high-band antennas
 - 110-250 MHz
 - presently not utilized

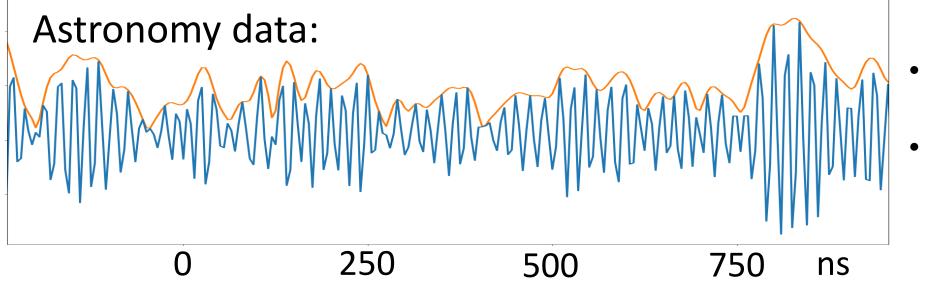
Also shown in Hare et al. [2018], JGR





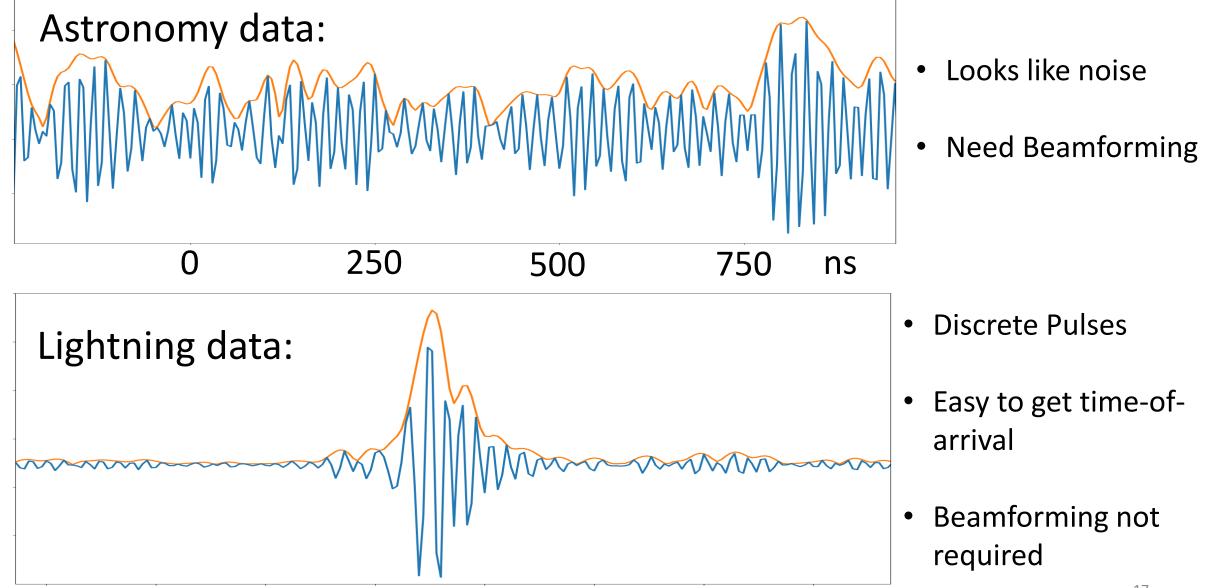


Lightning vs Astronomy Transient Buffer Data

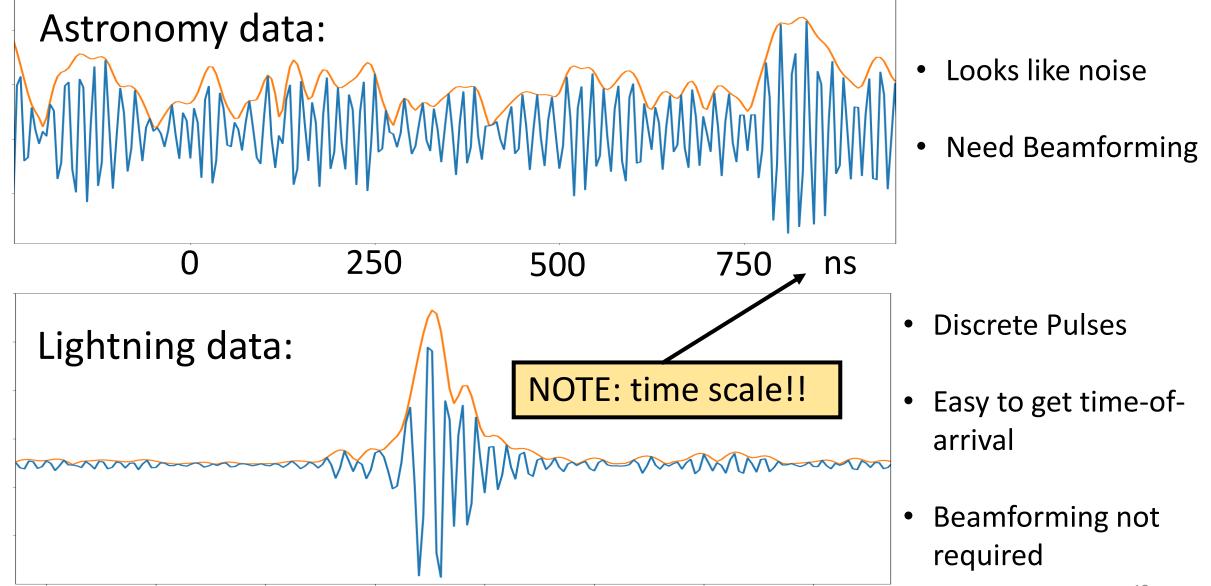


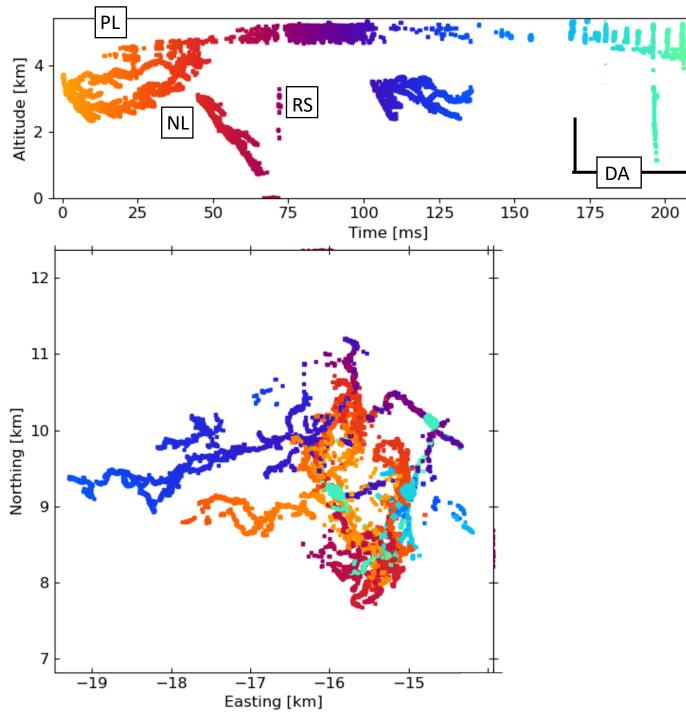
- Looks like noise
- Need Beamforming

Lightning vs Astronomy Transient Buffer Data



Lightning vs Astronomy Transient Buffer Data



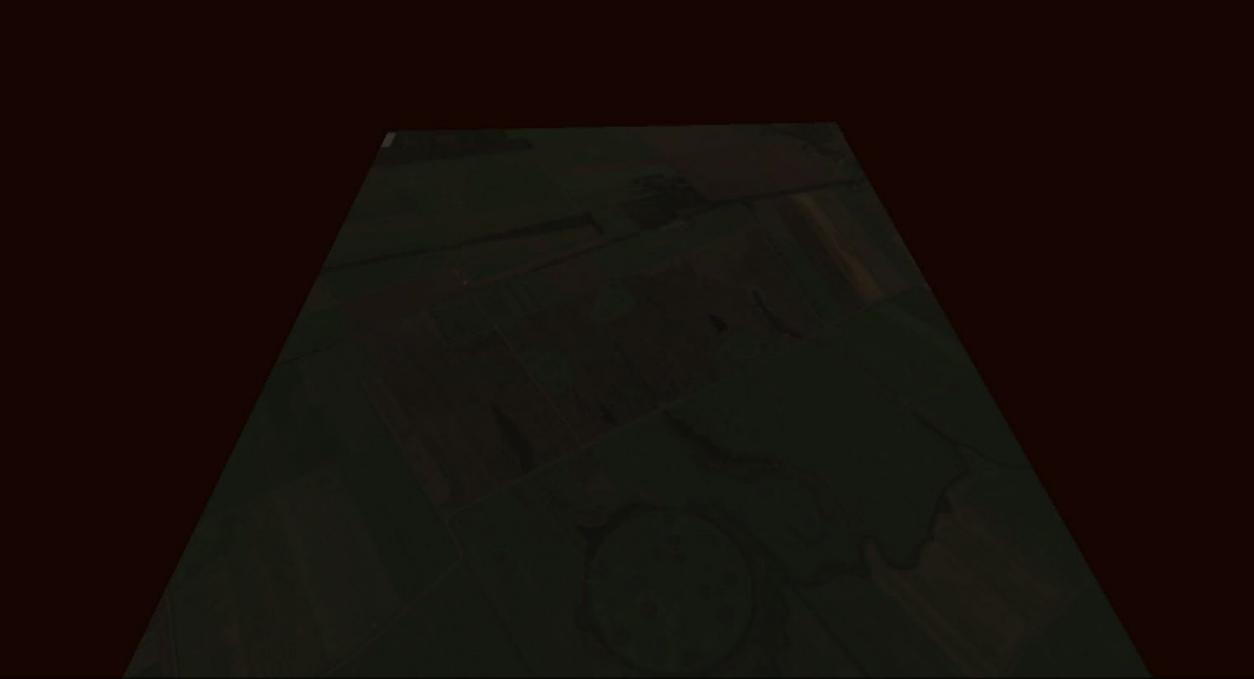


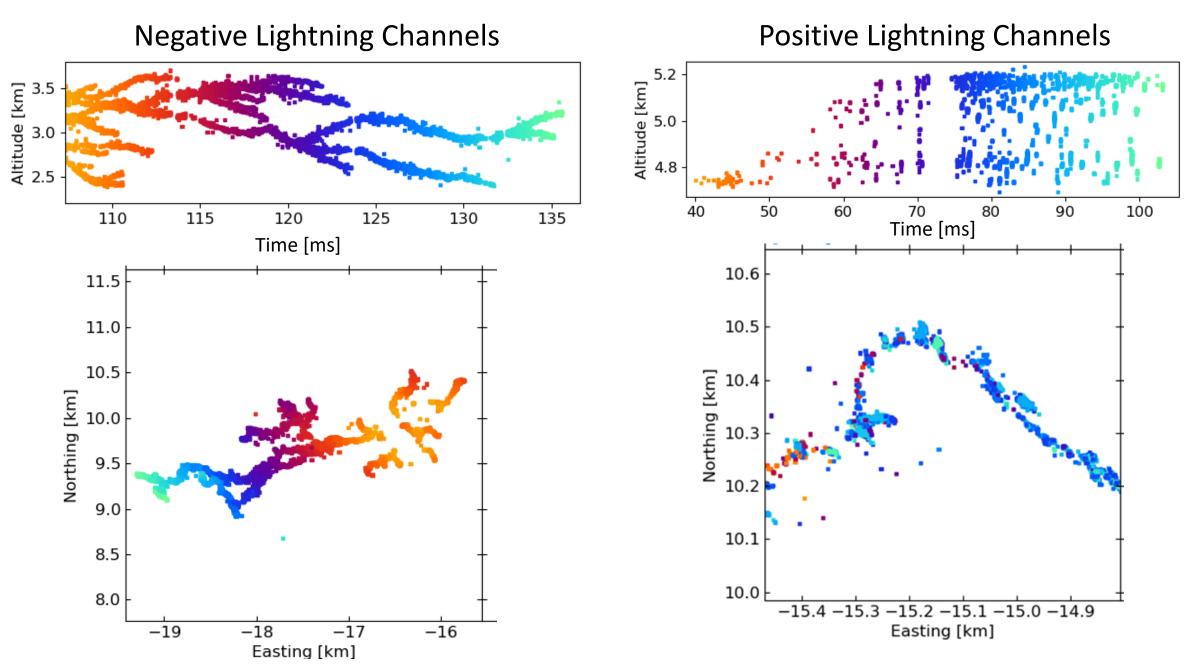
PL - positive leaderNL - negative leaderRS - return strokeDA - dart leader

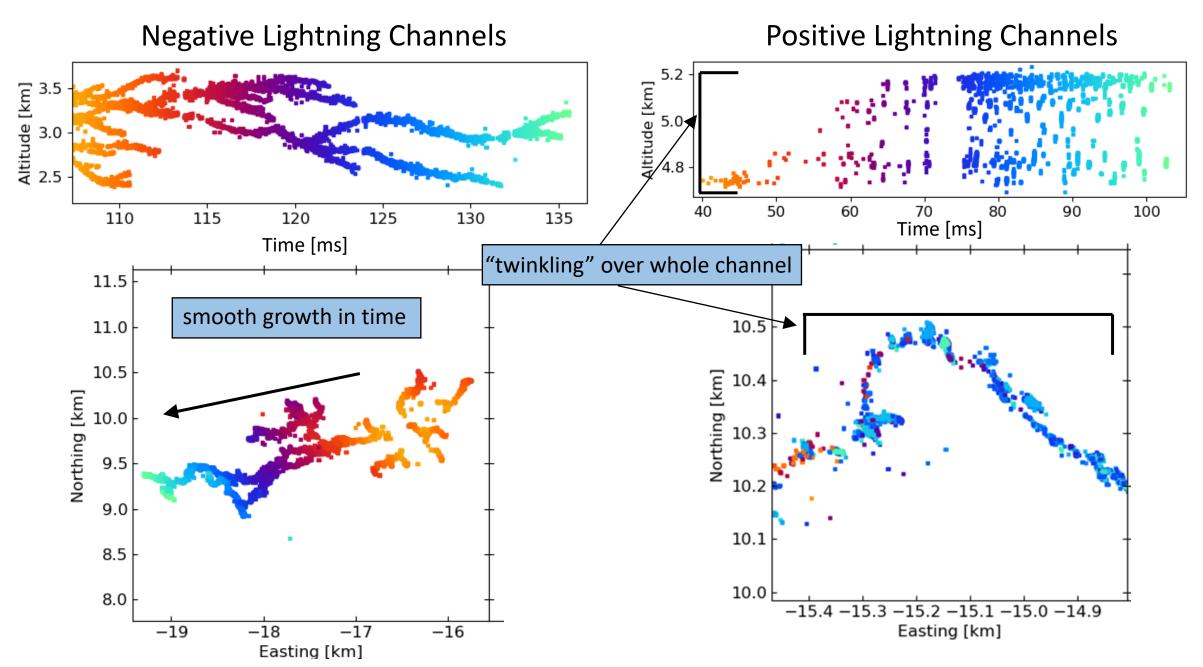
Lightning flash imaged with "impulsive techniques" (TOA) -CPU efficient, not as good as interferometry

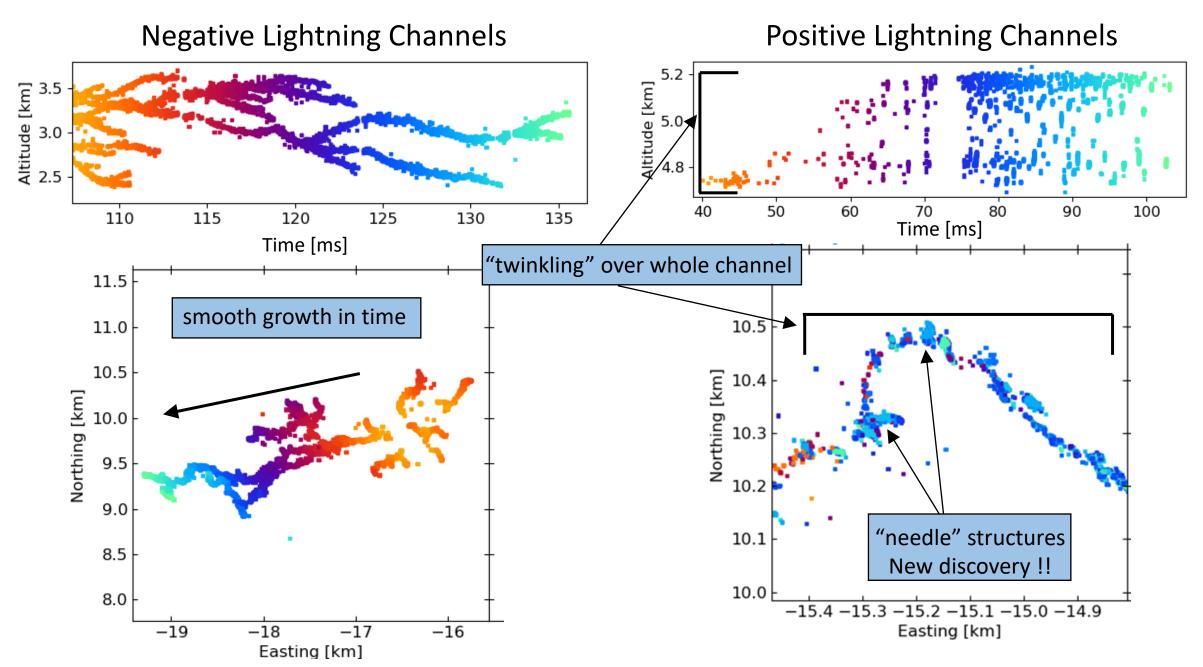
65,200 sources after cuts≈ 300 sources per ms

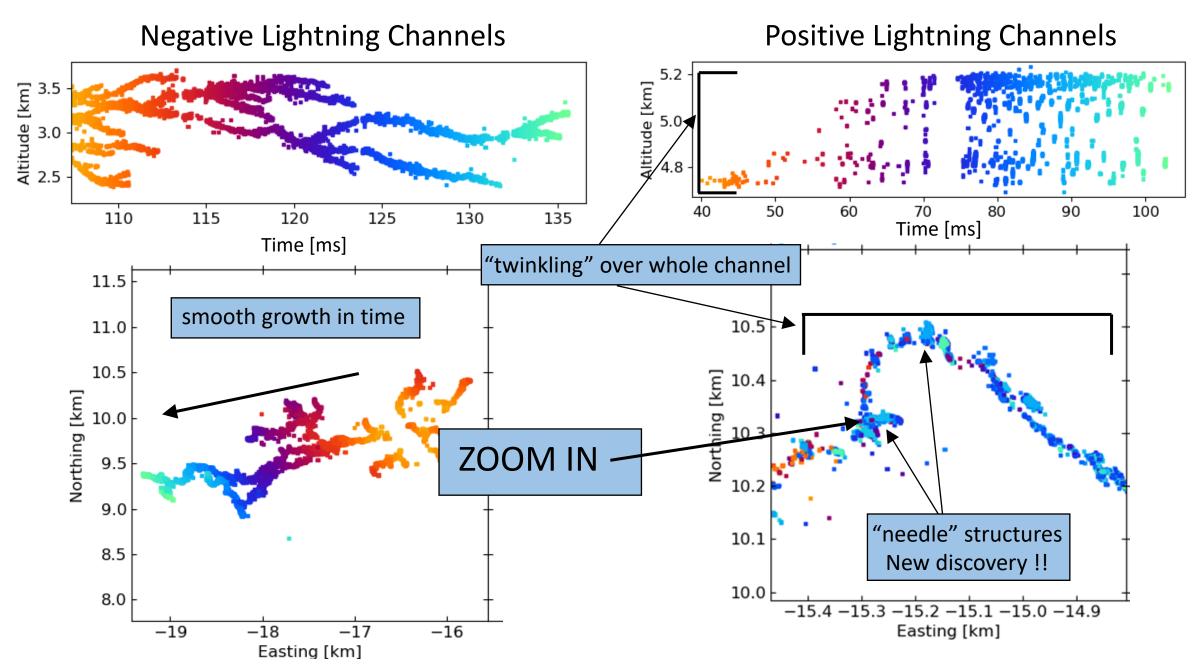
Horizontal accuracy around 1 m -limited by source-confusion





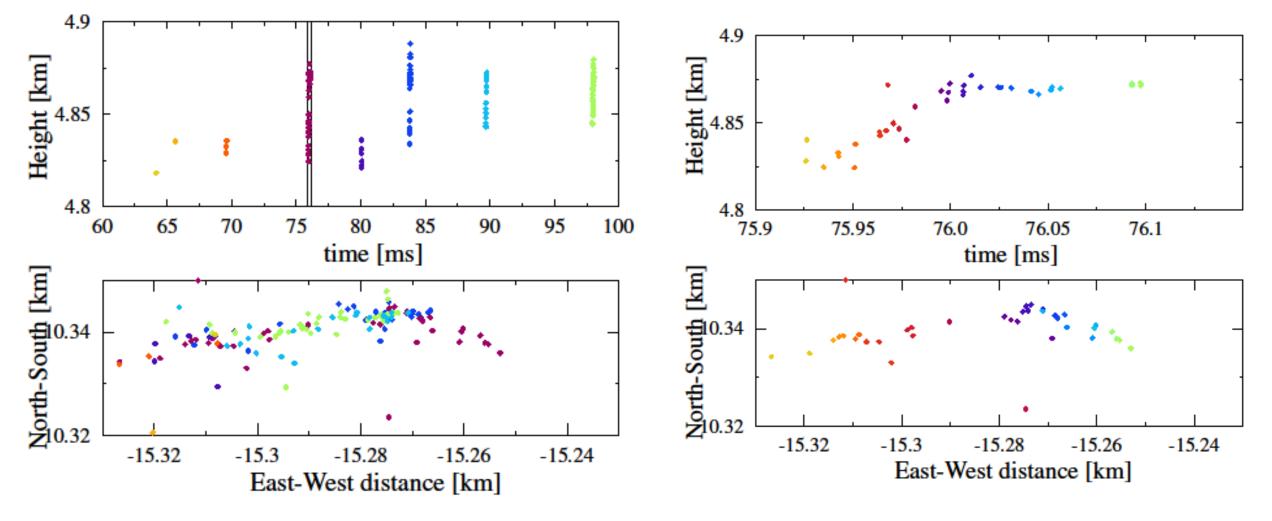




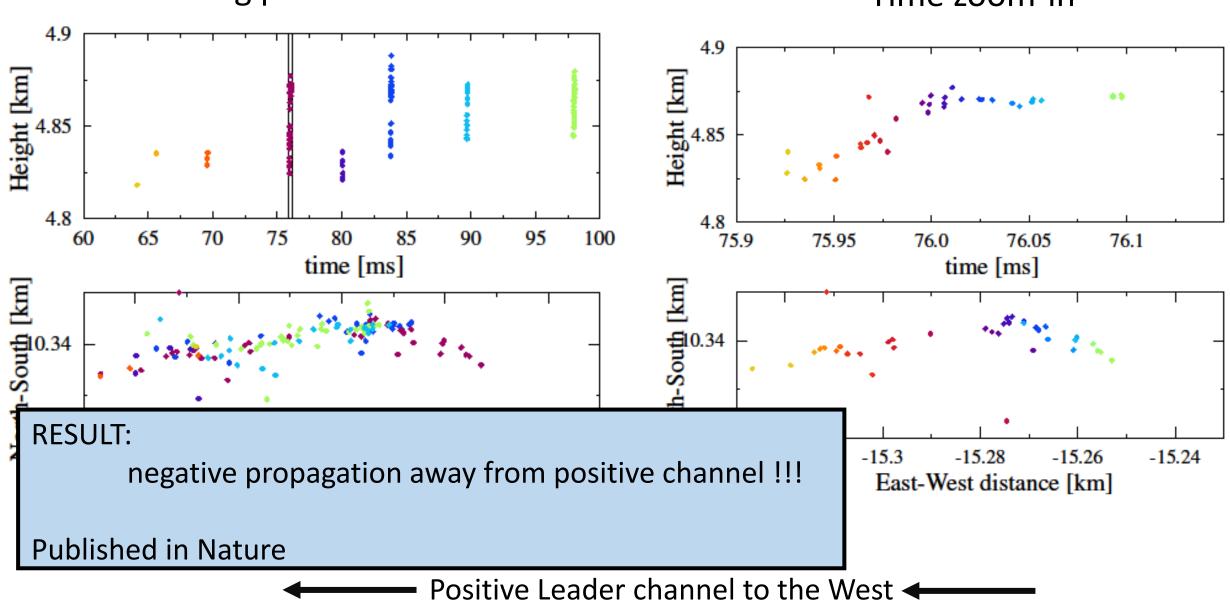


One flickering positive leader structure

Time zoom-in

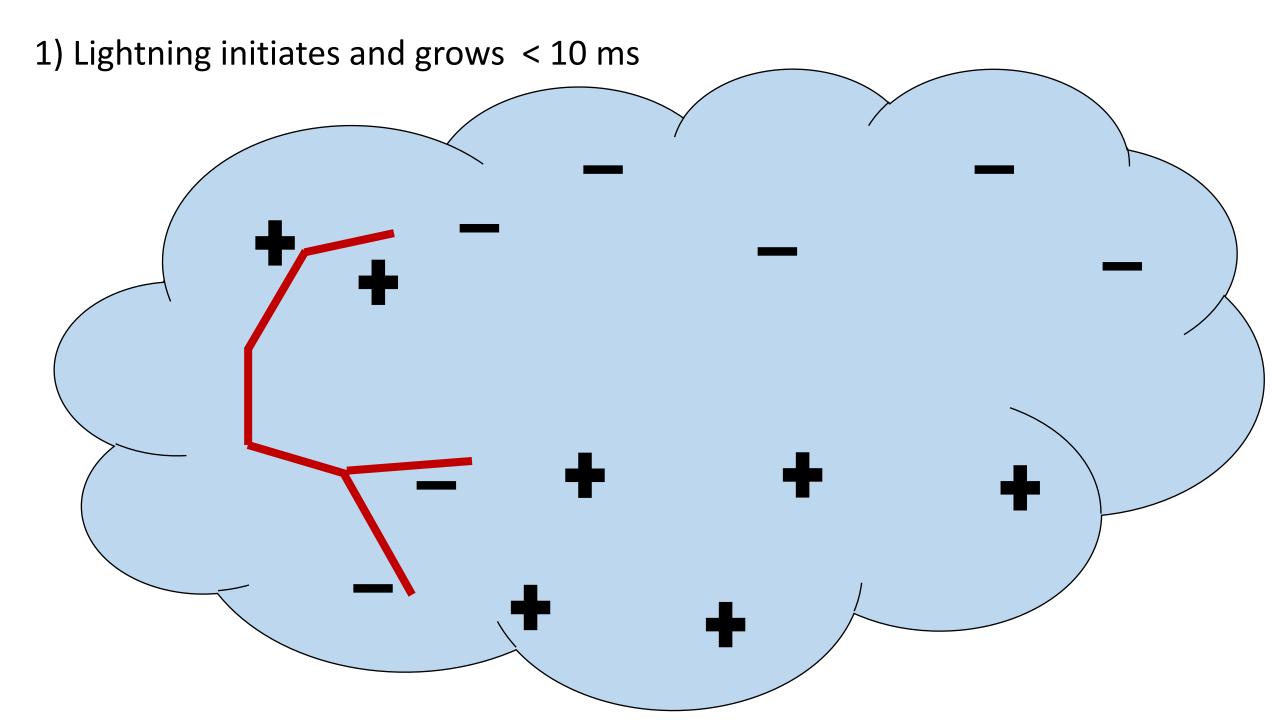


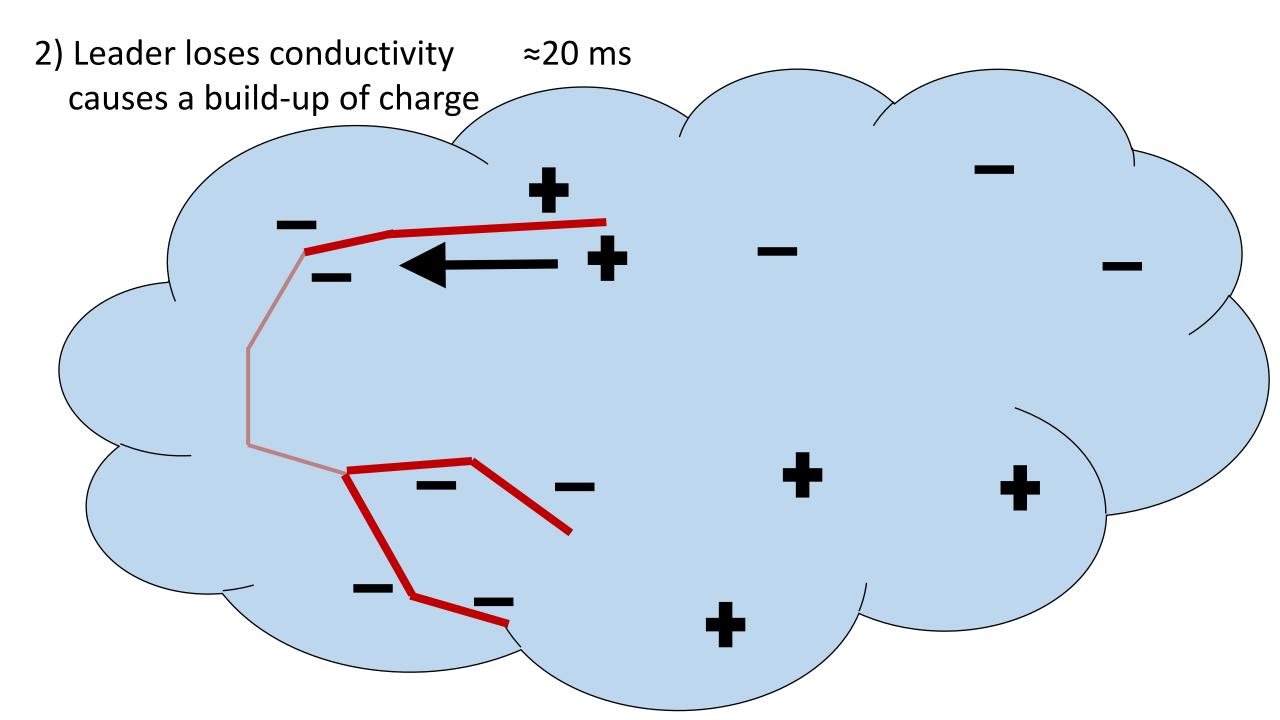
Positive Leader channel to the West

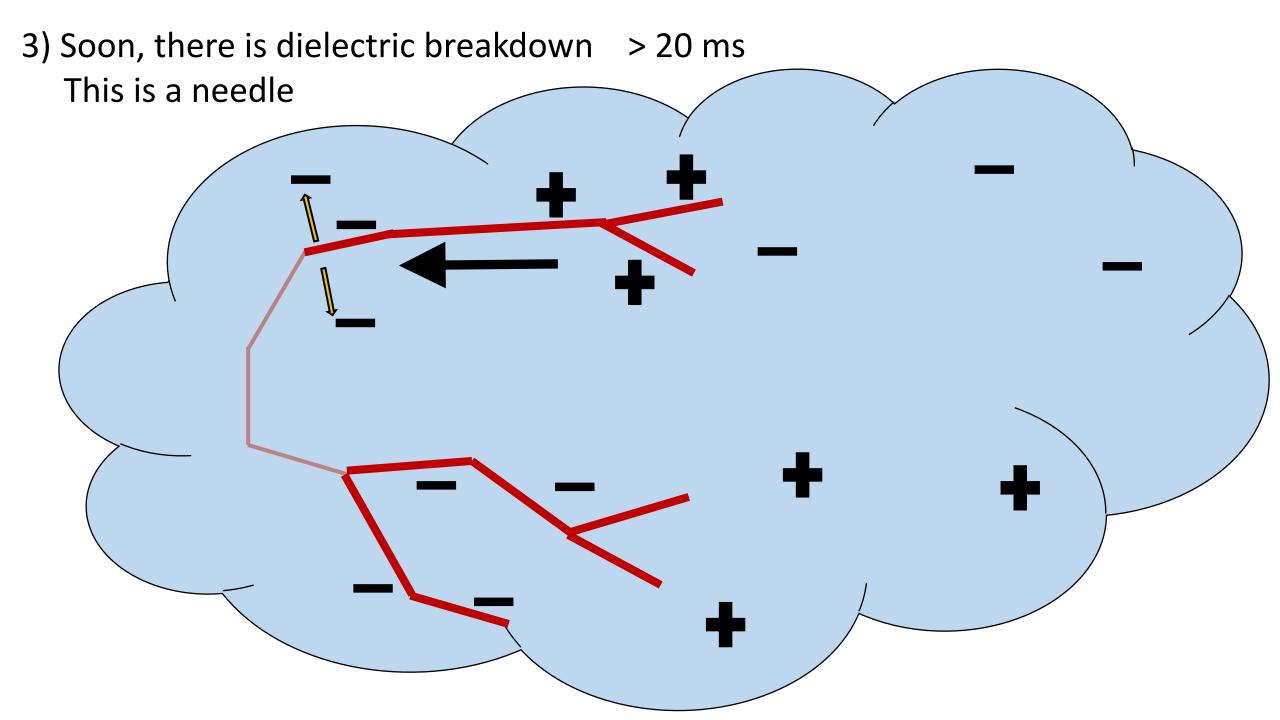


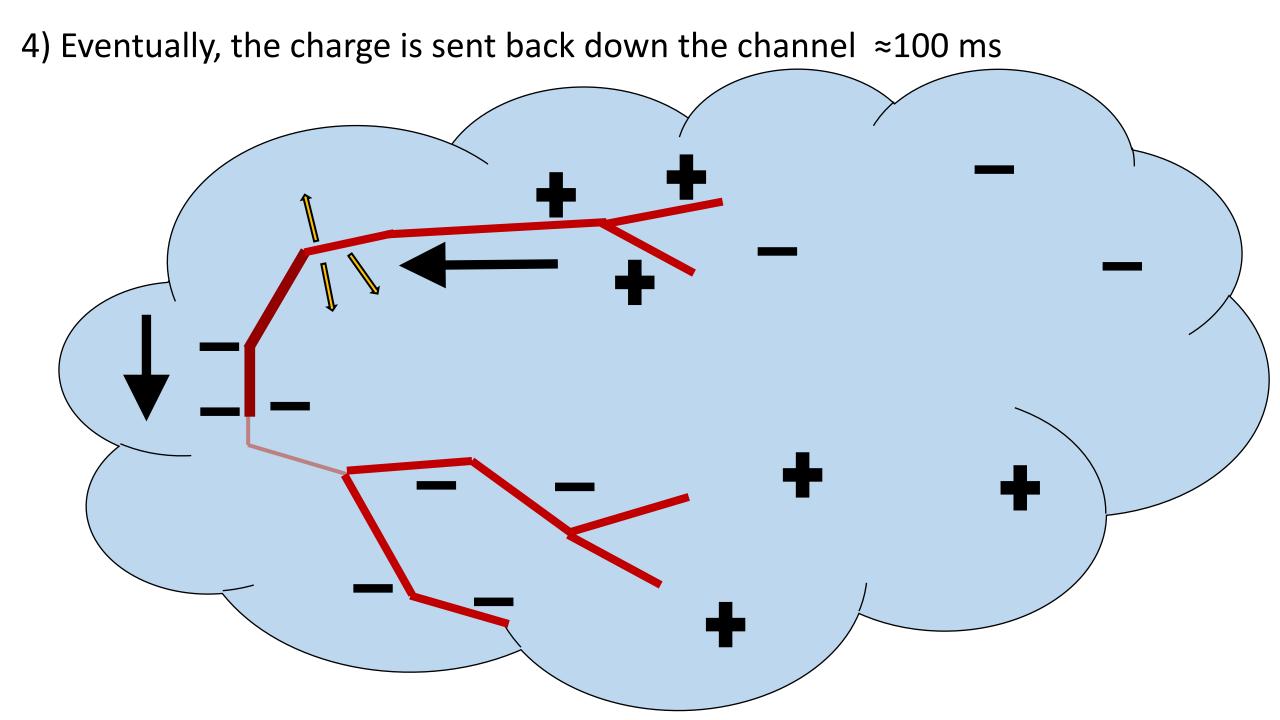
One flickering positive leader structure

Time zoom-in

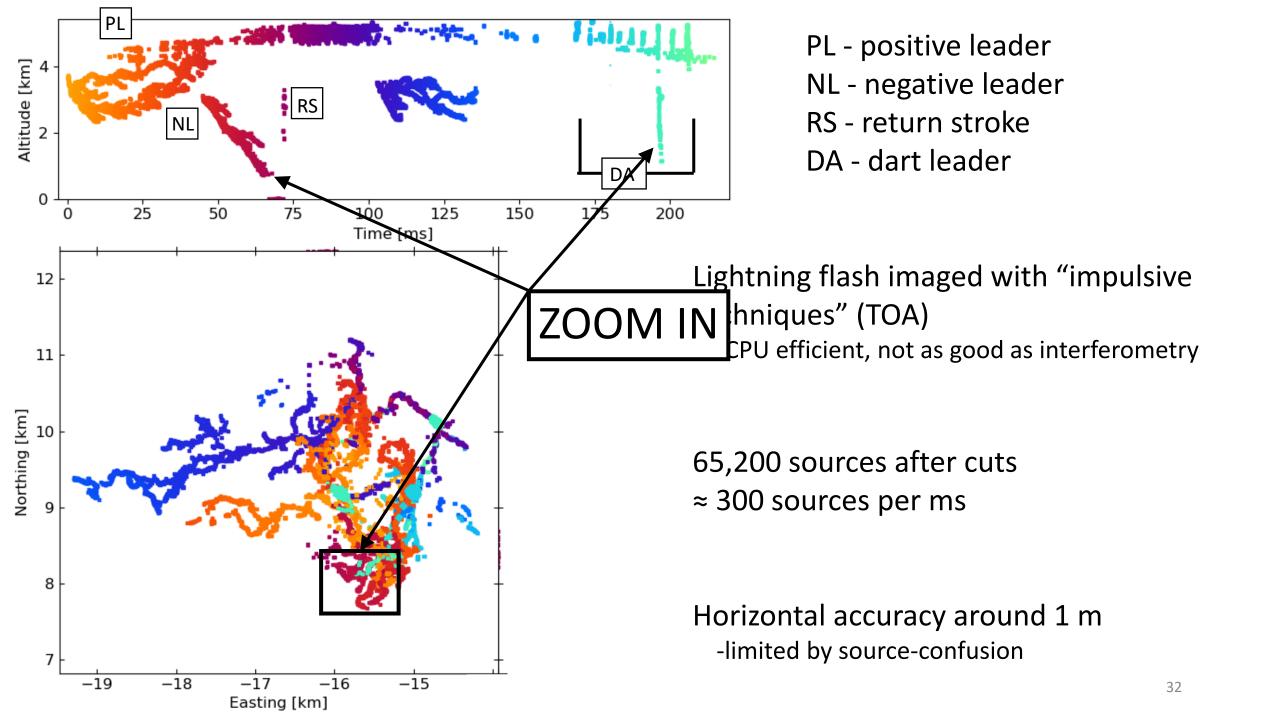




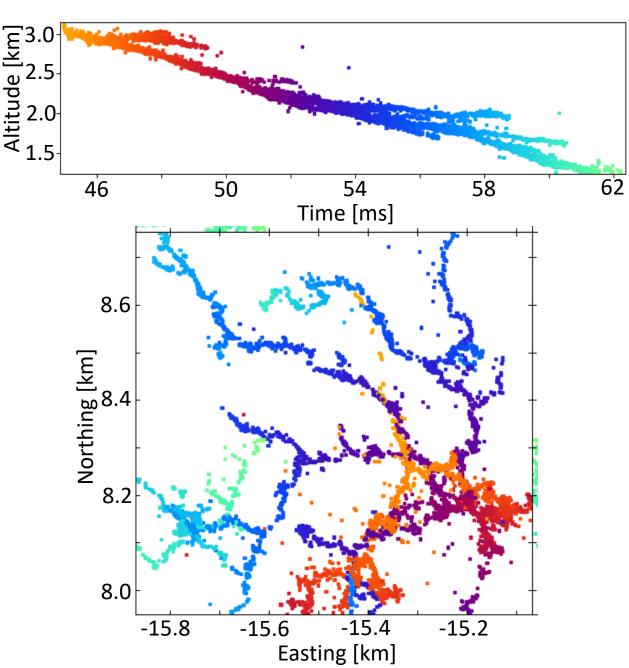




Current work: Dart Leaders



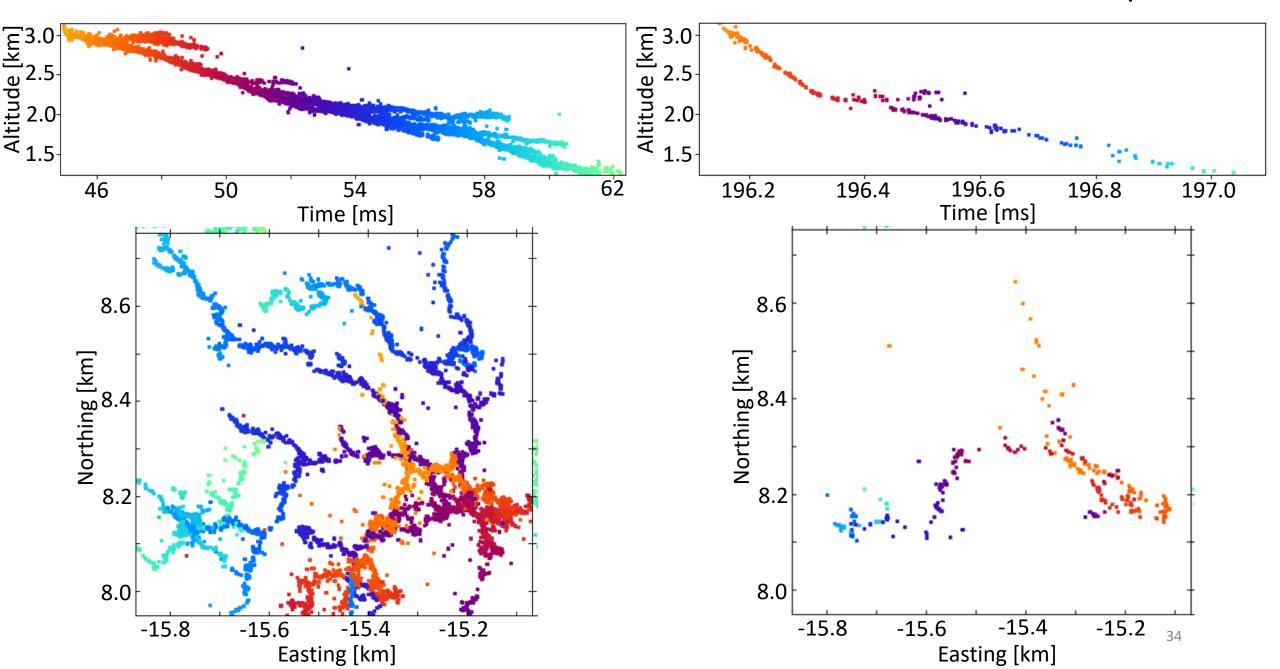
Negative Leader



Dart Leader over same spot

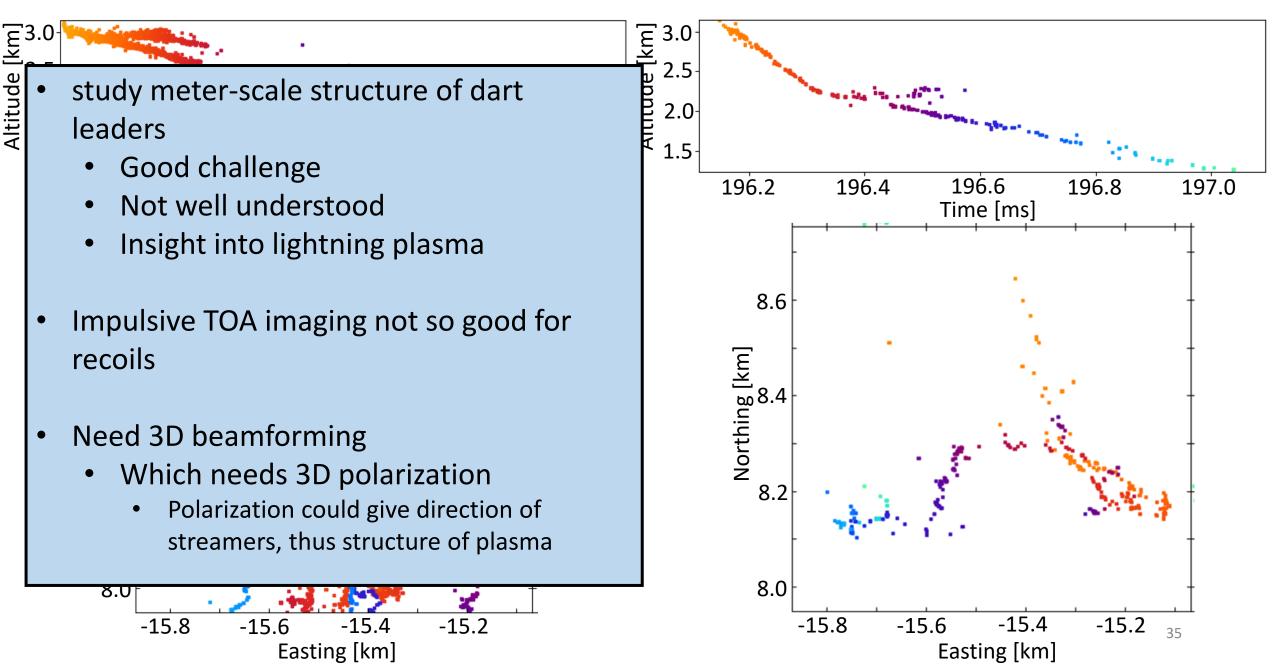
Negative Leader

Dart Leader over same spot

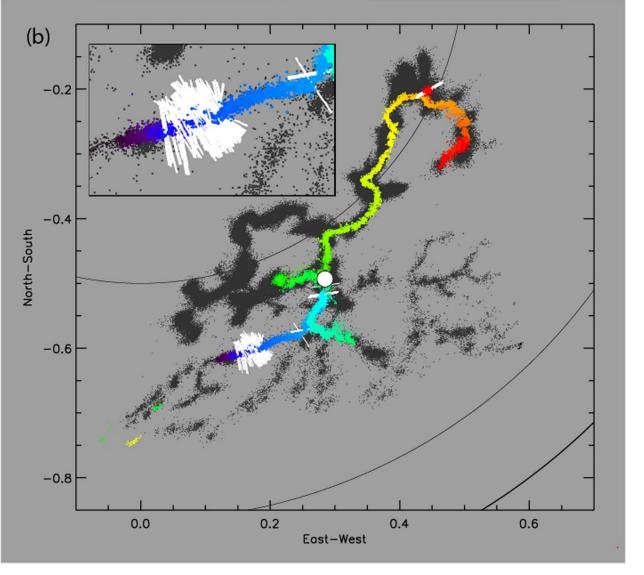


Negative Leader

Dart Leader over same spot



A Previous Work



X.M Shao (2018), JGR-Atmospheres

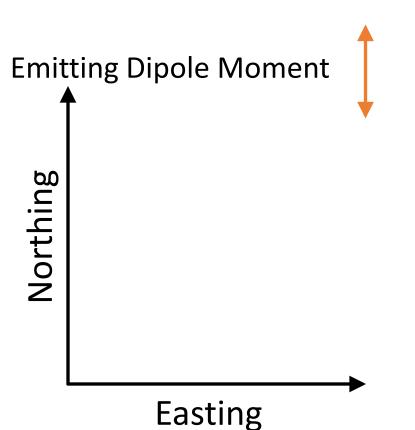
- Studied recoil leaders with 2D radio imaging
 - 1 µs integration time
 - 100 m resolution
- Extracted 2D polarization
- Black dots are whole flash
- Colored dots are a dart leader
- White lines is linear polarization direction
- RESULT:
 - Polarization is perpendicular to propagation
 - Implies dart strongly interacts with corona sheath around plasma channel

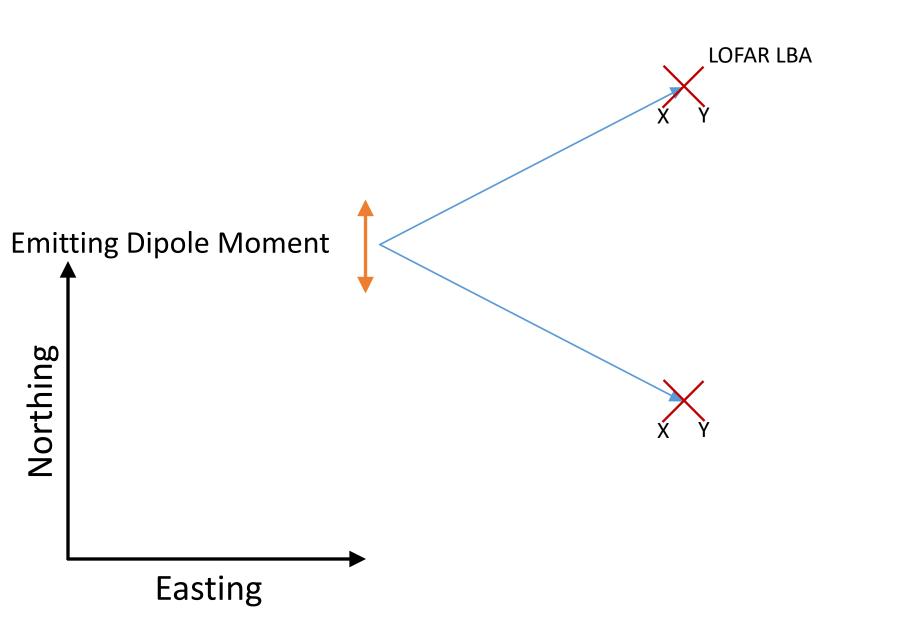
3D Beamforming

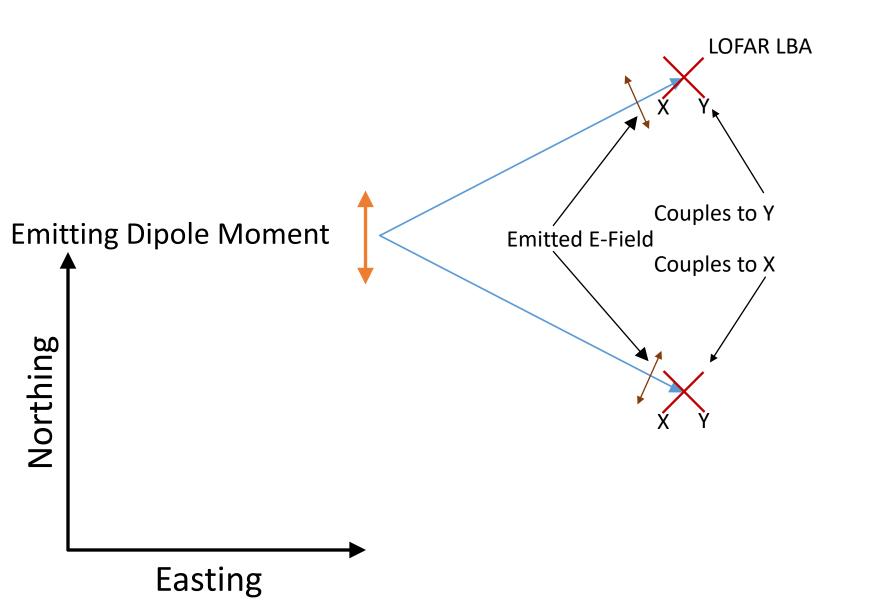
Impulsive imaging techniques struggle with interfering sources, which we want to improve by using beamforming

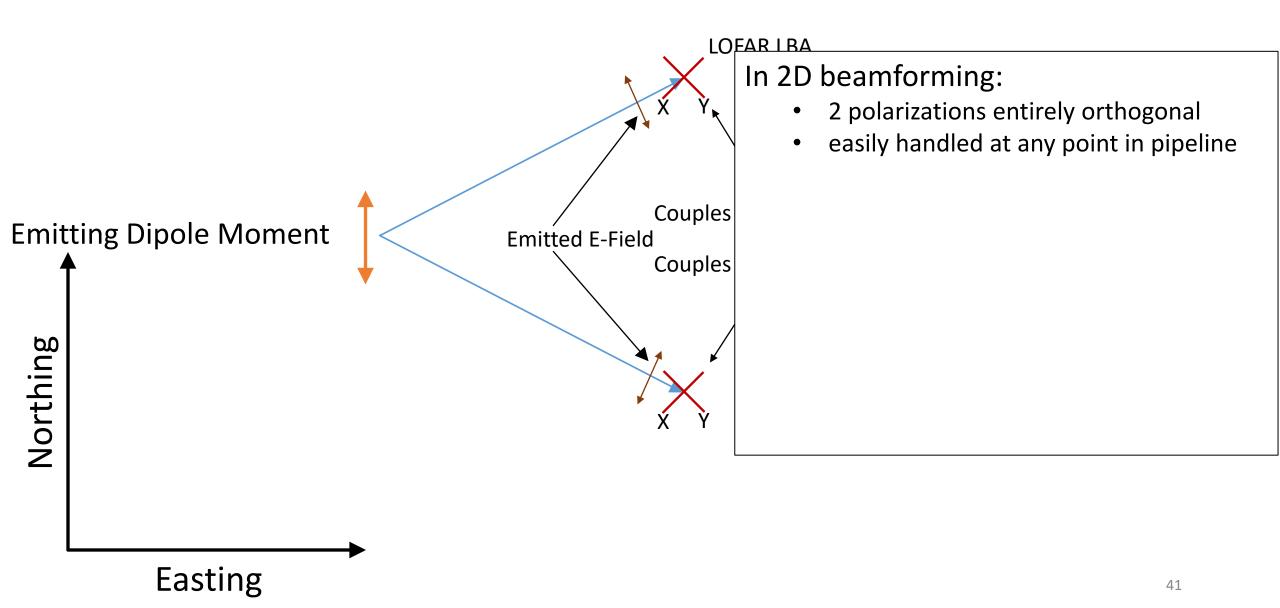
Challenges for 3D:

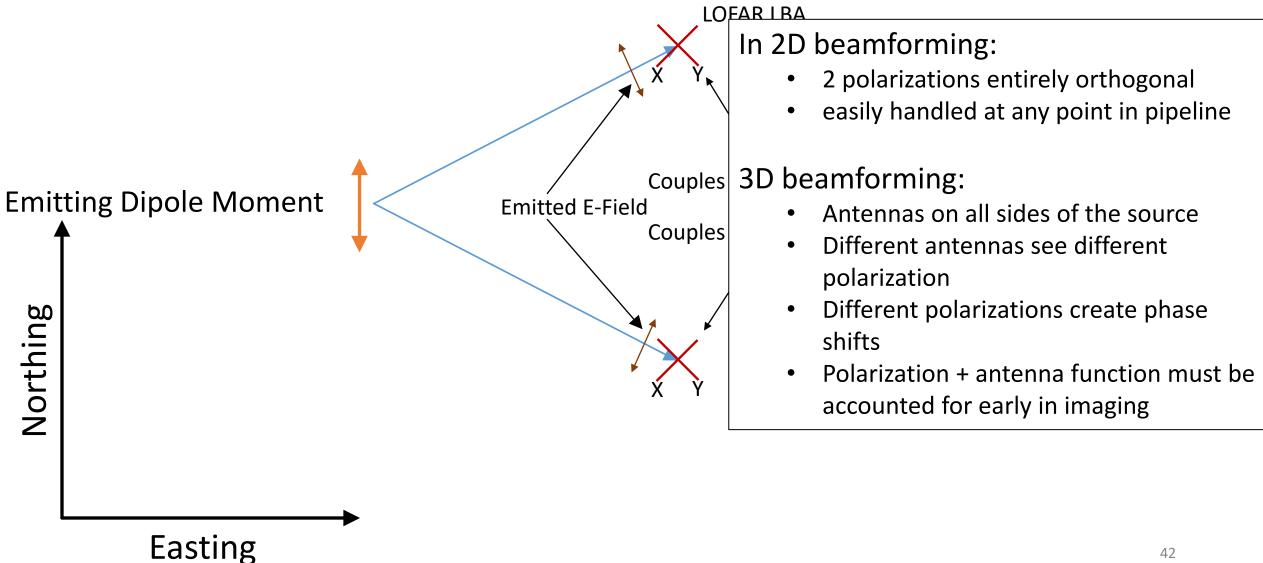
- Curved wave front
 - Solution: replace plane wave model with 3D point sources
- Curse of dimensionality (3D+time vs 2D)
 - Want integration times of 100 ns
 - Solution: only image small sections, not entire flash
- Polarization
 - This one is hard



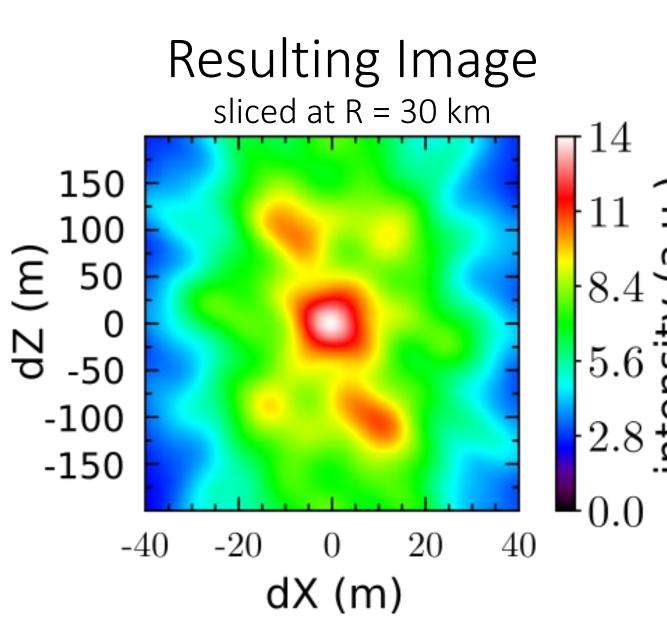








NEW3D Lightning Beamforming

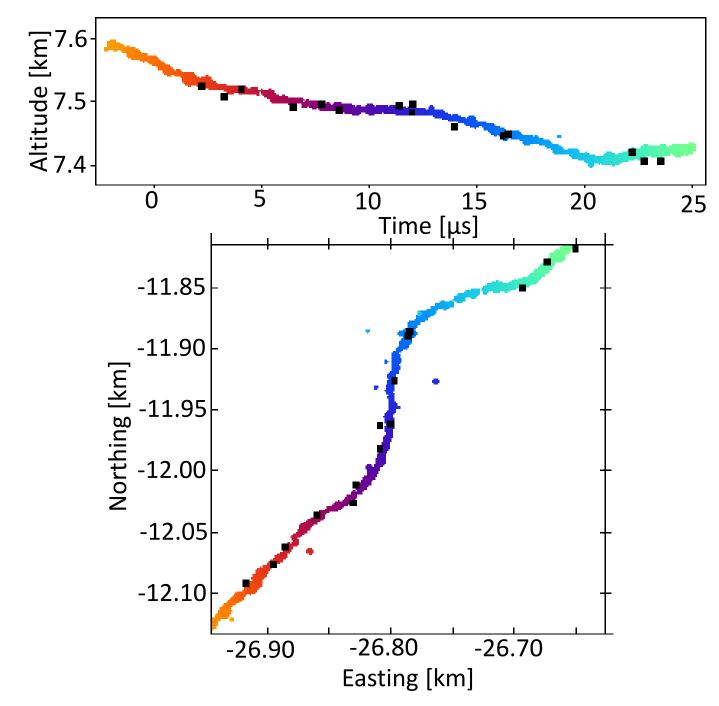


• Accounts for full 3D polarization

- Full 3D Image
 - Integrate over 100 ns

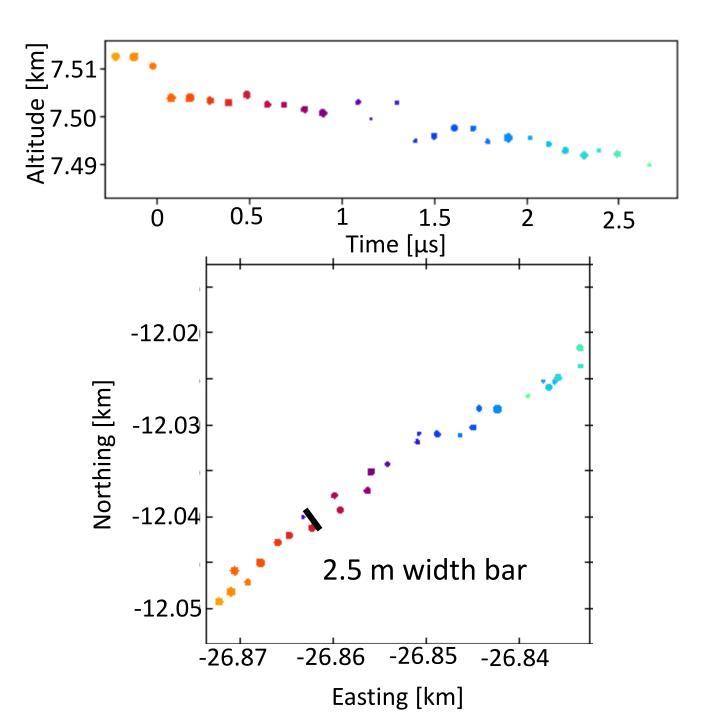
- Pick location of max. intensity as location of source
 - No deconvolution yet

• Repeat this over many time-slices

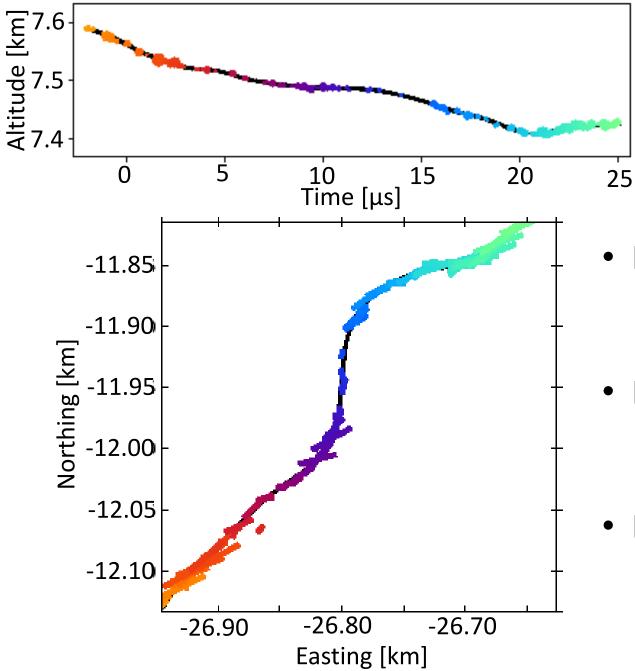


Small Section of Dart

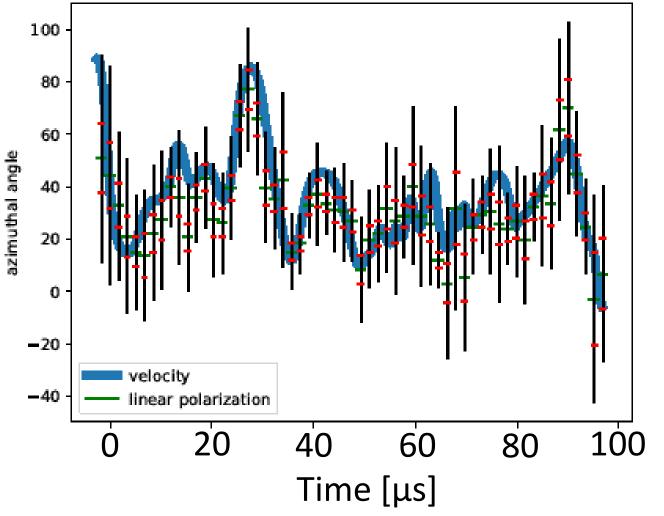
- Black data are previous imager
- Colored dots are new Interferometry
- One of cleanest events
- Tremendous detail
 - Physics results: very smooth



- Zoom-in to previous plot
- VHF-emitting region is about 2 m wide
 - Due to resolution
- Lightning channel have
 - Thin hot core (cm in size)
 - Cold corona sheath (20 m size)
- Thus VHF from hot core
 - Surprising!!



- Lines show linear polarization
 - Of strongest pulses
- Polarization is parallel leader track
 - Different from previous results
- But this type of plot can be misleading
 - Better statistical analysis next slide

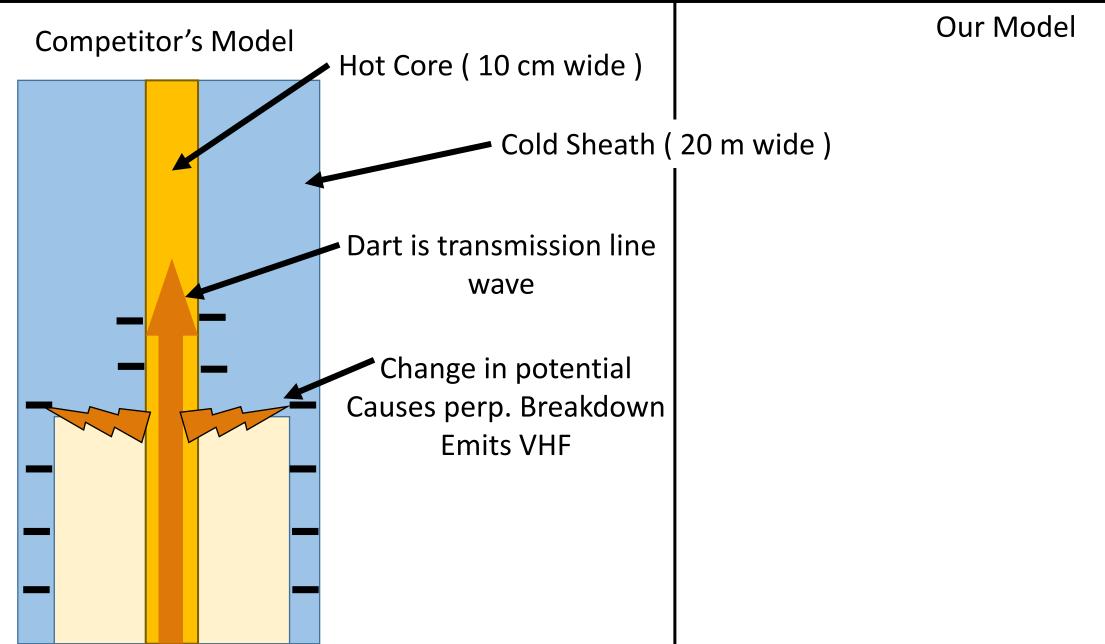


• Blue line : direction of propagation

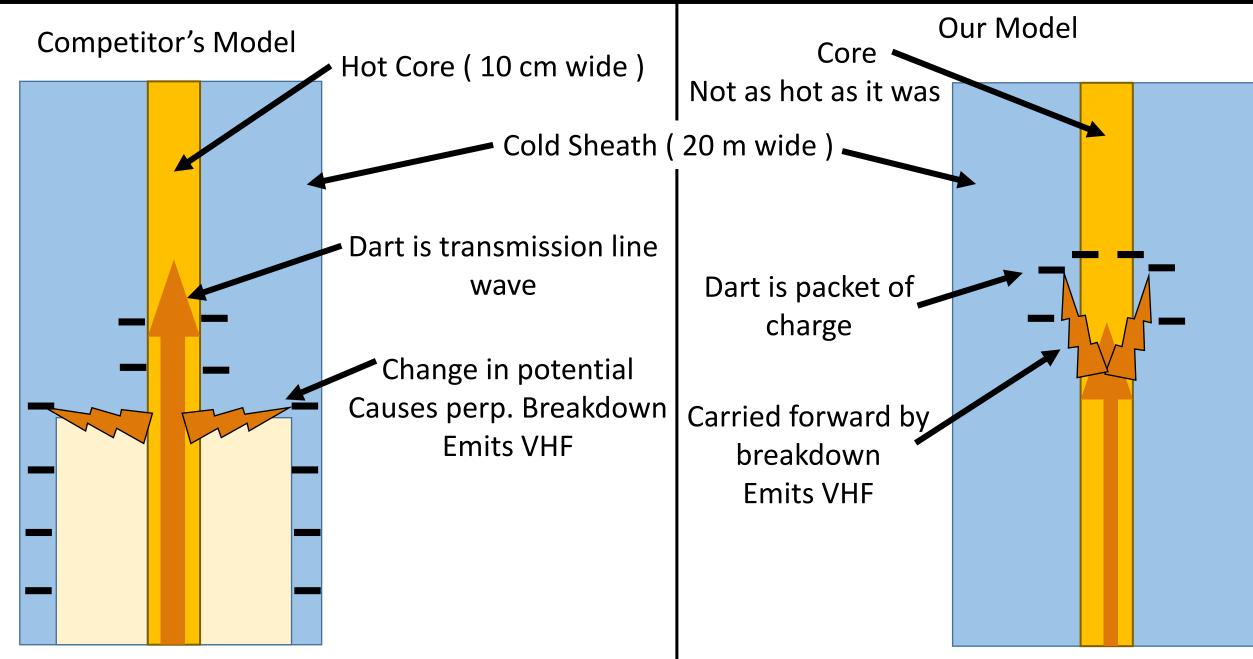
- Black bars : linear polarization
 - Center shows average
 - Width shows standard deviation

- Confirms previous plot
 - Polarization follows propagation

Dart Conclusions



Dart Conclusions



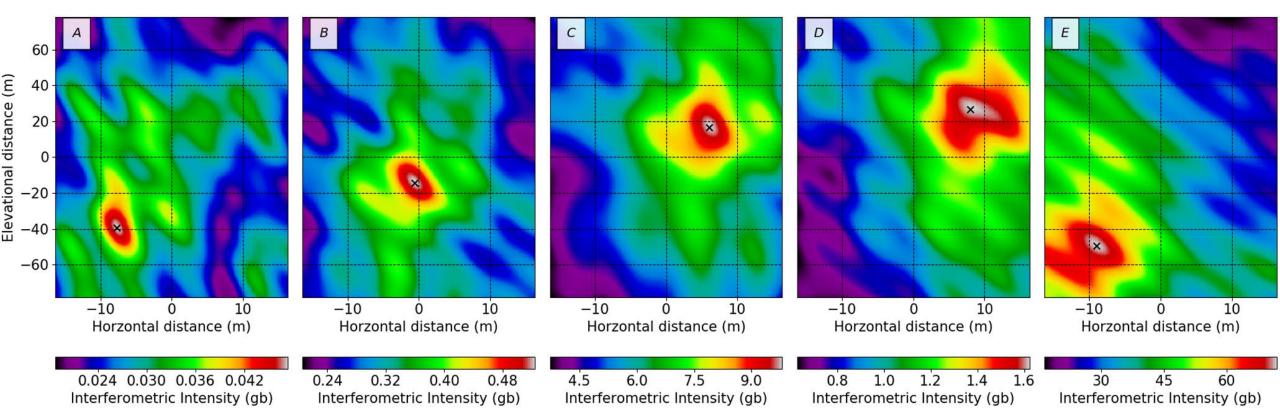
Initiation

New interferometry allows imaging weakly emitting processes

Initiation

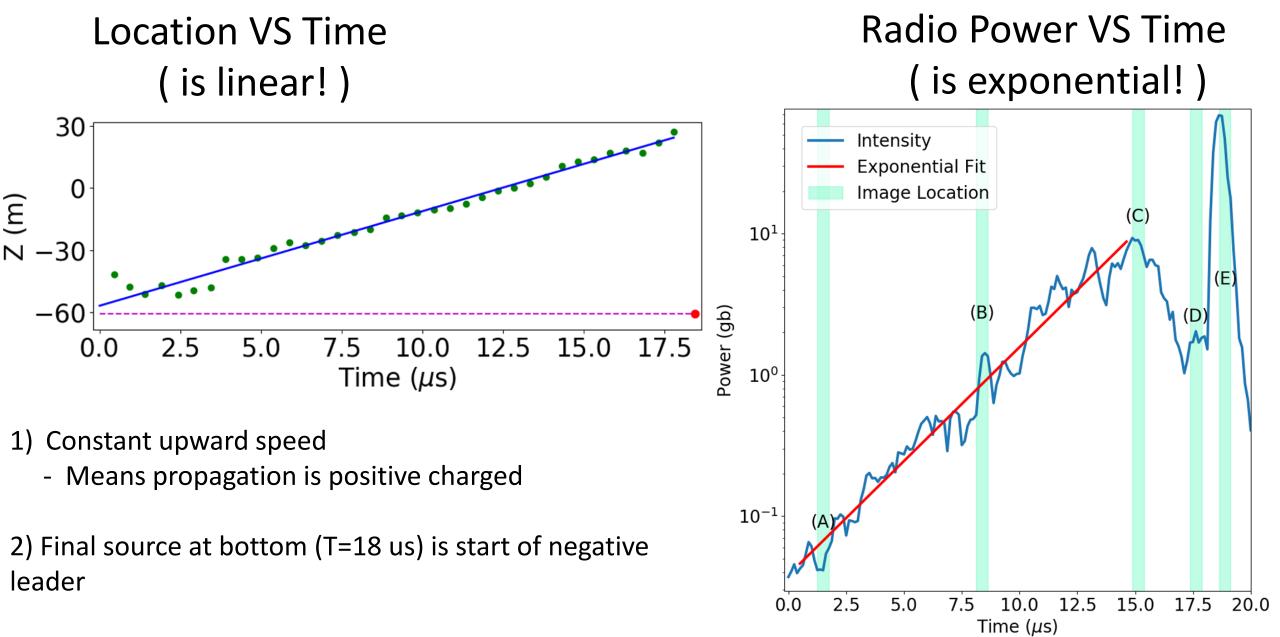
New interferometry allows imaging weakly emitting processes

First Image of first process in lightning

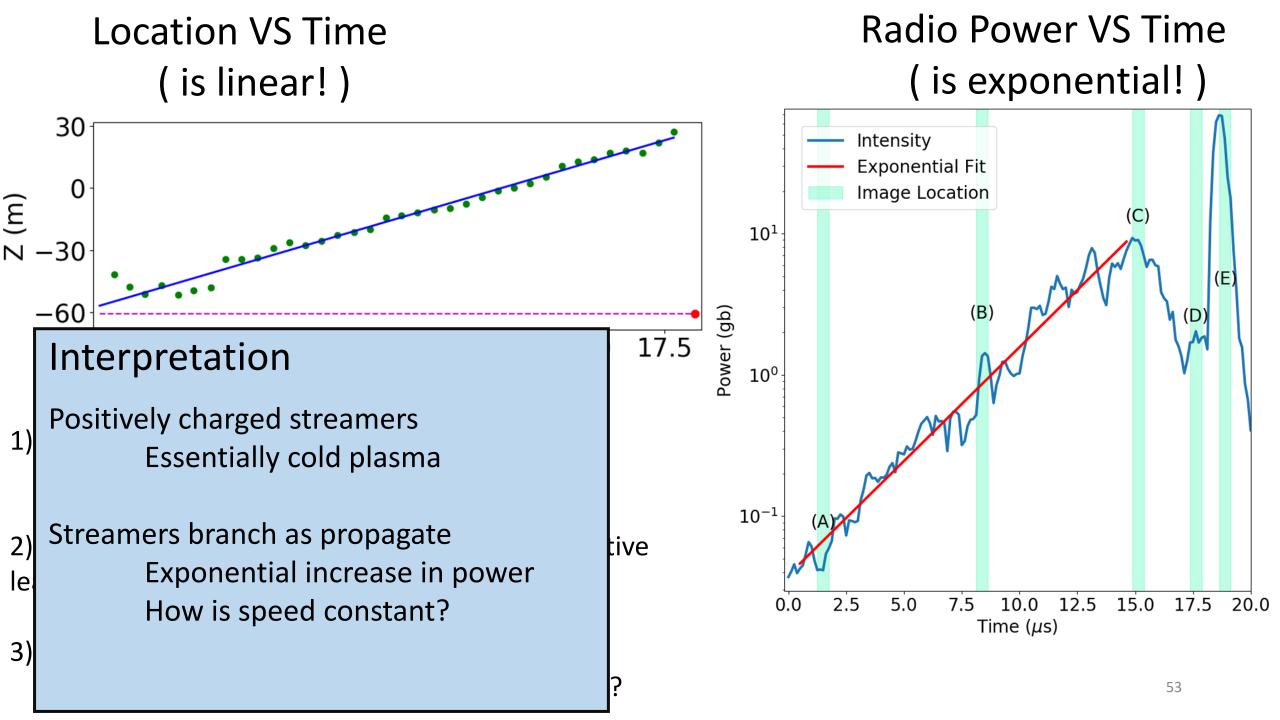


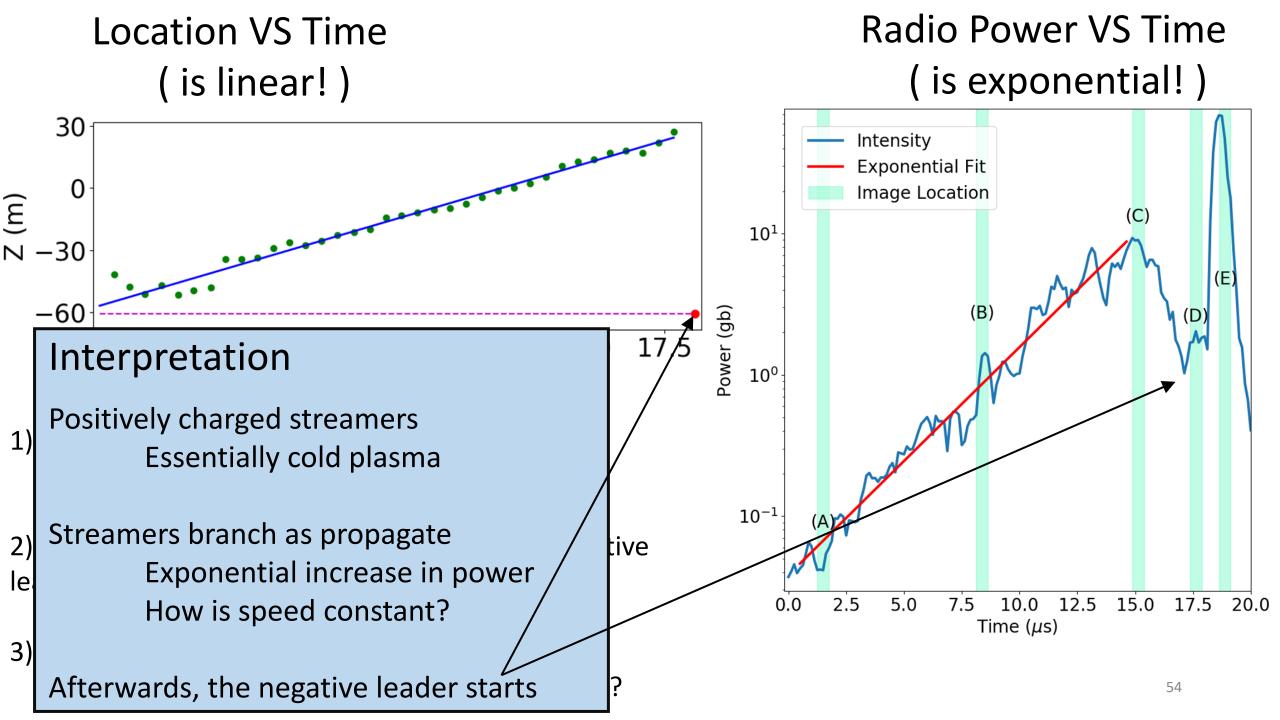
A) Process goes UP then back DOWN

B) VHF power increases exponentially



- 3) Power increase exponential
 - How is speed constant but power exponential?





Conclusion

- LOFAR is probing fundamental lightning questions
 - Discovering new lightning phenomena (needles)
 - Exploring how lightning propagates (dart leaders)
 - Imaging lightning initiation

EXTRA



• For each pulse to locate:

1) Use 1 LOFAR station

• Get initial direction



• For each pulse to locate:

1) Use 1 LOFAR station

• Get initial direction

2) Select next station

- Guess which pulse
- Improve location guess



• For each pulse to locate:

Use 1 LOFAR station
Get initial direction

2) Select next station

- Guess which pulse
- Improve location guess

3) Repeat for all LOFAR stations



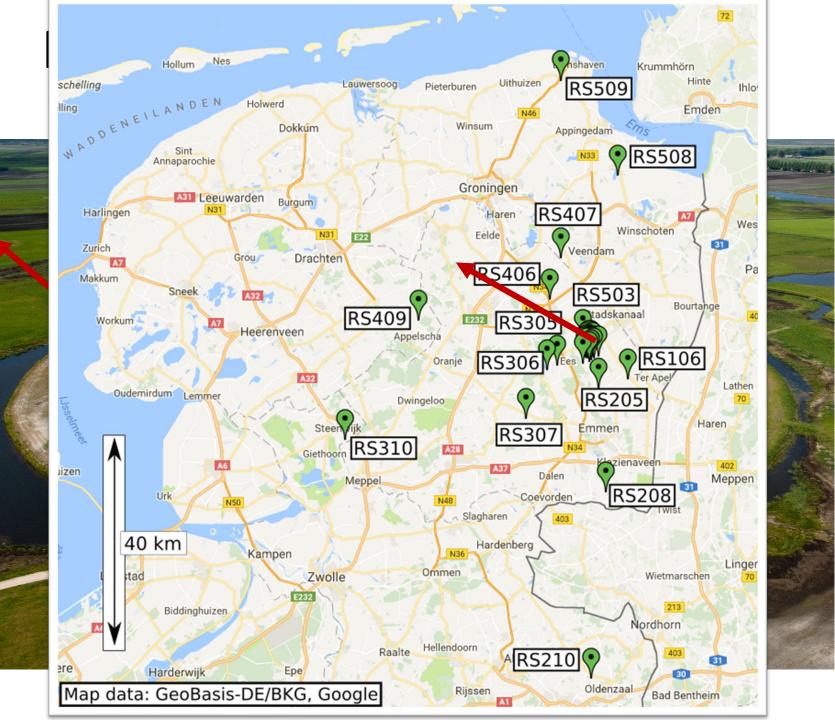
• For each pulse to locate:

1) Use 1 LOFAR stationGet initial direction

2) Select next station

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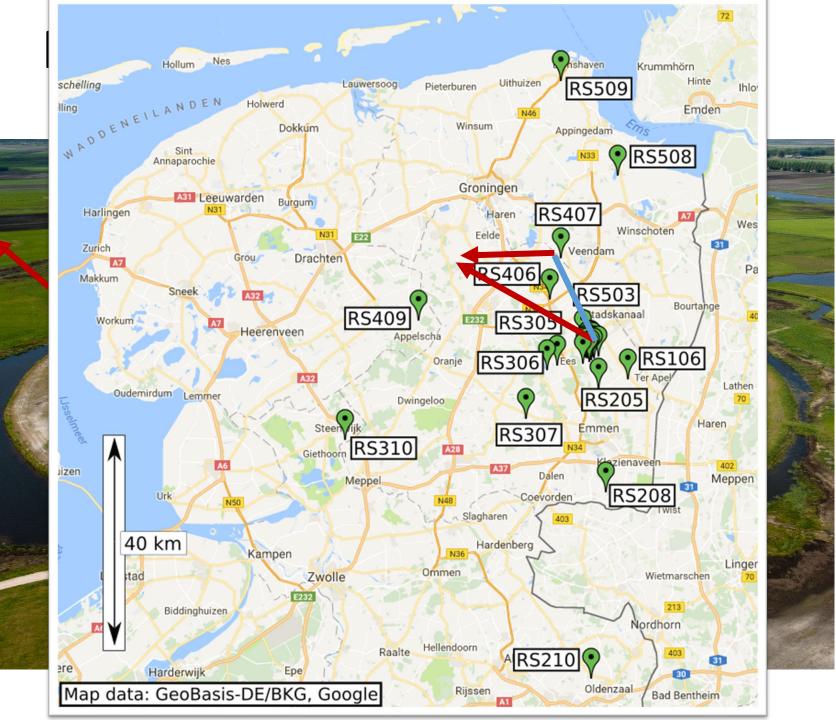
1) Use 1 LOFAR station

Get initial direction

2) Select next station

- Guess which pulse
- Improve location guess

3) Repeat for all LOFAR stations



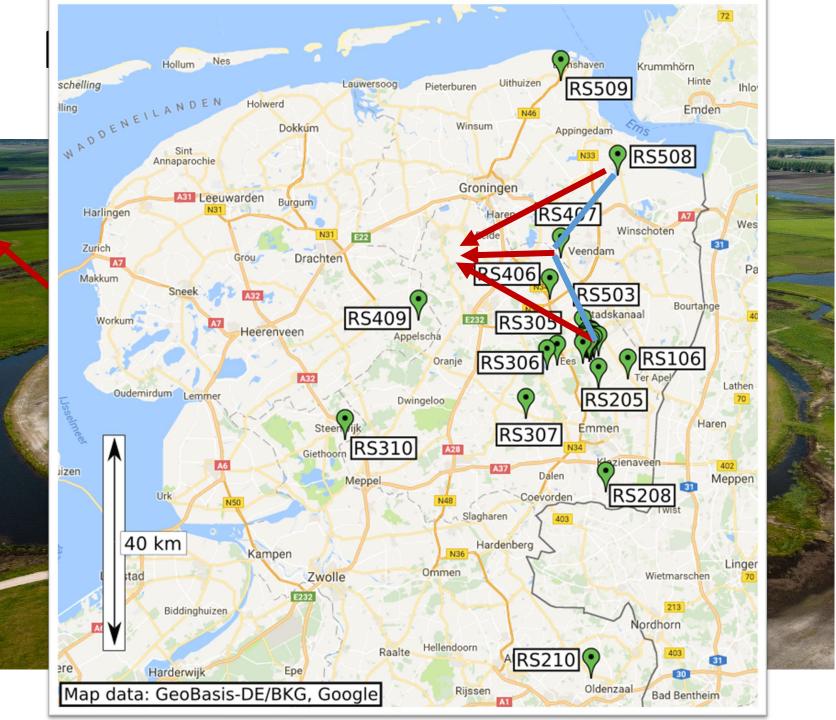
1) Use 1 LOFAR station

Get initial direction

2) Select next station

- Guess which pulse
- Improve location guess

3) Repeat for all LOFAR stations



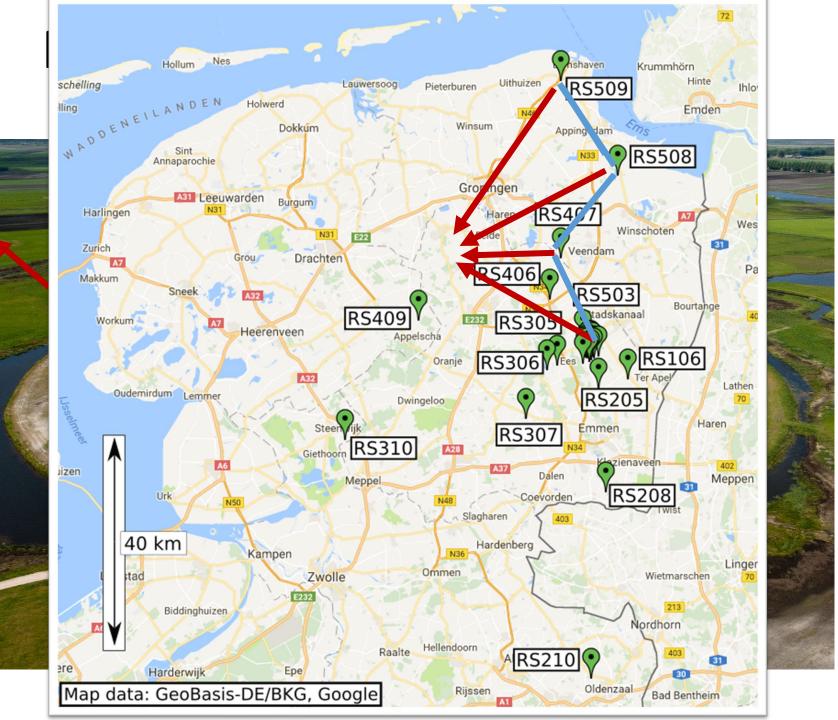
1) Use 1 LOFAR station

Get initial direction

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- Guess which pulse
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3) Repeat for all LOFAR stations



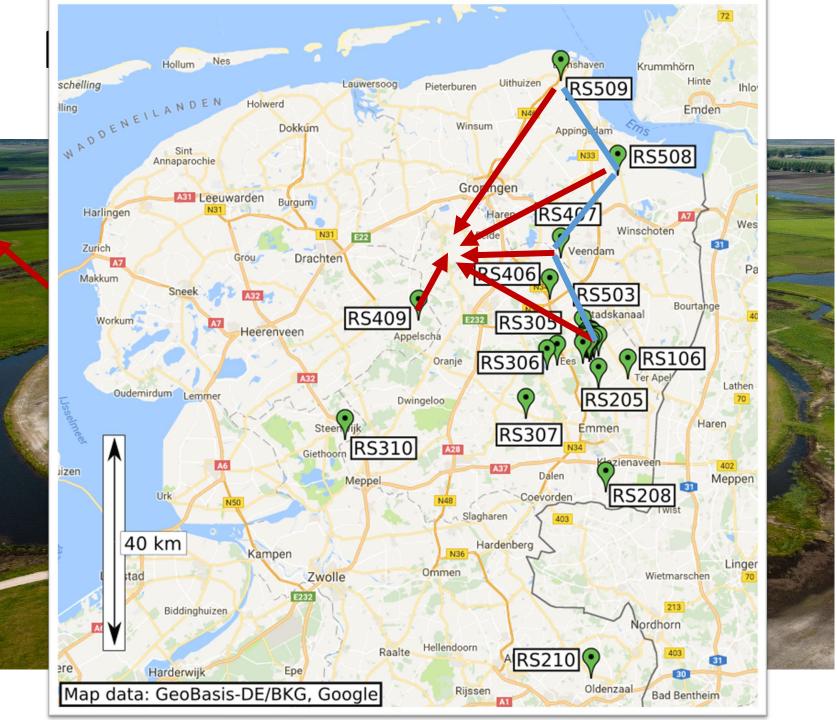
1) Use 1 LOFAR station

Get initial direction

2) Select next station

- Guess which pulse
- Improve location guess

3) Repeat for all LOFAR stations



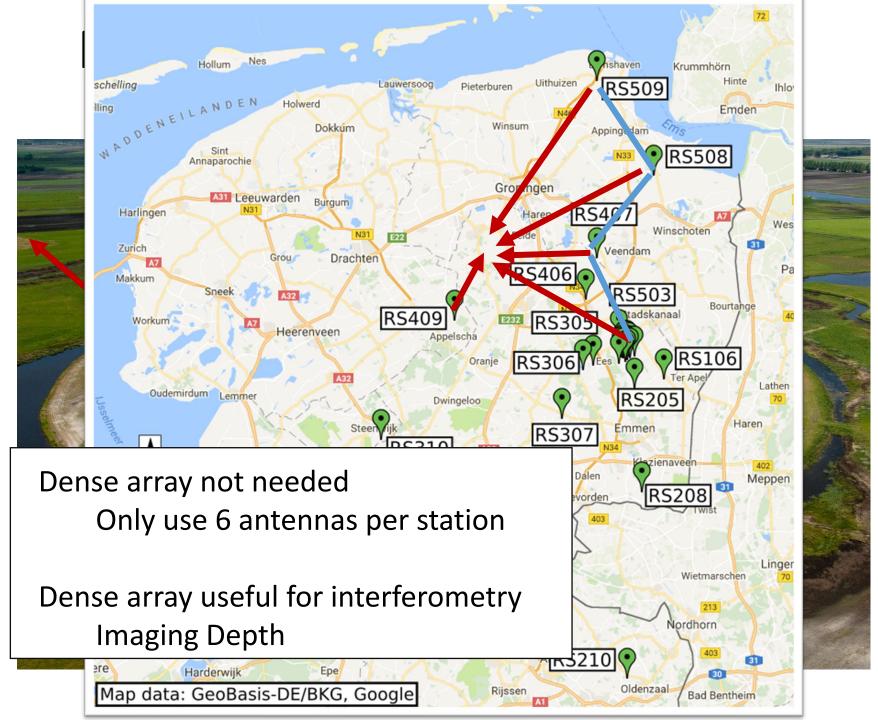
1) Use 1 LOFAR station

Get initial direction

2) Select next station

- Guess which pulse
- Improve location guess

3) Repeat for all LOFAR stations



1) Use 1 LOFAR station

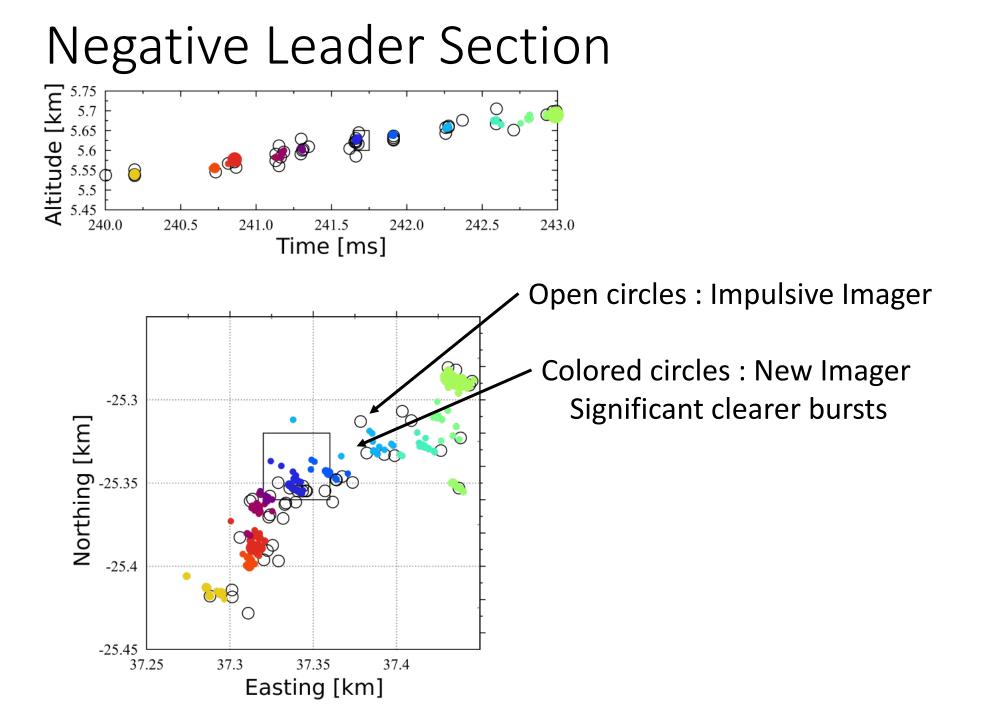
Get initial direction

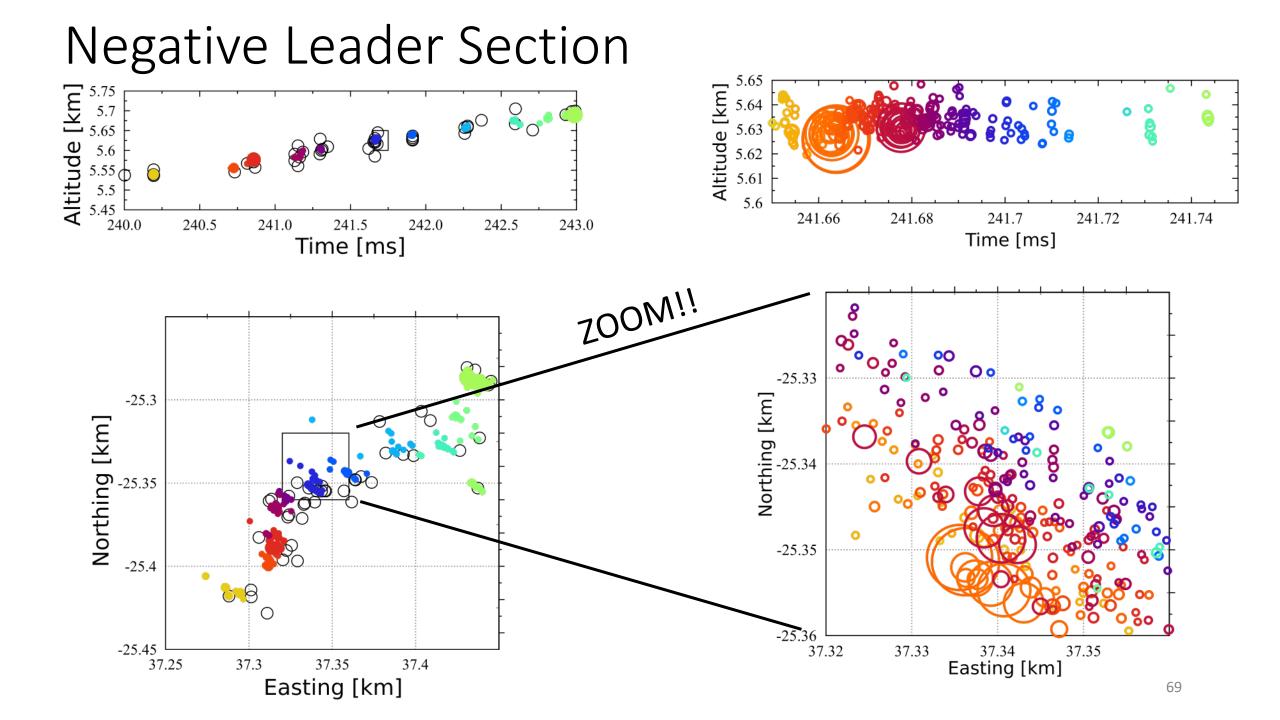
2) Select next station

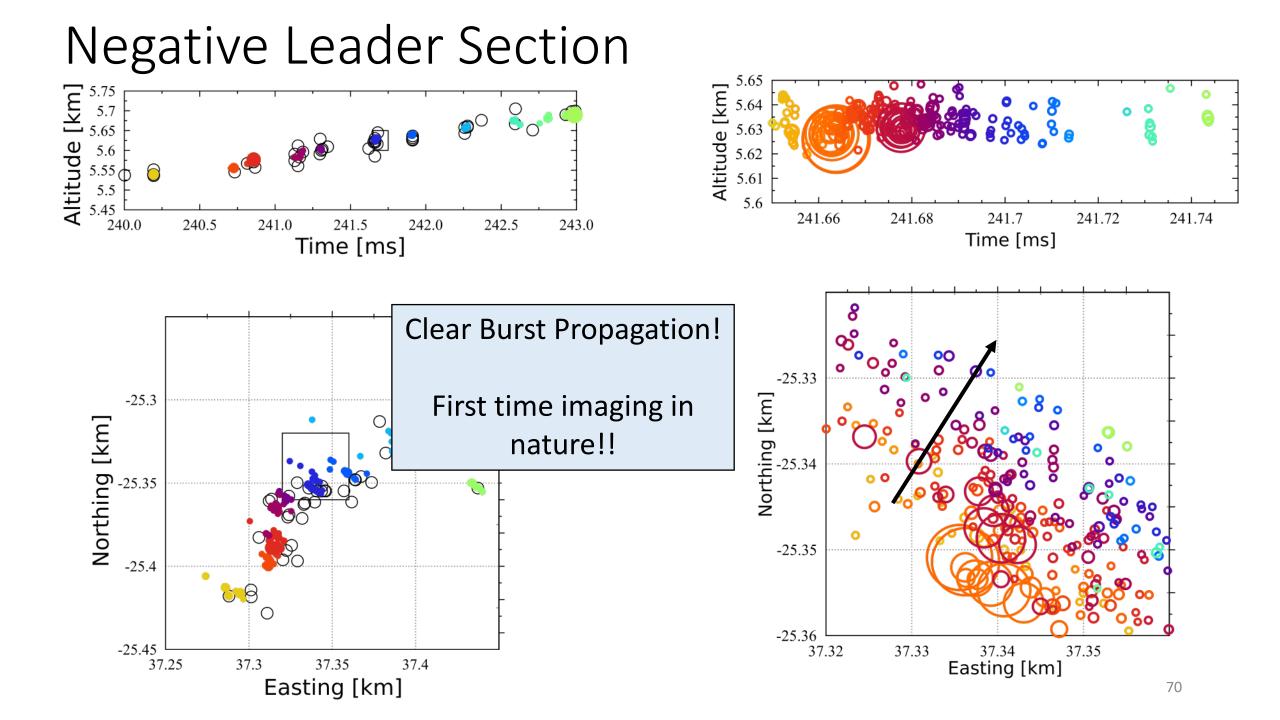
- Guess which pulse
- Improve location guess

3) Repeat for all LOFAR stations

Negative Leader







Polarization problem