

PROBING THE HIGH-ENERGY UNIVERSE:

DETECTING THE HIGHEST ENERGY COSMIC MESSENGERS – BELGIAN CONTRIBUTIONS

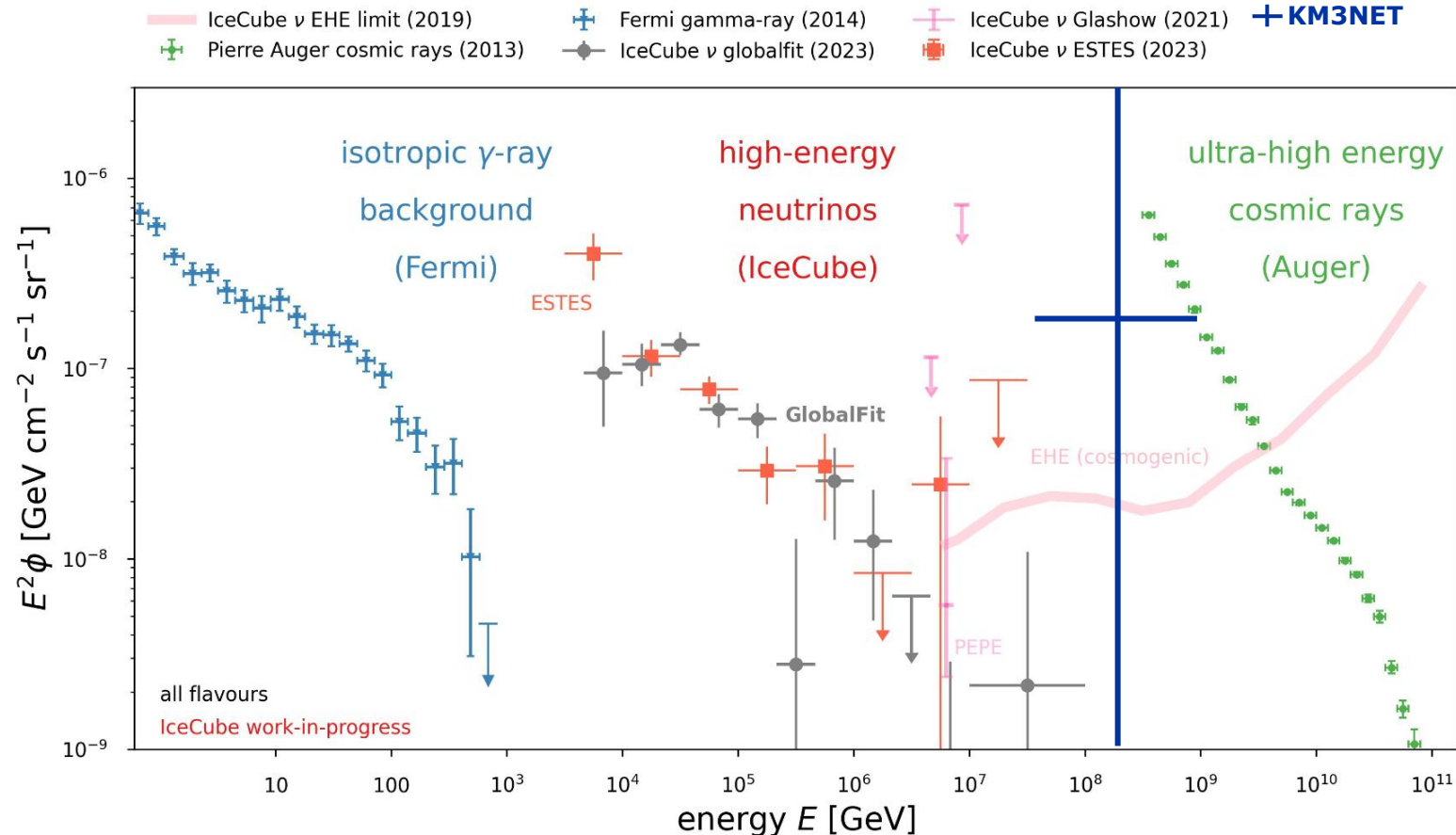
KRIJN D DE VRIES

VRIJE UNIVERSITEIT BRUSSEL
INTERUNIVERSITY INSTITUTE FOR HIGH ENERGIES



HIGH-ENERGY COSMIC MESSENGERS

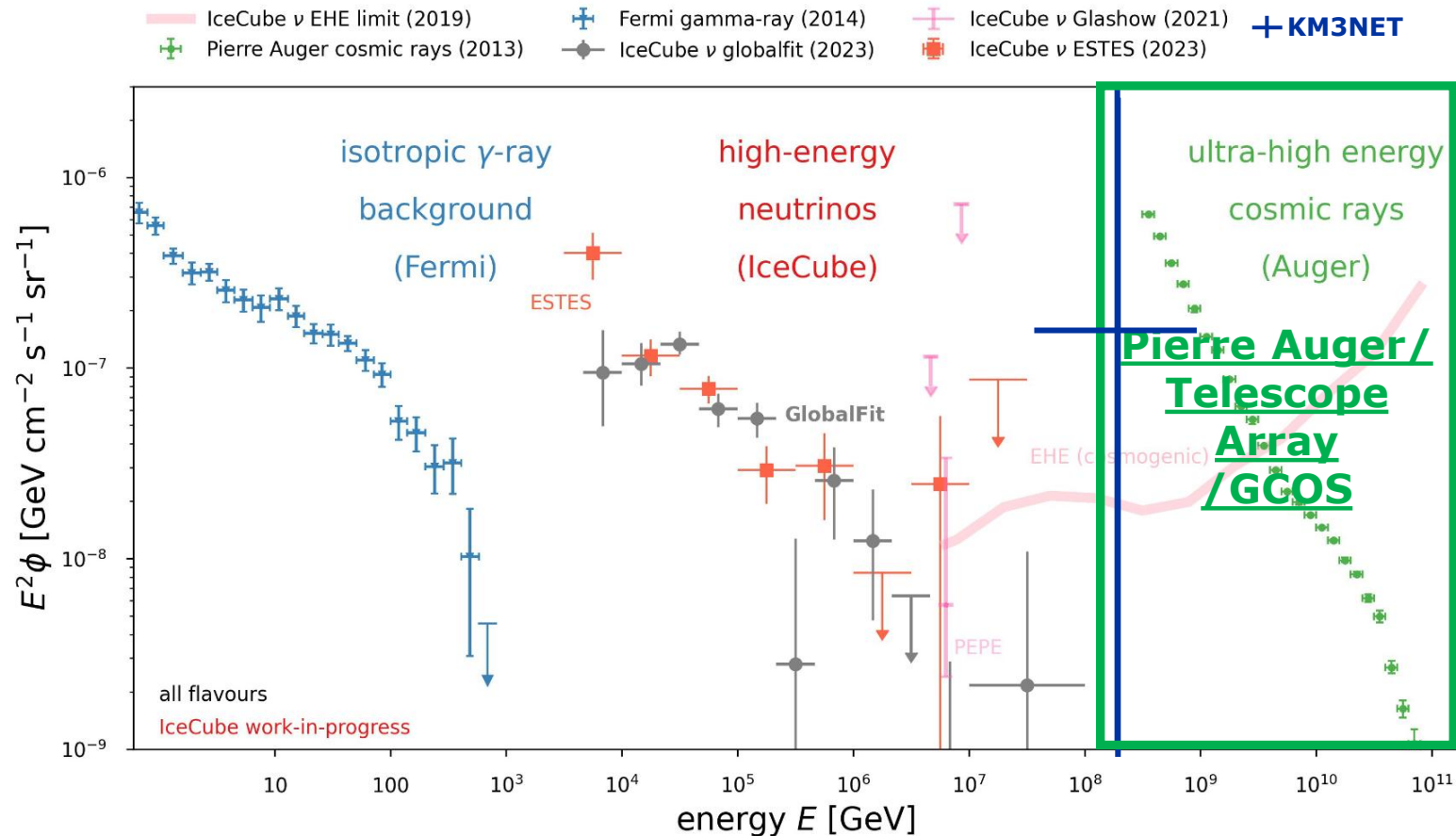
GAMMA-RAYS, NEUTRINOS AND COSMIC RAYS



C. Haack, IceCube Collaboration, The 43rd International Symposium on Physics in Collision PIC 2024.

HIGH-ENERGY COSMIC MESSENGERS

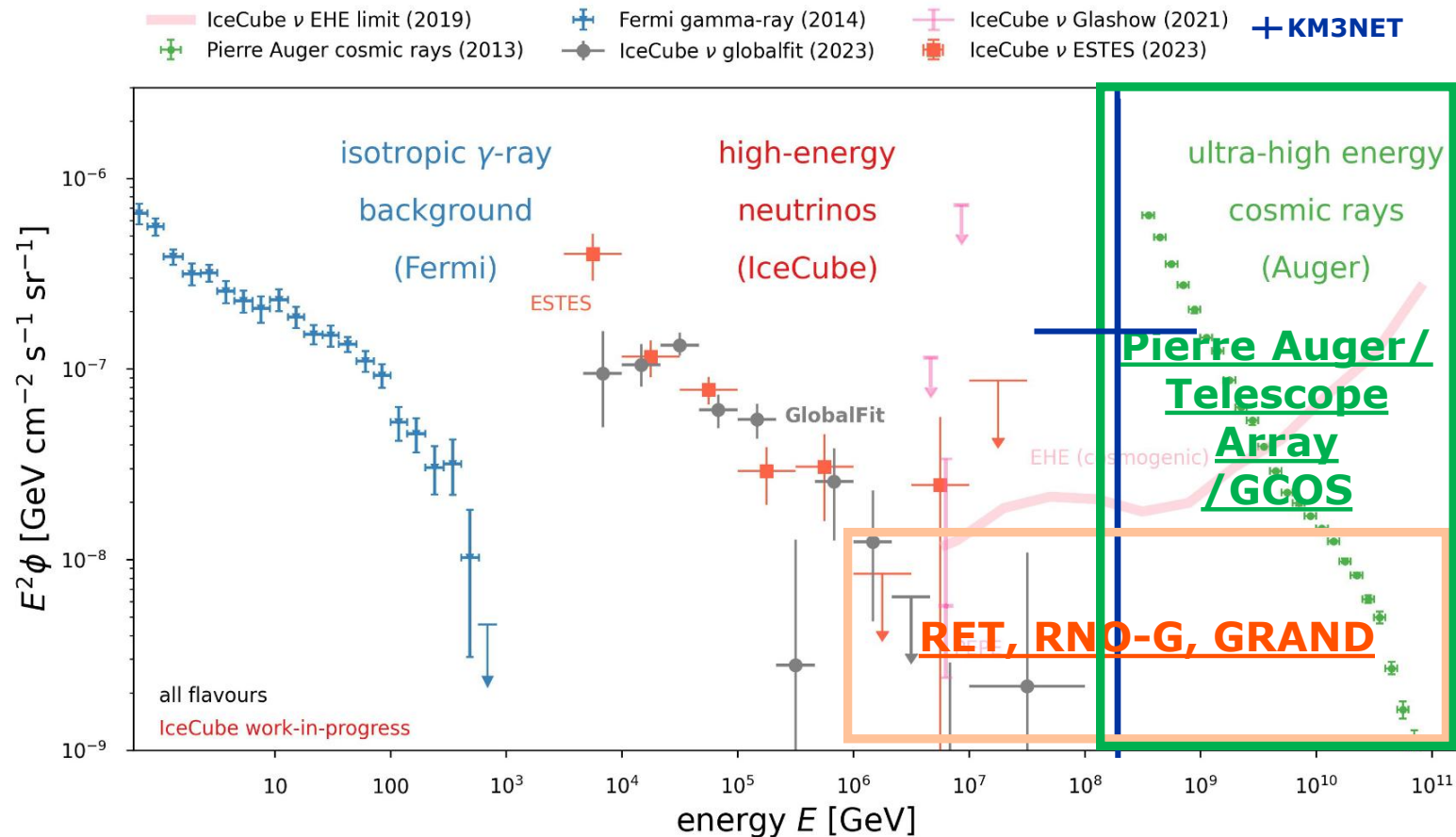
GAMMA-RAYS, NEUTRINOS AND COSMIC RAYS



C. Haack, IceCube Collaboration, The 43rd International Symposium on Physics in Collision PIC 2024.

HIGH-ENERGY COSMIC MESSENGERS

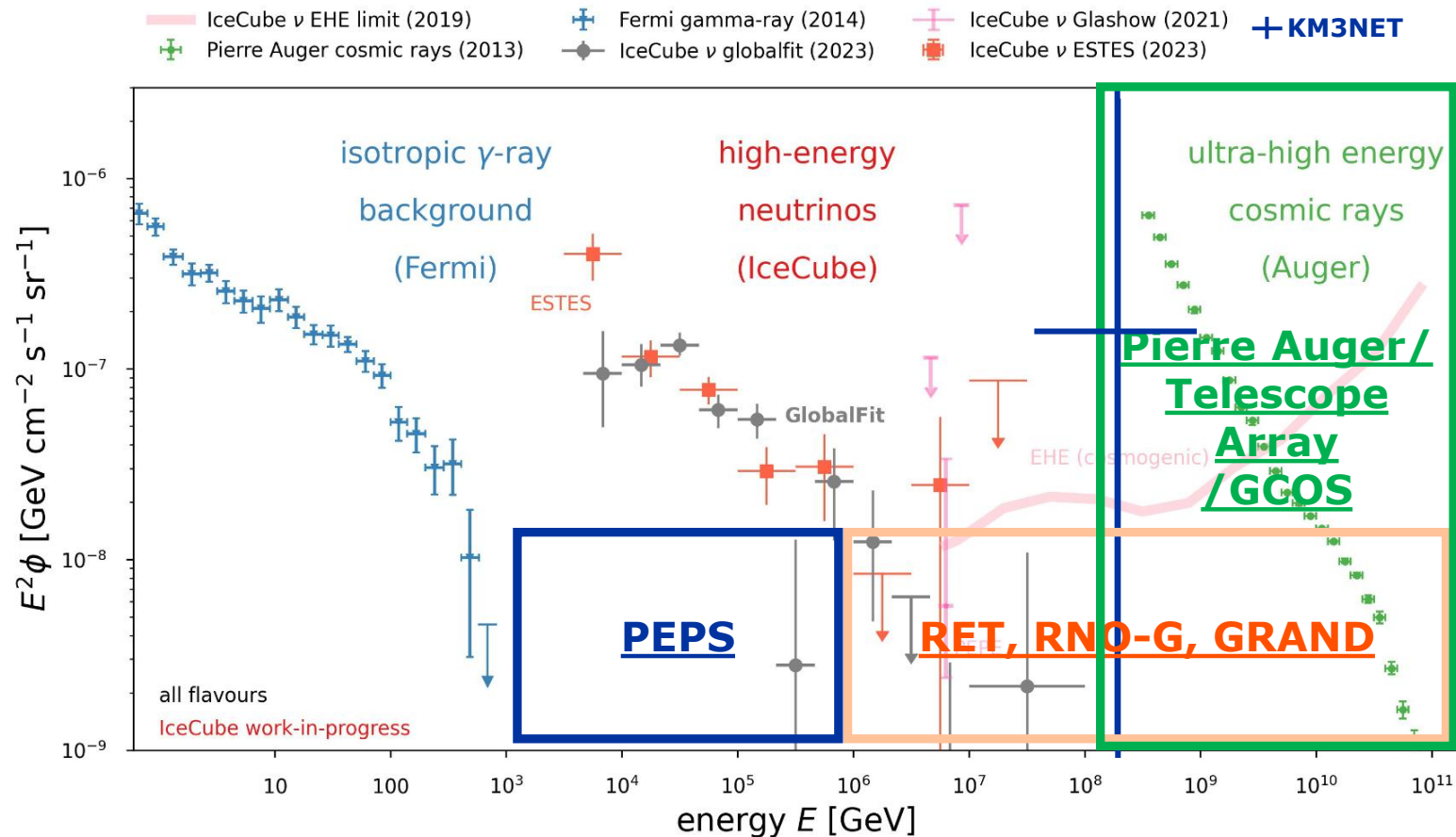
GAMMA-RAYS, NEUTRINOS AND COSMIC RAYS



C. Haack, IceCube Collaboration, The 43rd International Symposium on Physics in Collision PIC 2024.

HIGH-ENERGY COSMIC MESSENGERS

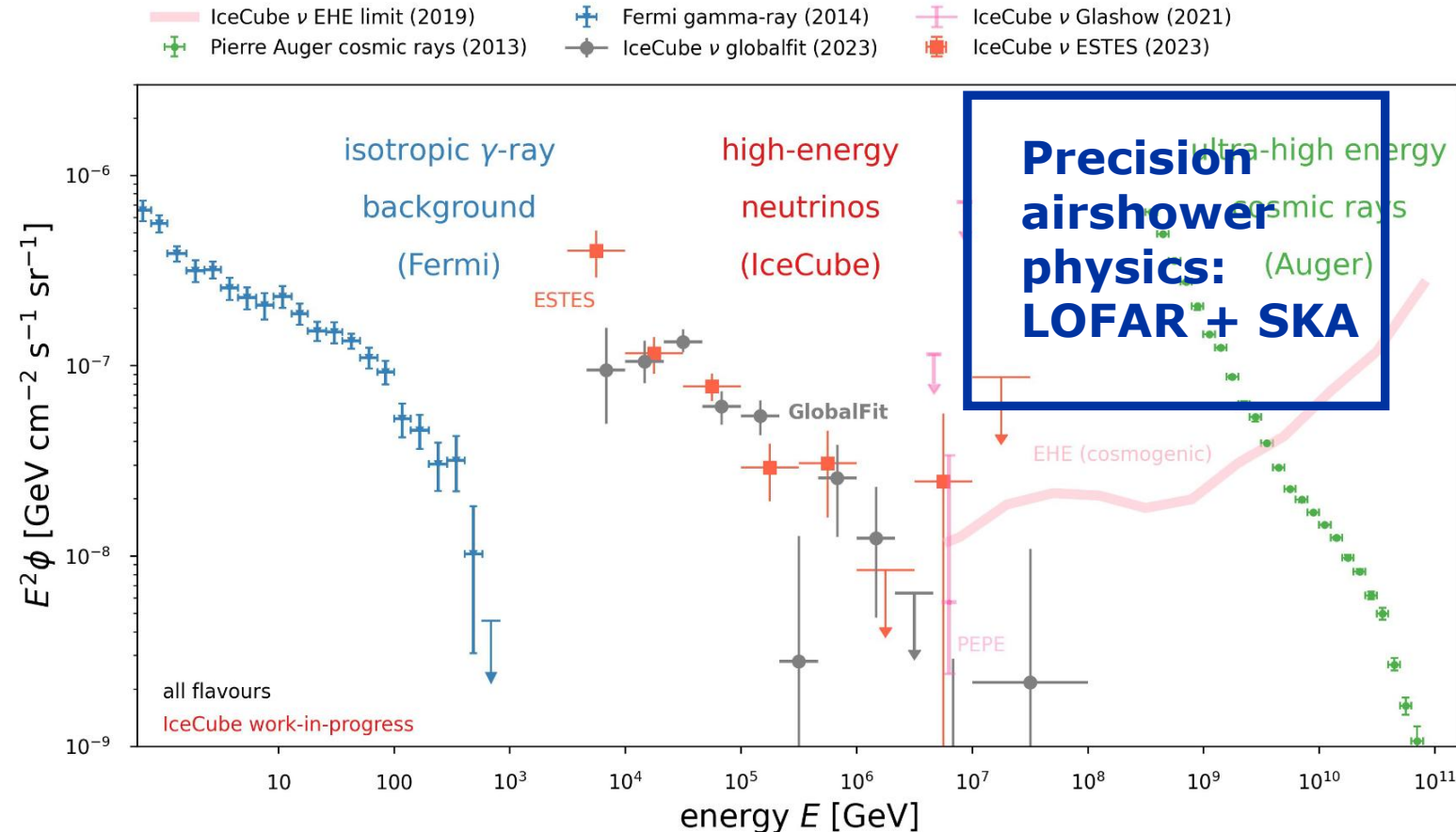
GAMMA-RAYS, NEUTRINOS AND COSMIC RAYS



C. Haack, IceCube Collaboration, The 43rd International Symposium on Physics in Collision PIC 2024.

HIGH-ENERGY COSMIC MESSENGERS

GAMMA-RAYS, NEUTRINOS AND COSMIC RAYS



C. Haack, IceCube Collaboration, The 43rd International Symposium on Physics in Collision PIC 2024.

HIGH-ENERGY COSMIC MESSENGERS – CURRENT PROJECTS

COSMIC RAYS – THE PIERRE AUGER /TELESCOPE ARRAY OBSERVATORIES ULB/VUB

Ultra-High-Energy Cosmic Rays

- Particles (p, nuclei) with $E > 10^{18}$ eV
- Pierre Auger and TA have a large accumulated exposure with >20 years of operation
- Fluorescence and water Cherenkov (PA) /scintillator particle detectors (TA).
- Recently added at Auger:
Radio and scintillator components -> AugerPrime.
- Auger MOU assures **operation till at least 2035**

Belgian groups:

- **Auger: I. Maris (ULB), S. Buitink (VUB)**
- **TA: P. Tinyakov (ULB)**

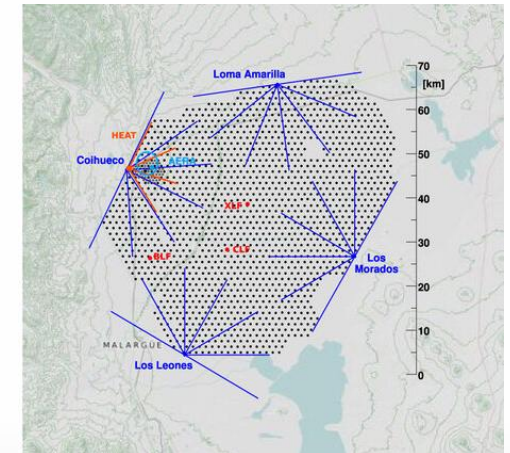


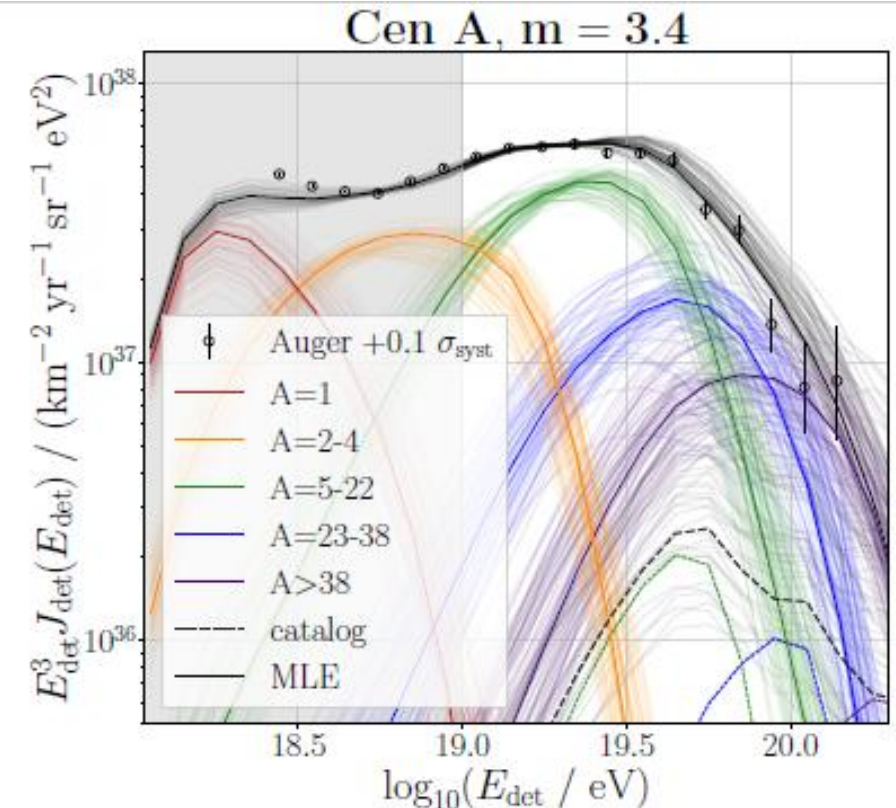
Figure credits: Pierre Auger Collaboration / Telescope Array collaboration

HIGH-ENERGY COSMIC MESSENGERS – CURRENT PROJECTS

COSMIC RAYS – THE PIERRE AUGER OBSERVATORY ULB/VUB-IIHE

Ultra-High-Energy Cosmic Rays

- Precise measurement of the spectral features with discovery of a new feature above 10^{19} eV.
- Establishment of the presence of intermediate-mass nuclei at the highest energies.
- Test of hadronic interaction models at $\sqrt{s} > 14$ TeV and measurement of the p-air cross-section

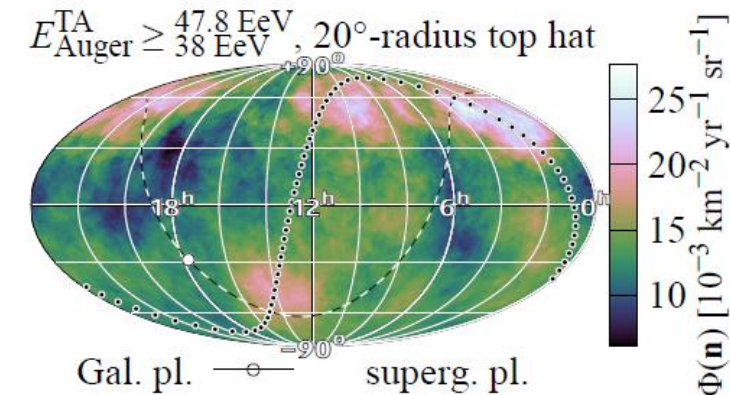
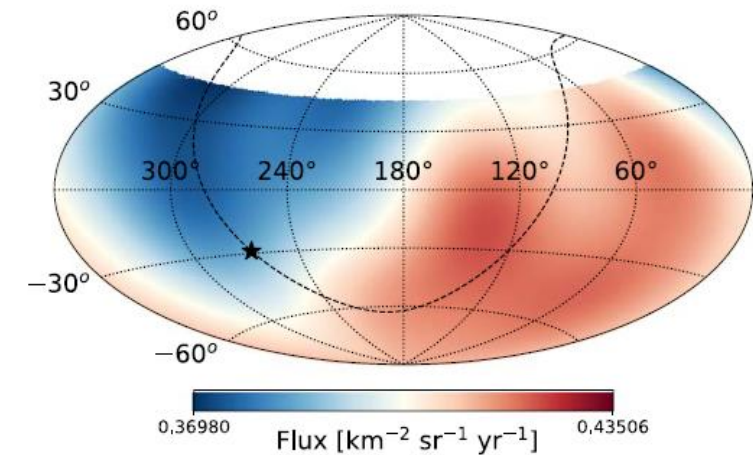


HIGH-ENERGY COSMIC MESSENGERS – CURRENT PROJECTS

COSMIC RAYS – THE PIERRE AUGER OBSERVATORY ULB/VUB-IIHE

The UHECR sky is anisotropic

- The dipole structure ($>7\sigma$ above 10^{18} eV) gives evidence of an **extra-galactic origin**
- Hints ($>3\sigma$) for intermediate scales: Centaurus A, Starburst Galaxies, Perseus Pieces Cluster, TA hotspot

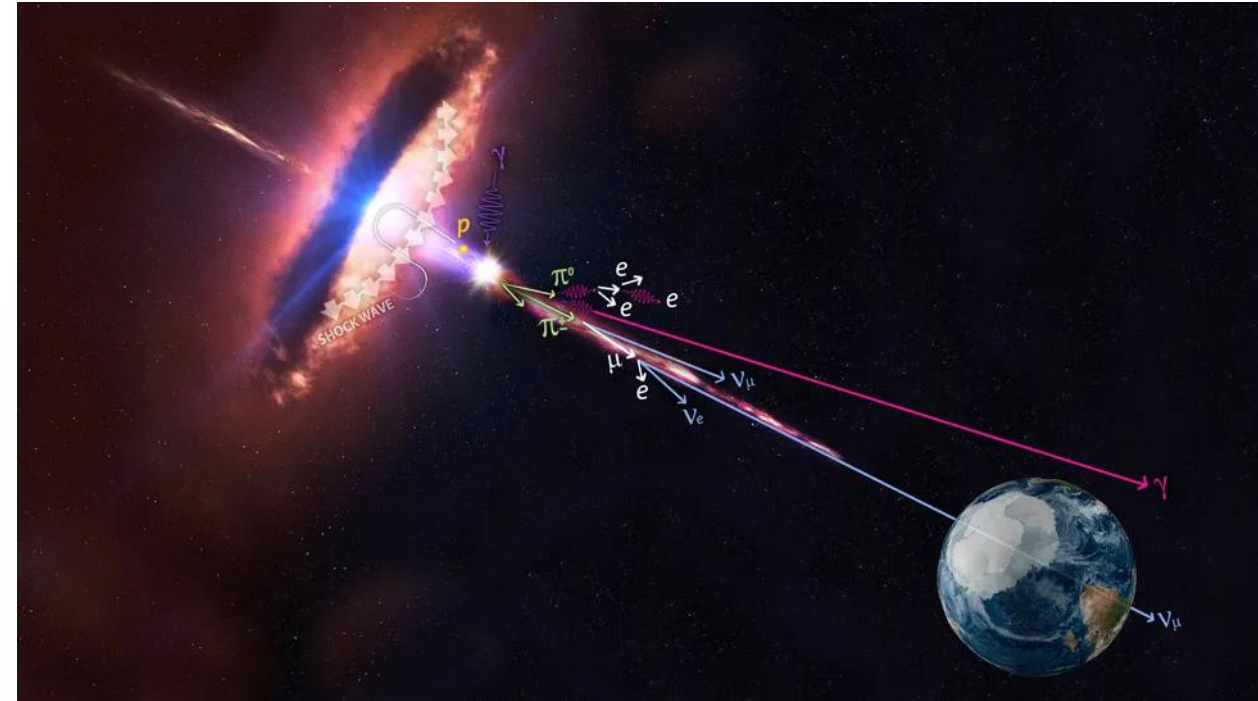


HIGH-ENERGY COSMIC MESSENGERS – CURRENT PROJECTS

COSMIC RAYS – THE PIERRE AUGER OBSERVATORY – ULB/VUB-IIHE

Multimessenger astrophysics

- Continuous search for **ultra-high energy neutrinos and photons** with limits excluding several theoretical models
- Constrains on the **lifetime and mass of super-heavy dark matter**
- Providing infrastructure and a **test environment for: FAST, IceCube-Gen2, GCOS, GRAND, SWGO, PEPS**



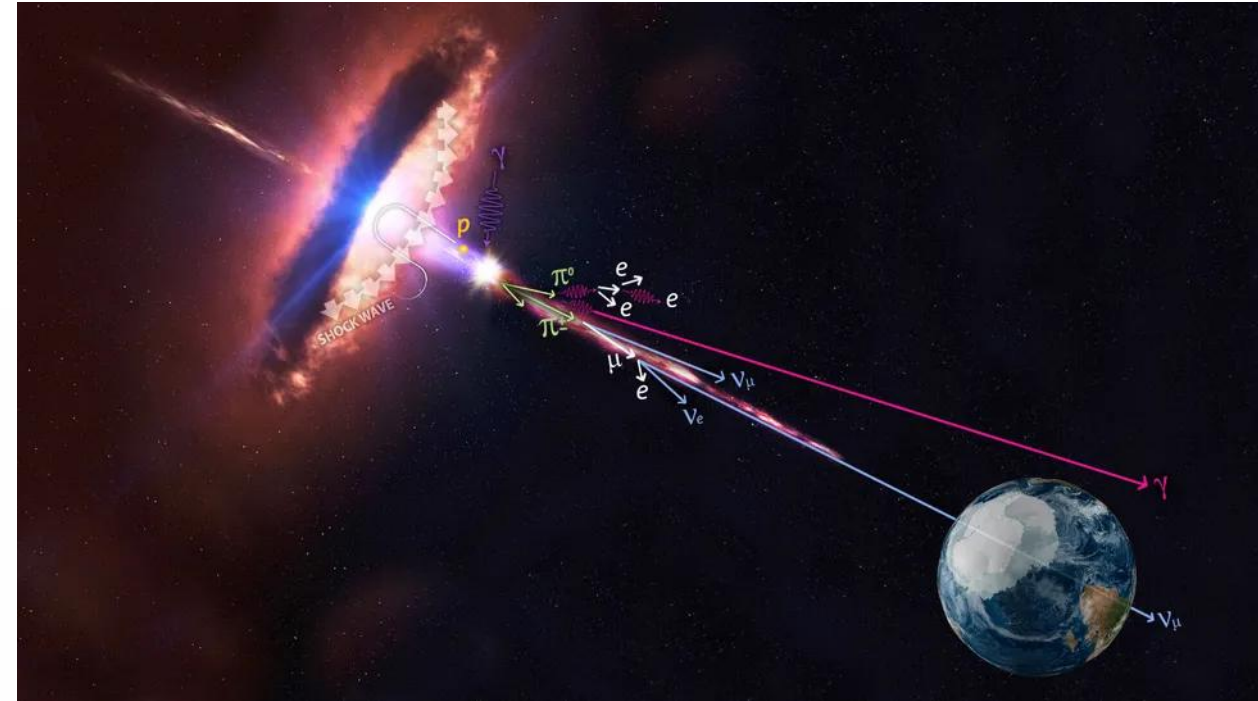
HIGH-ENERGY COSMIC MESSENGERS – CURRENT PROJECTS

COSMIC RAYS – THE PIERRE AUGER OBSERVATORY – ULB/VUB-IIHE

Belgian leadership/responsibilities:

Ioana Maris (ULB):

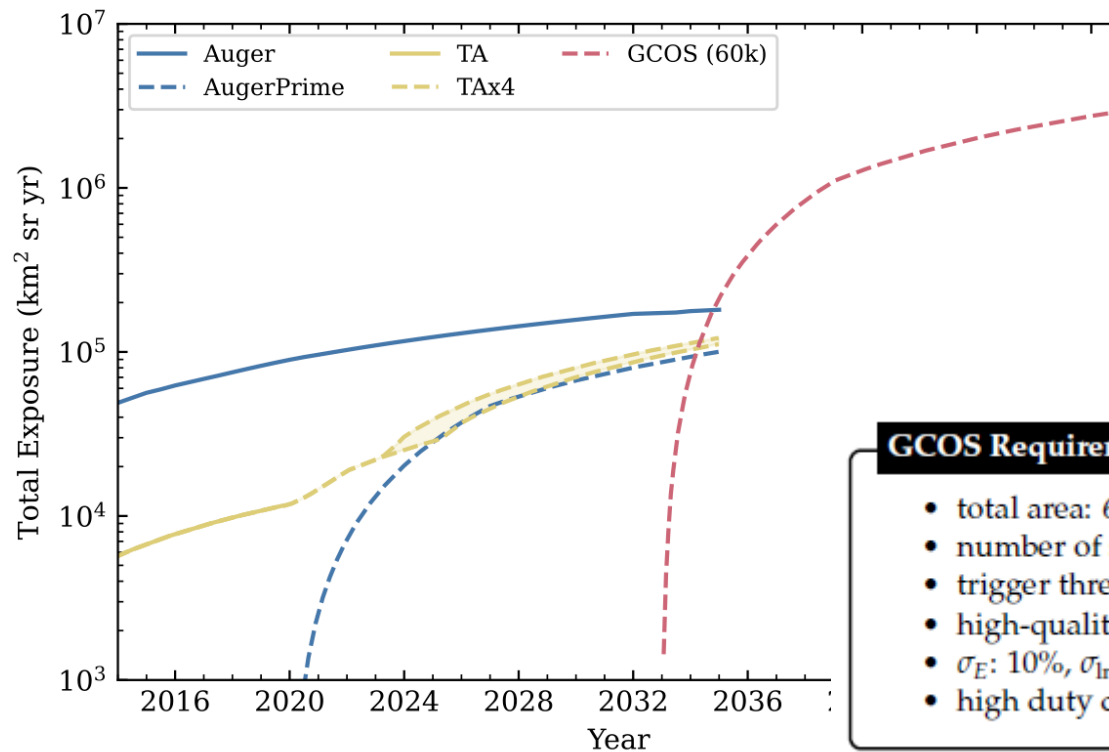
- Publication Committee Chair Pierre Auger collaboration.
- Task leader: Auger as test environment.



HIGH-ENERGY COSMIC MESSENGERS – FUTURE PROJECTS

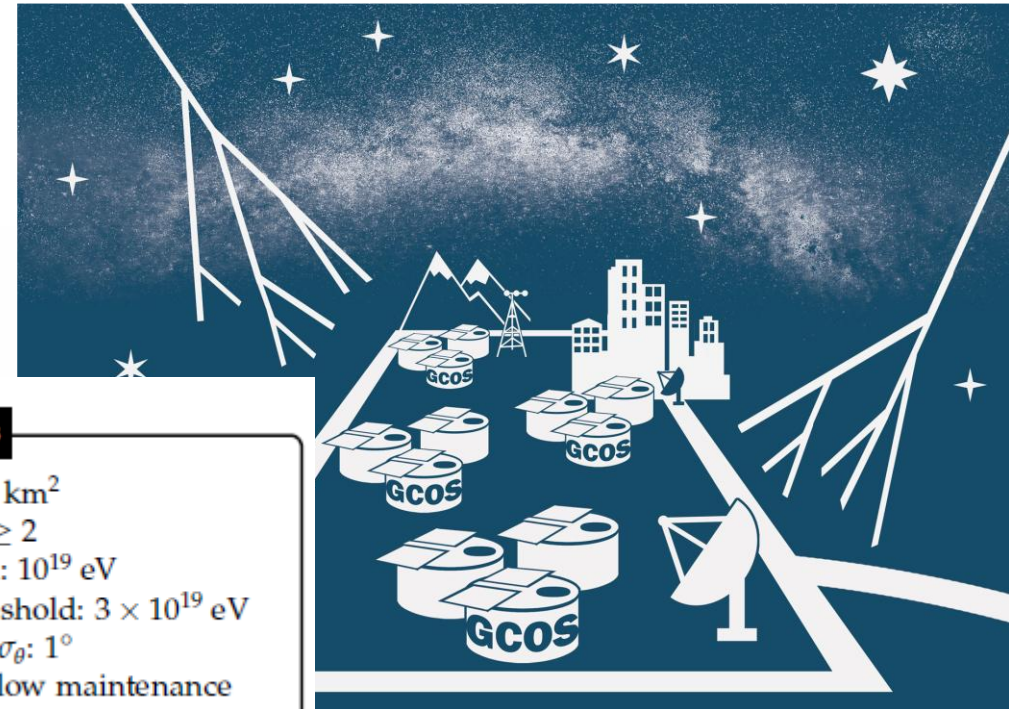
COSMIC RAYS – GCOS – ULB-IIHE

Figures from and more information at: <https://arxiv.org/abs/2502.05657>
(GCOS white paper)



GCOS Requirements

- total area: 60 000 km²
- number of sites ≥ 2
- trigger threshold: 10^{19} eV
- high-quality threshold: 3×10^{19} eV
- σ_E : 10%, $\sigma_{\ln A}$: 1, σ_θ : 1°
- high duty cycle, low maintenance



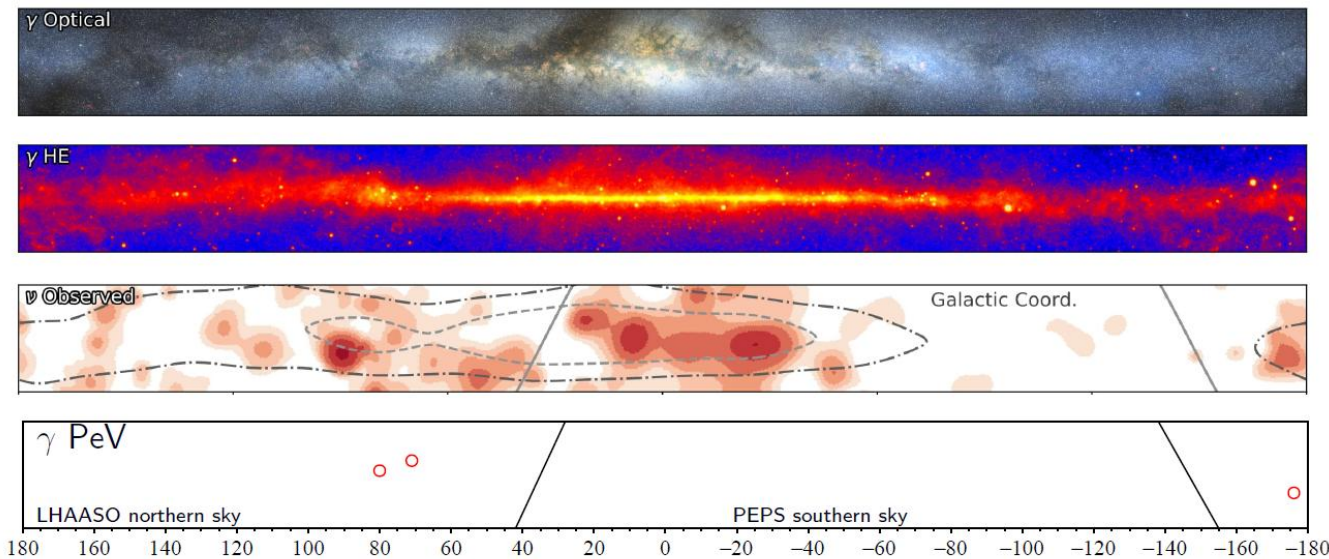
Belgian contributions and responsibilities:
Ioana Maris member of the GCOS core group.

HIGH-ENERGY COSMIC MESSENGERS – FUTURE PROJECTS

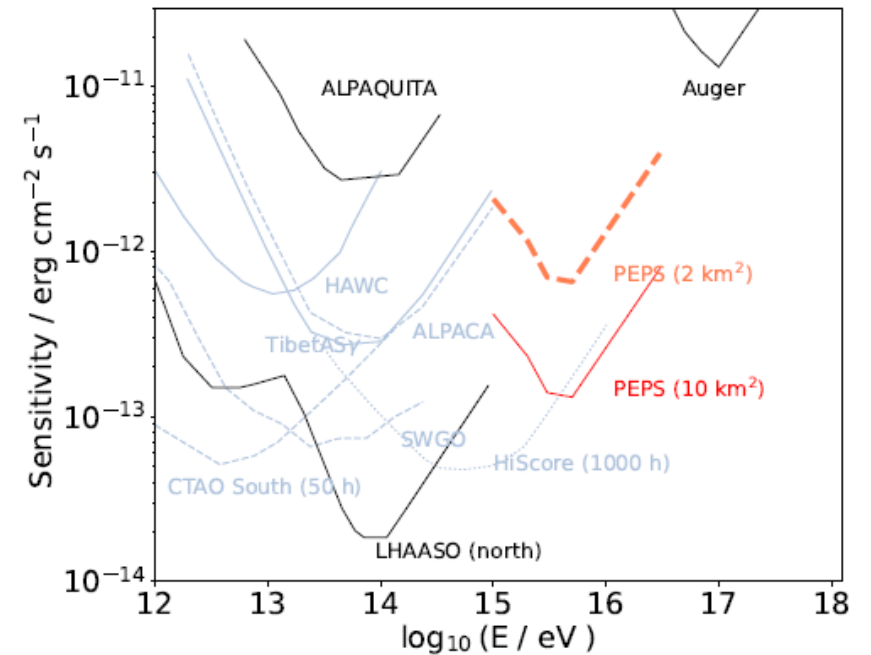
GAMMA-RAYS – PEPS – ULB-IIHE

Probing Extreme PeVatron Sources (PEPS) ; Project PI: Ioana Maris (ULB-IIHE)

10 km² detector to measure gamma-rays between 1 PeV and 50 PeV



Discovery potential in the new energy range with
a unique view of the Galactic Center and southern galactic sky!



HIGH-ENERGY COSMIC MESSENGERS – FUTURE PROJECTS

GAMMA-RAYS – PEPS – ULB-IIHE

Probing Extreme PeVatron Sources (PEPS)



Prototype stations:

- Components build and tested in Brussels.
- First 2 prototype layered Water Cherenkov Detectors installed at Auger, 3 under construction.

Envisioned array: 55 stations/km².

- Phase 1: 110 stations
- Phase 2: 550 stations

Timeline since funding – 2-2.5 yr. Completion phase 1 aimed at for 2027-2028.

Total hardware cost < 10 M€



(2024)

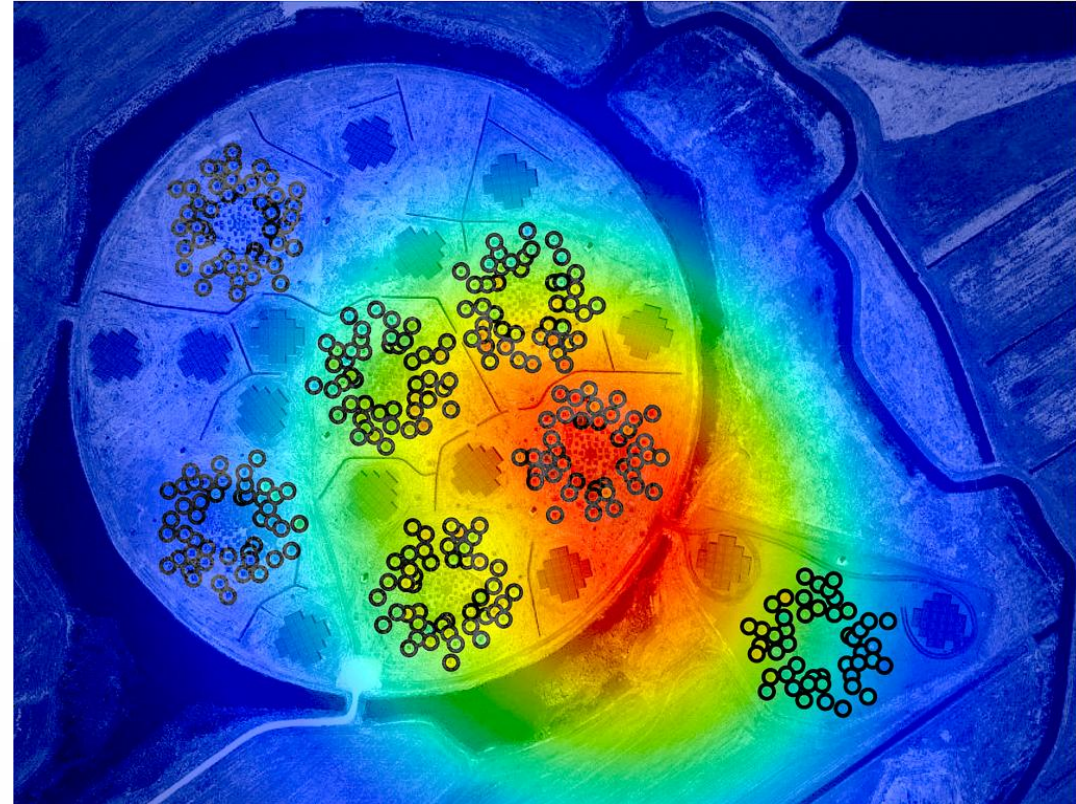


HIGH-ENERGY COSMIC MESSENGERS – CURRENT PROJECTS

COSMIC RAYS – PRECISION PHYSICS USING RADIO: LOFAR AND SKA – VUB-IIHE



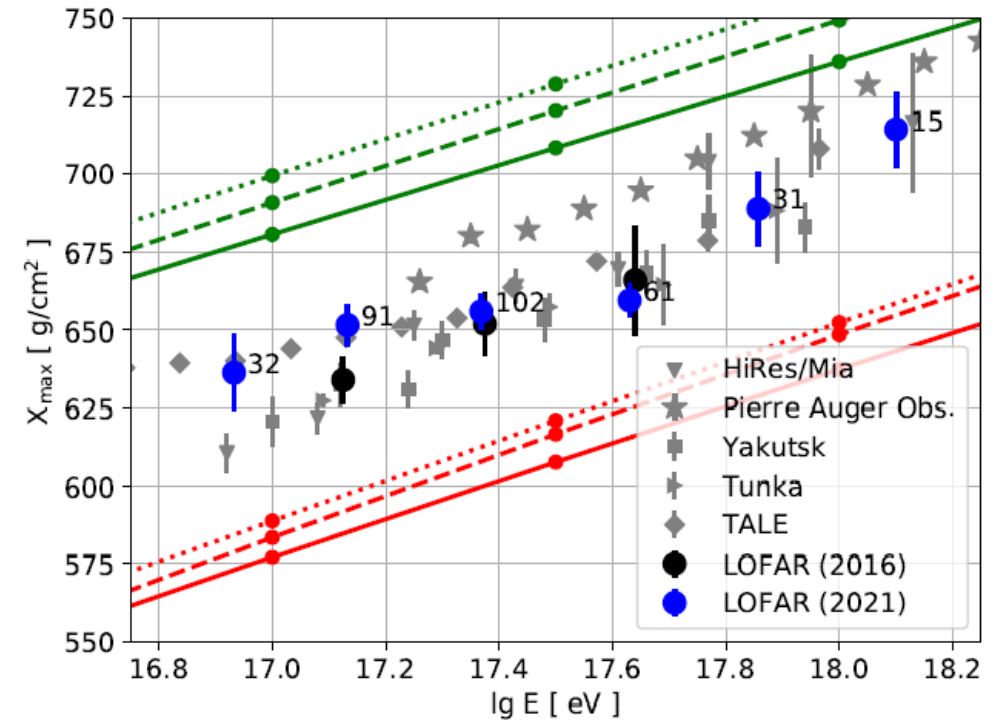
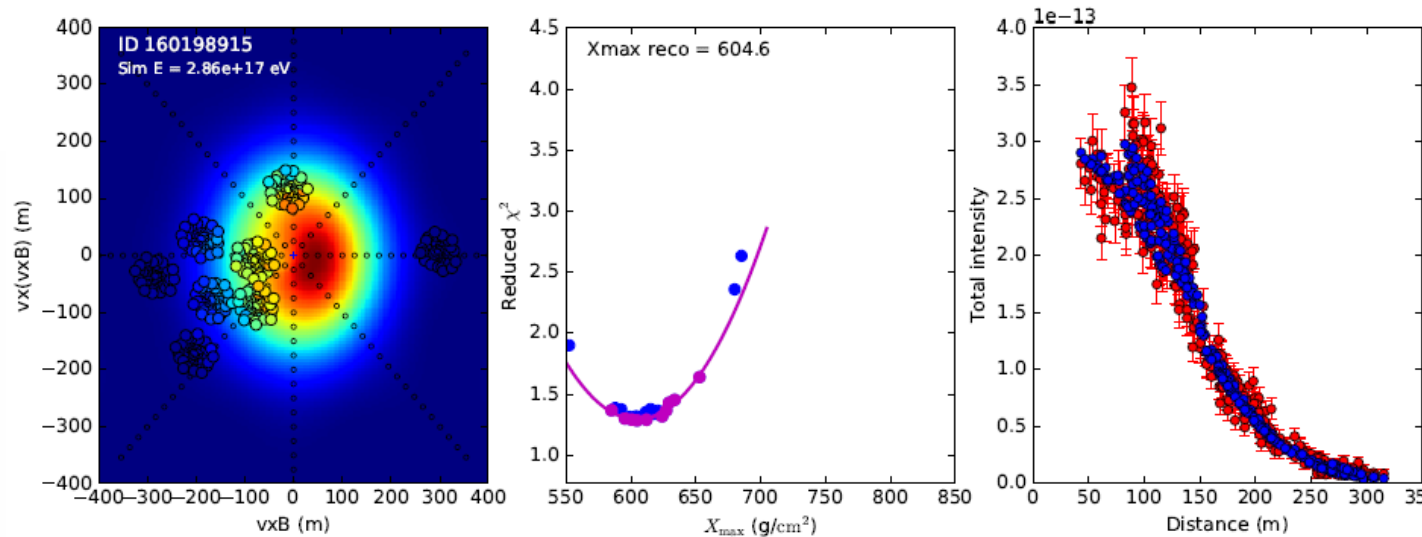
**Cosmic-ray trigger
from particle
detectors (LORA)**



LOFAR CR detection project PI: Stijn Buitink (VUB-IIHE)

HIGH-ENERGY COSMIC MESSENGERS – CURRENT PROJECTS

COSMIC RAYS – PRECISION PHYSICS USING RADIO: LOFAR AND SKA – VUB-IIHE



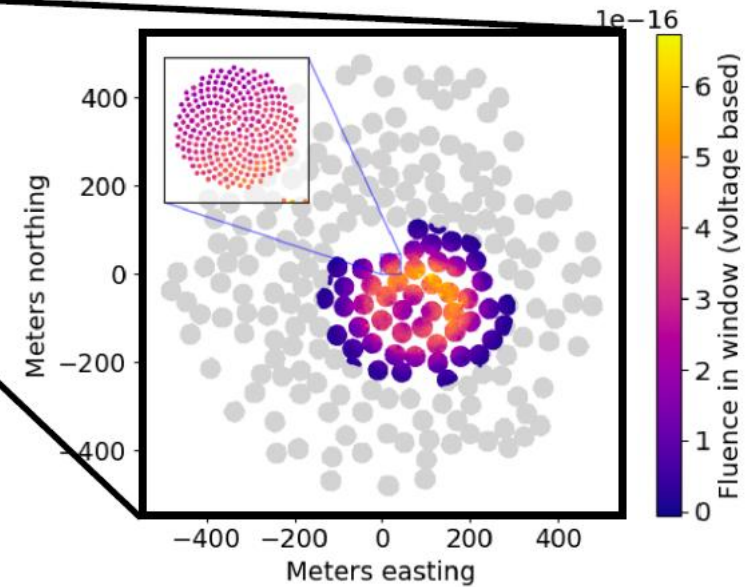
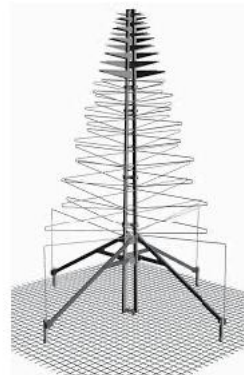
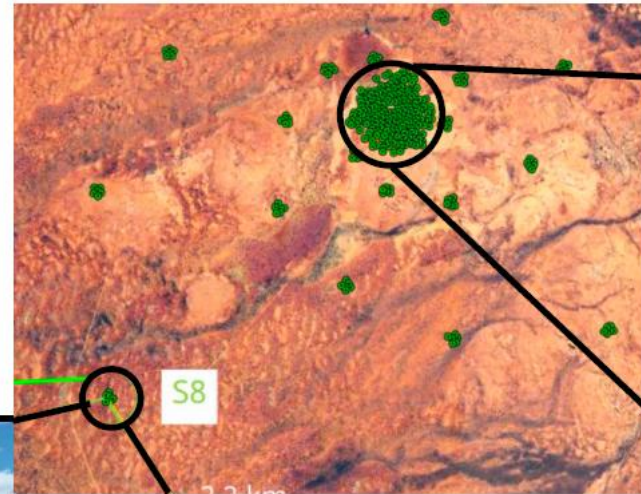
Phys. Rev. D 103, 102006 (2021)

HIGH-ENERGY COSMIC MESSENGERS – CURRENT PROJECTS

COSMIC RAYS – RADIO: SKA-LOW CORE CONSTRUCTION STARTED – VUB-IIHE



Western Australia
700 km north of Perth



core region: ~60.000 antennas within 1 km²
clustered in stations of 256 antennas
SKALA4 LPDA antennas 50-350 MHz

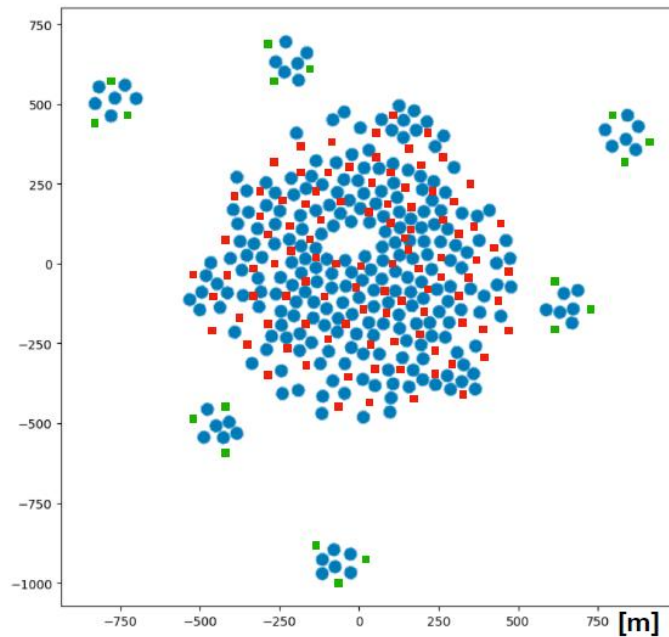
2

Stijn Buitink ICRC2025 Geneva

HIGH-ENERGY COSMIC MESSENGERS – CURRENT PROJECTS

COSMIC RAYS – RADIO: SKA-LOW CORE CONSTRUCTION STARTED – VUB-IIHE

The SKA Particle detector array



- Antenna station
- Particle detectors dense array (~100 units)
- Particle detectors satellite (~18 units, optional)

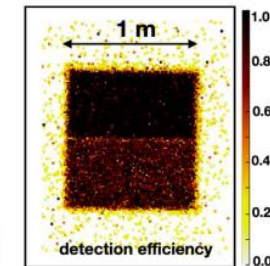
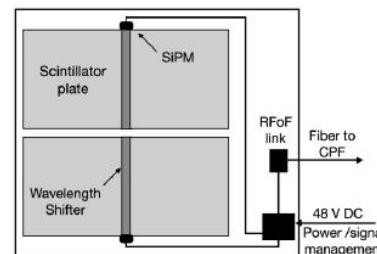


8-station prototype array installed at Murchison Widefield Array

Design: **Univ. of Manchester**
Deployment: **Curtin Univ.**
DAQ: **CSIRO**
J. Bray et al., NIMPA 973, id. 164168 (2020)

Construction of scintillator array now fully funded:

FWO Belgium, **NOVA** Netherlands + in-kind contributions **KIT** (Karlsruhe, Germany)
3rd generation design developed at **KIT** and **Radboud Nijmegen**



Same power cables & RFoF links as antennas
Very strict regulations on radio quietness
On-board SiPM voltage control to correct for temperature fluctuations

4

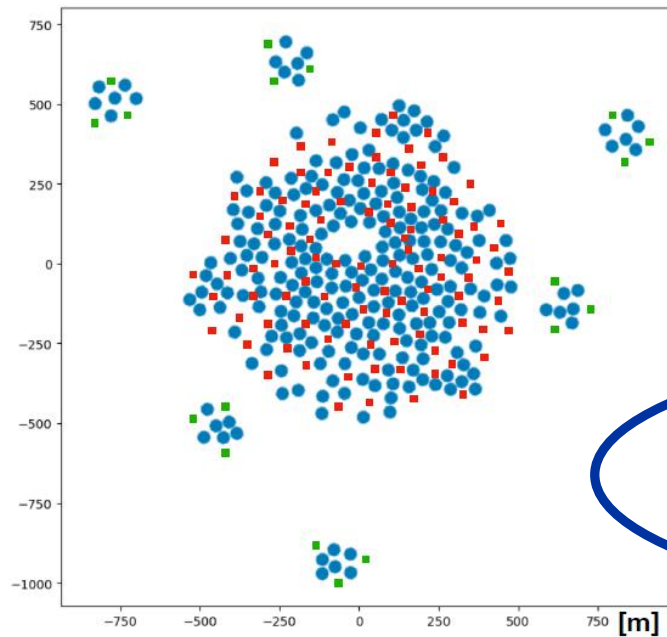
first light at KIT muon tower

Stijn Buitink ICRC2025 Geneva

HIGH-ENERGY COSMIC MESSENGERS – CURRENT PROJECTS

COSMIC RAYS – RADIO: SKA-LOW CORE CONSTRUCTION STARTED – VUB-IIHE

The SKA Particle detector array



- Antenna station
- Particle detectors dense array (~100 units)
- Particle detectors satellite (~18 units, optional)

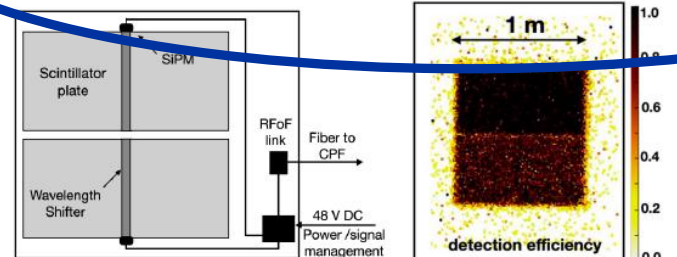


8-station prototype array installed at Murchison Widefield Array

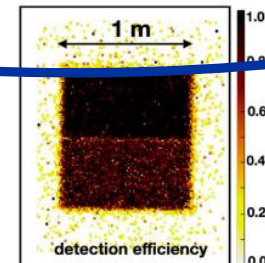
Design: **Univ. of Manchester**
Deployment: **Curtin Univ.**
DAQ: **CSIRO**
J. Bray et al., NIMPA 973, id. 164168 (2020)

Construction of scintillator array now fully funded:

FWO Belgium, **NOVA** Netherlands + in-kind contributions **KIT** (Karlsruhe, Germany)
3rd generation design developed at **KIT** and **Radboud Nijmegen**



first light at KIT muon tower

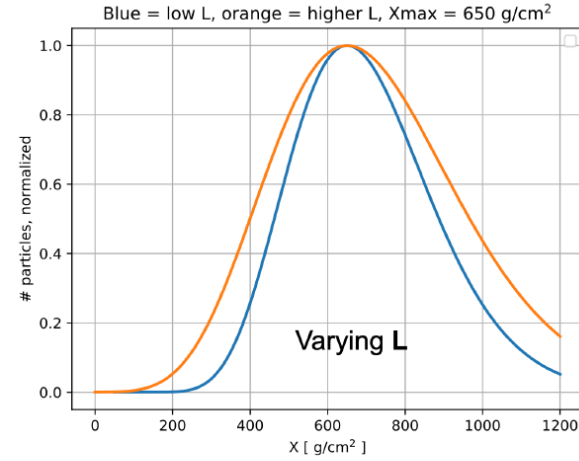
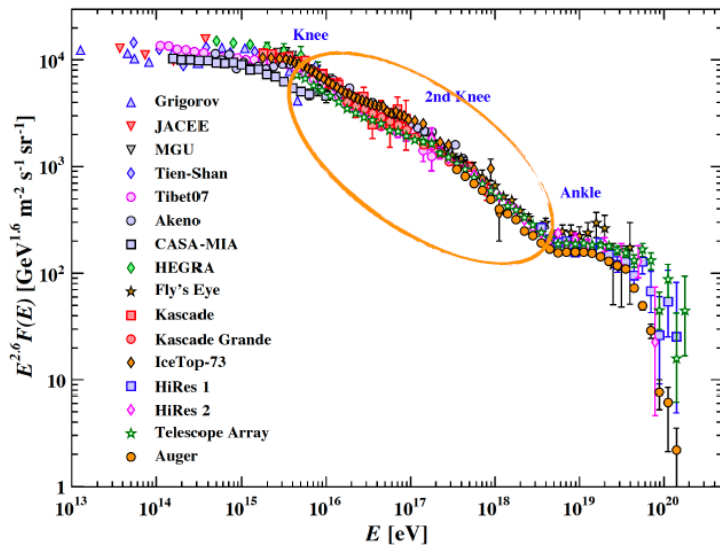


Same power cables & RFoF links as antennas
Very strict regulations on radio quietness
On-board SiPM voltage control to correct for temperature fluctuations

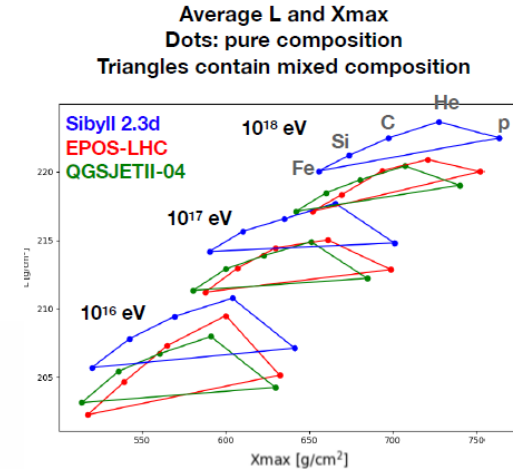
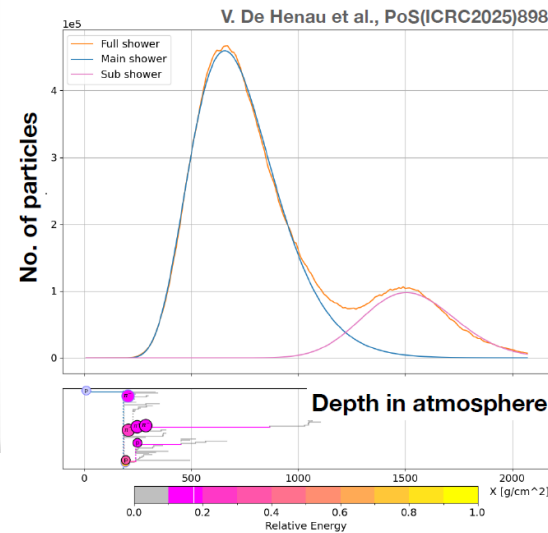
Stijn Buitink ICRC2025 Geneva

HIGH-ENERGY COSMIC MESSENGERS – CURRENT PROJECTS

COSMIC RAYS – RADIO: SKA-LOW CORE CONSTRUCTION STARTED – VUB-IIHE



A. Corstanje et al., PoS(ARENA2022)024



- Extremely detailed mass-composition studies in the galactic-extragalactic transition region.
- Precision tests of hadronic interaction models.

HIGH-ENERGY COSMIC MESSENGERS

NEUTRINOS – RNO-G (VUB/ULB-IIHE - UGENT) AND RET (VUB/ULB-IIHE)

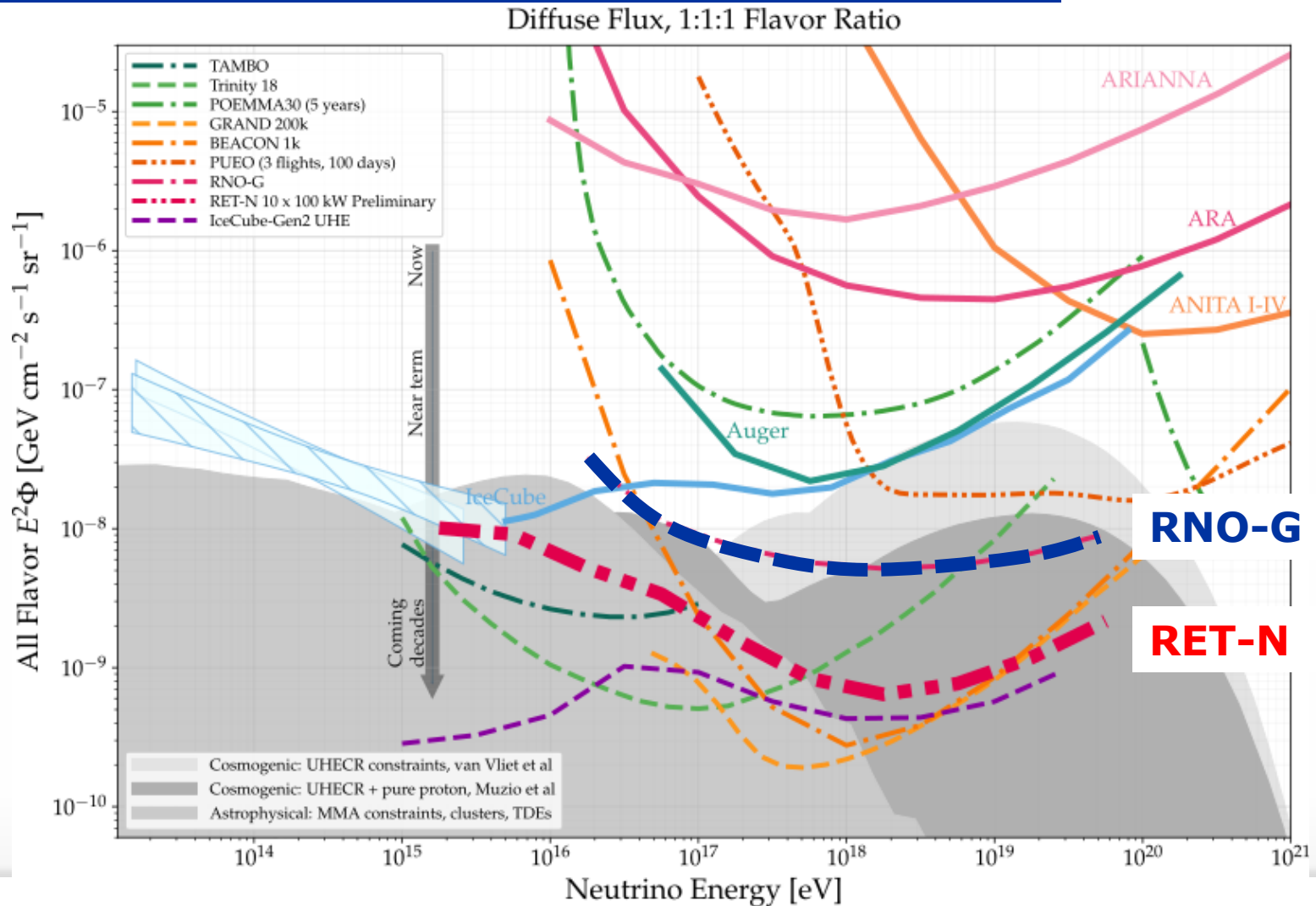


- **Currently under construction:**
8/35 stations taking data
- **Completion in coming 2-3 years**

RADAR ECHO TELESCOPE



- **RET-CR pathfinder operated in 2023-2024**
- **RET-N first station installation foreseen ~2028**
- **Completion 2030-2035**



Probing the high-energy universe-- KD de Vries (VUB)

Sept, 12, 2025 | 21

HIGH-ENERGY COSMIC MESSENGERS

NEUTRINOS – RNO-G AND RET VUB/ULB-IIHE - UGENT



Belgian groups RNO-G:

VUB-IIHE:

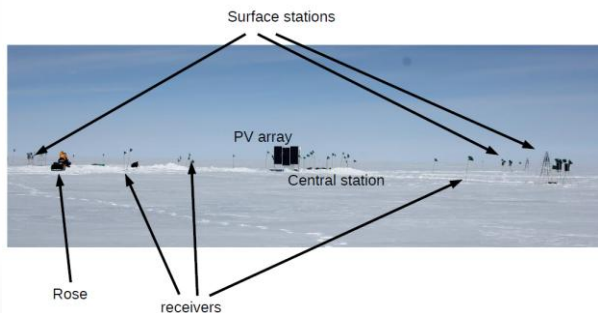
N. van Eijndhoven – project PI

K.D. de Vries

ULB-IIHE: S. Toscano, J.A. Aguilar

Ugent: J. Stachurska

RADAR ECHO TELESCOPE



Belgian groups RET:

VUB-IIHE:

K.D. de Vries – project PI

N. Van Eijndhoven

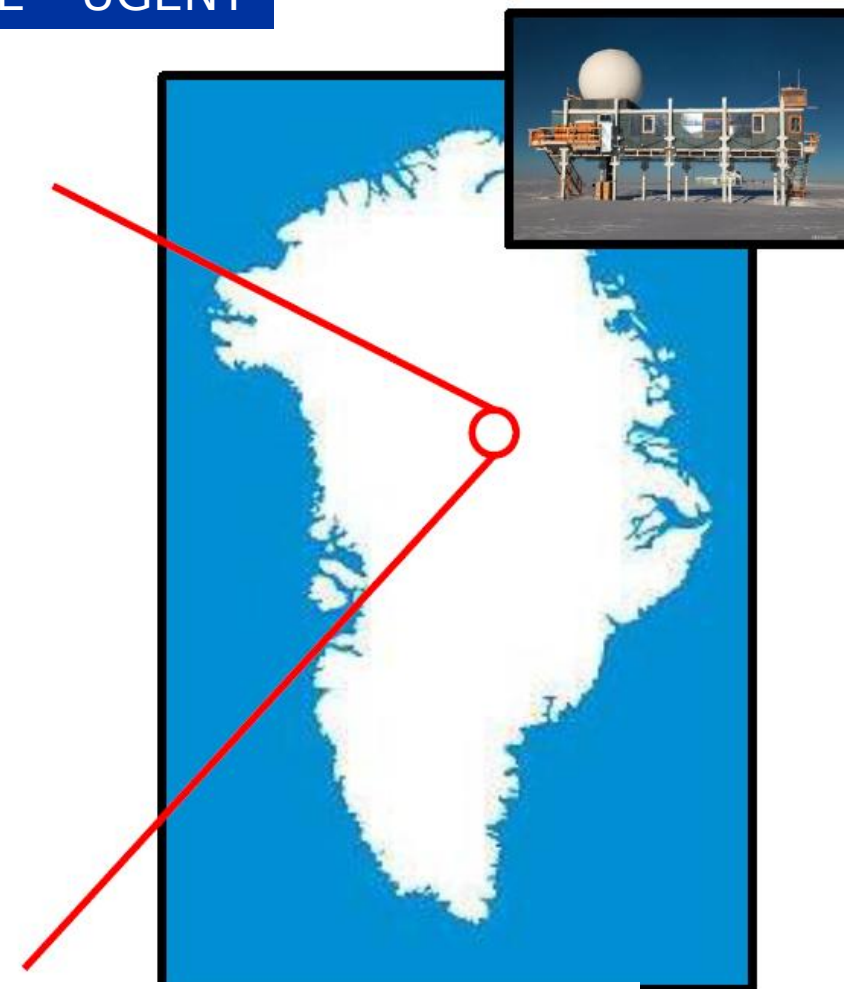
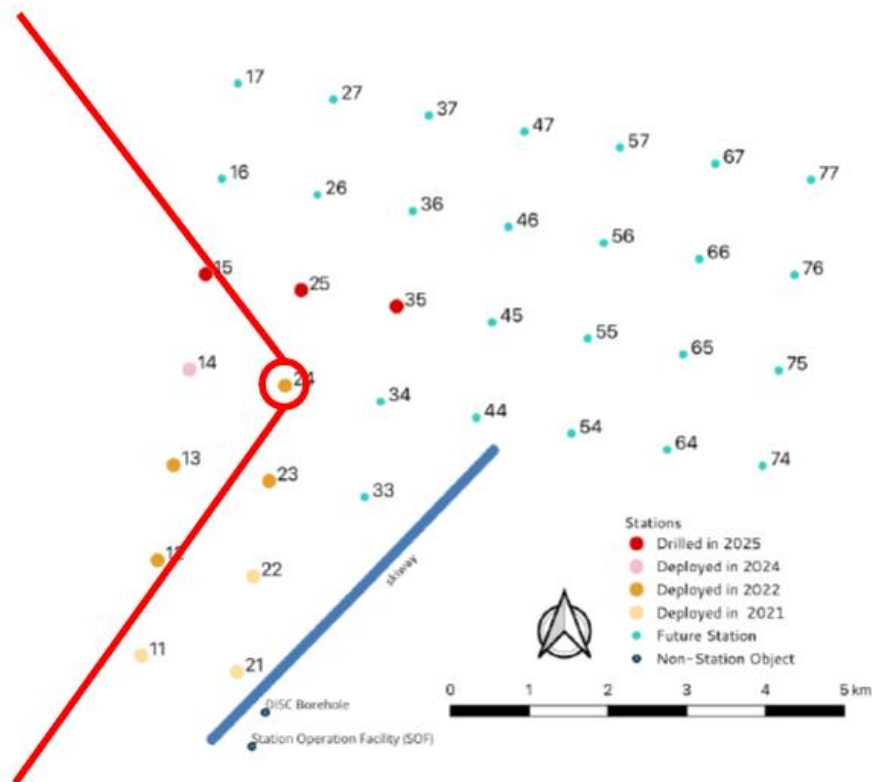
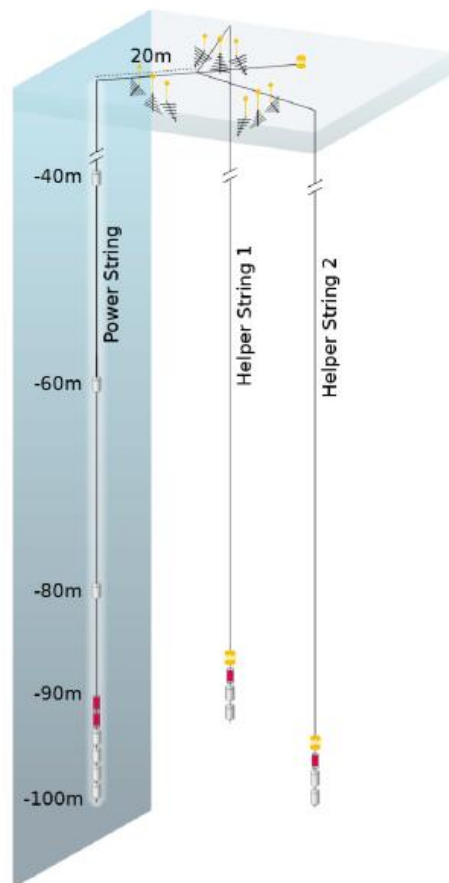
ULB-IIHE: S. Toscano



<https://geo-summit.org/summit-station>

HIGH-ENERGY COSMIC MESSENGERS

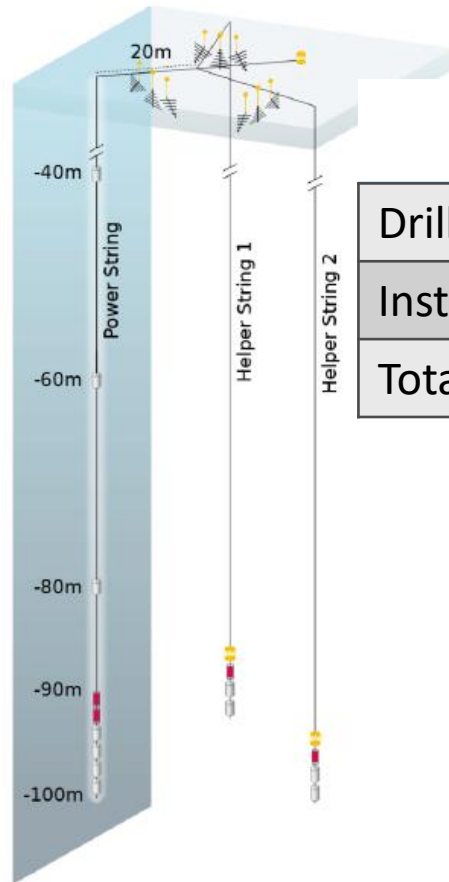
NEUTRINOS – RNO-G VUB/ULB-IIHE - UGENT



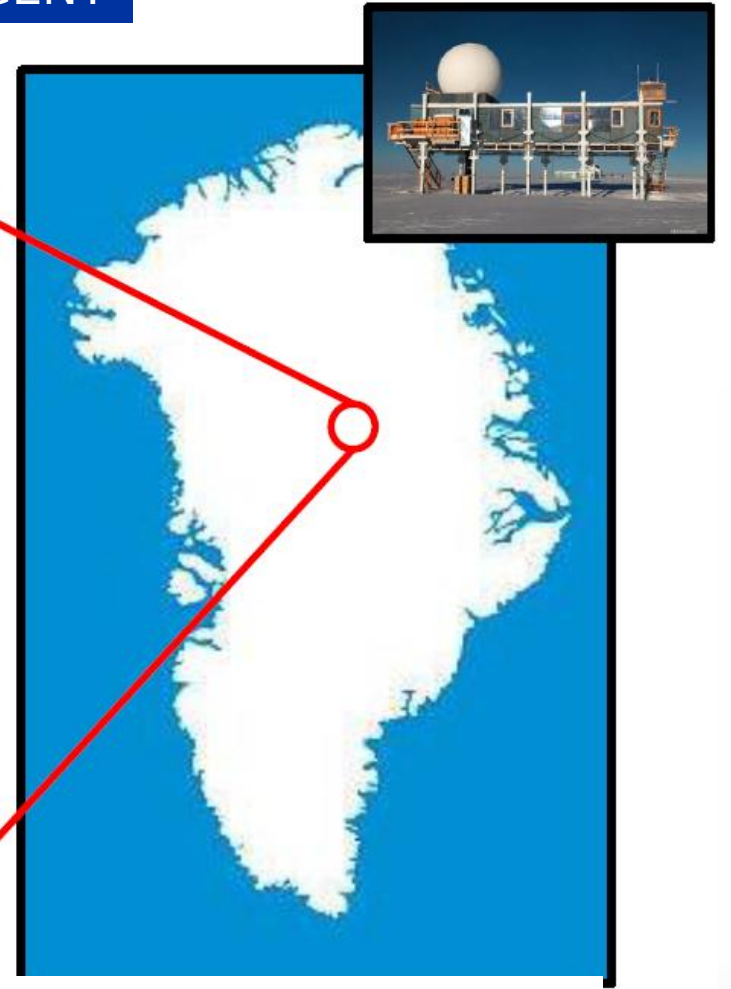
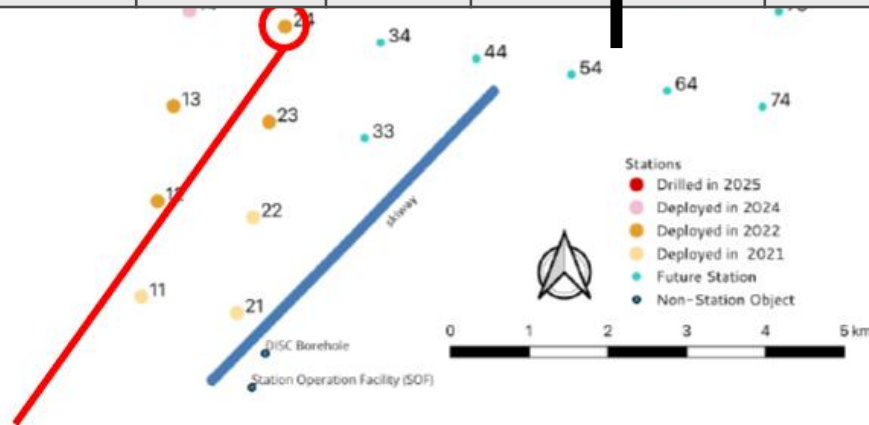
J. Stoffels (VUB-IIHE), TAUP 2025

HIGH-ENERGY COSMIC MESSENGERS

NEUTRINOS – RNO-G VUB/ULB-IIHE - UGENT



	≤ 2023	2024	2025	2026	2027	
Drilled	7	1	4	>4	>6	...
Installed	7	1	0	>8	>6	...
Total	7	8	8	>16	>22	...



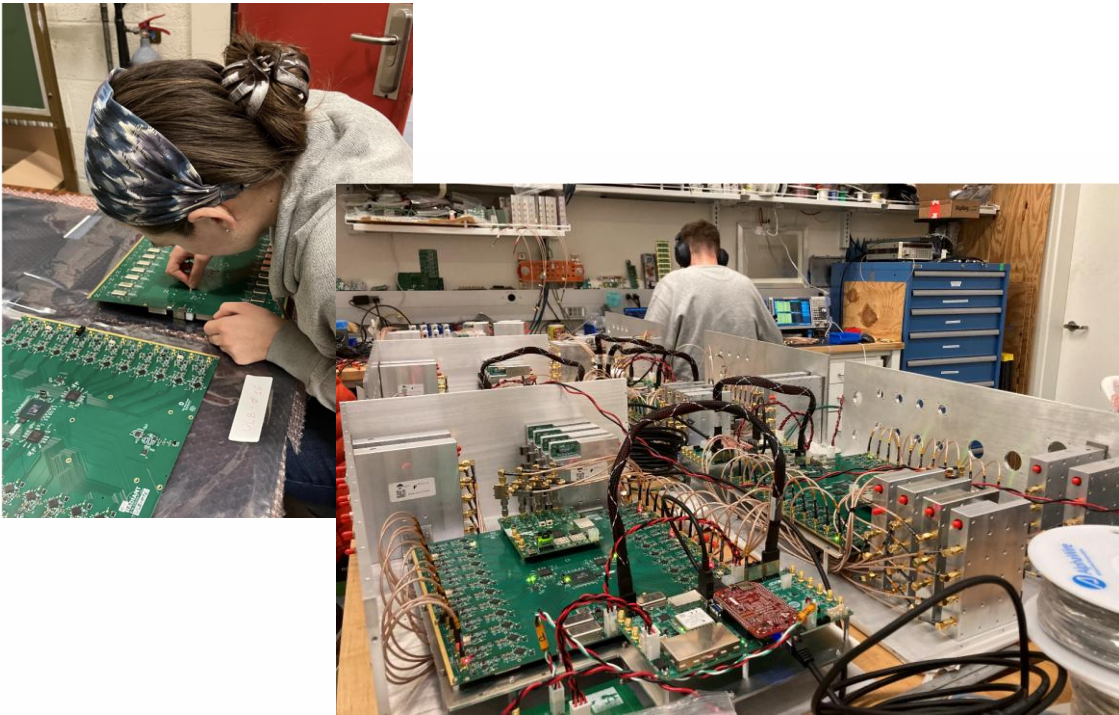
J. Stoffels (VUB-IIHE), TAUP 2025

5

HIGH-ENERGY COSMIC MESSENGERS

NEUTRINOS – RNO-G – BELGIAN INVOLVEMENT

Hardware efforts at the IIHE (ULB-VUB): Radiant board testing and development

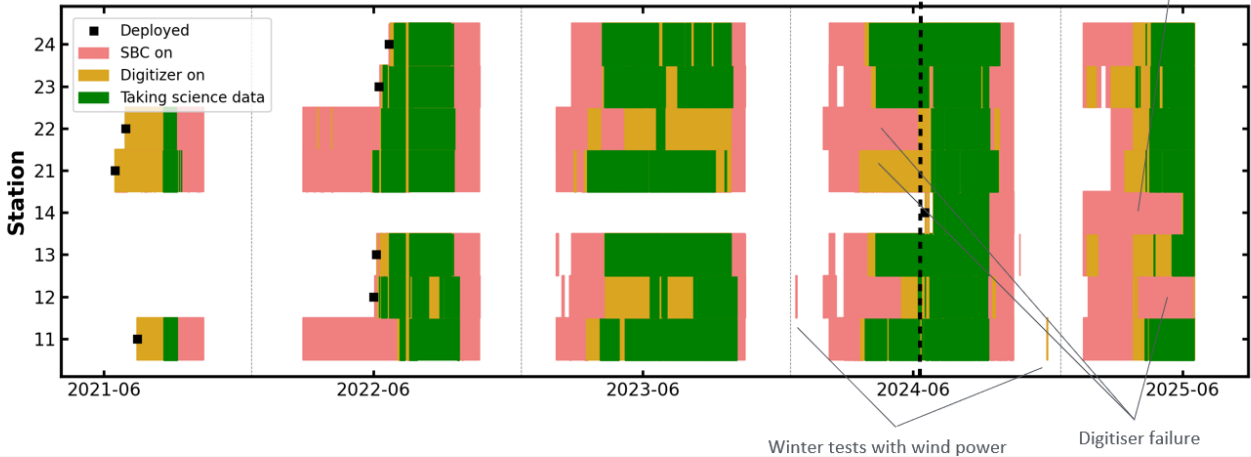


Commisioning and deployment (ULB-VUB):



Major maintenance (replaced electronics in 6/7 stations)

Broken comms. connection



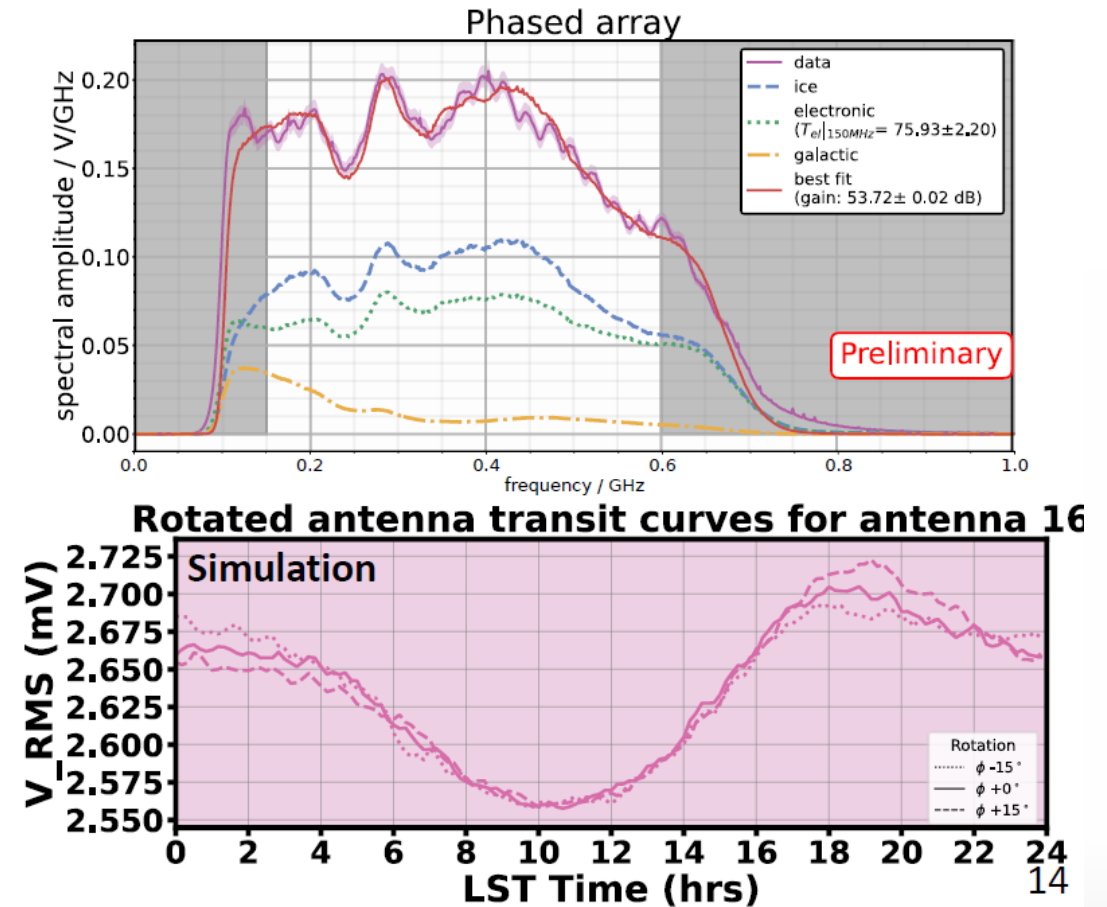
HIGH-ENERGY COSMIC MESSENGERS

NEUTRINOS – RNO-G – BELGIAN INVOLVEMENT

Detector calibration (VUB/ULB-IIHE)

- Absolute calibration of signal amplitude using thermal noise
- Investigating surface antenna orientation using Galactic noise

Thermal and Galactic noise



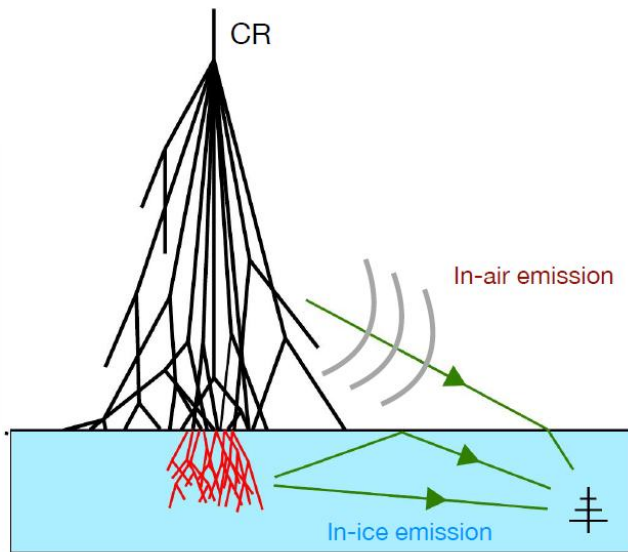
HIGH-ENERGY COSMIC MESSENGERS

NEUTRINOS – RNO-G – BELGIAN INVOLVEMENT

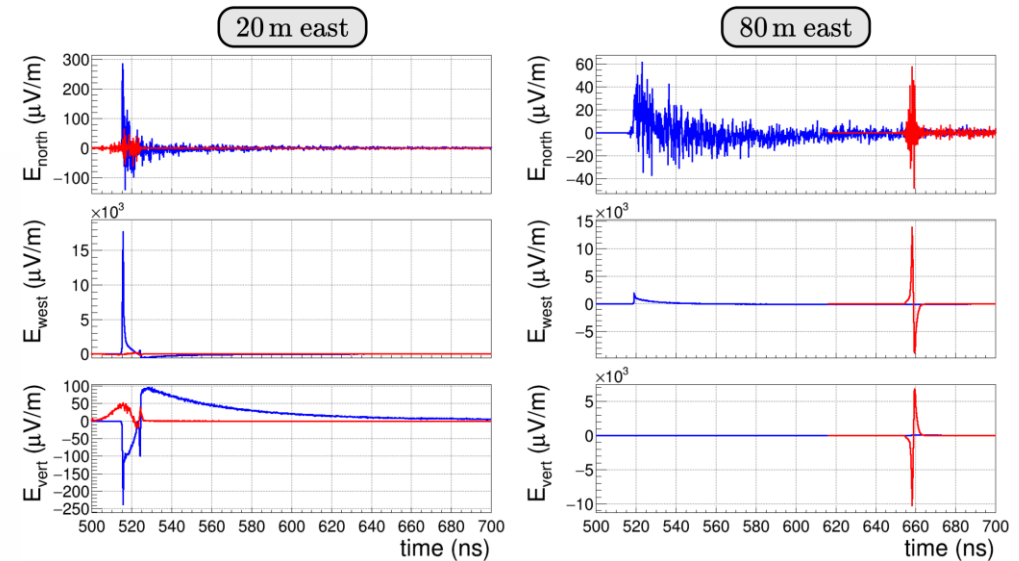
Cosmic-Ray analysis and simulations: Focus on Cosmic-ray in-ice signals

FAERIE simulation framework developed at VUB-IIHE

RNO-G In-ice Cosmic-Ray analysis working group lead by S. Toscano (ULB-IIHE)



credits:
Simon de Kockere



Data analysis ongoing: first signals expected soon

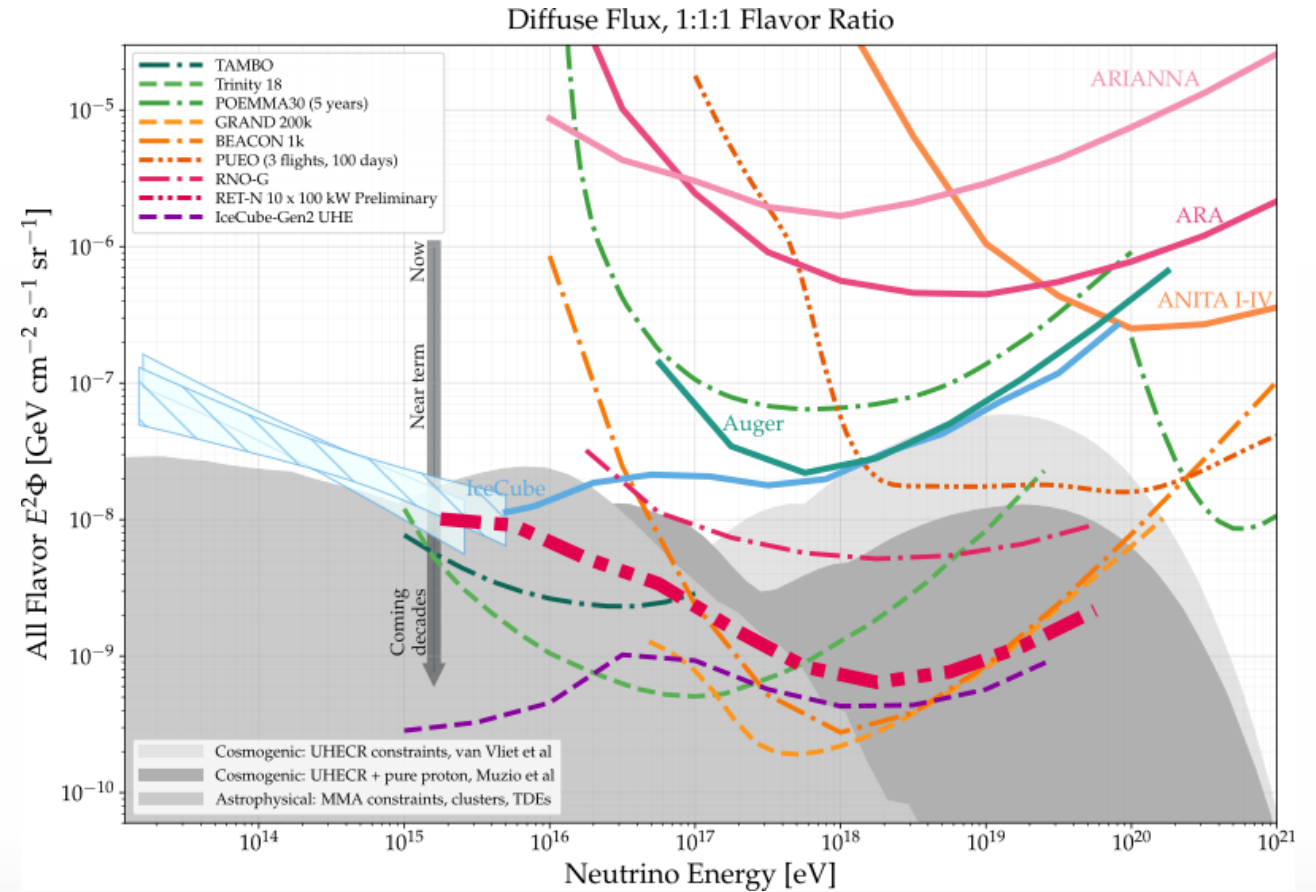
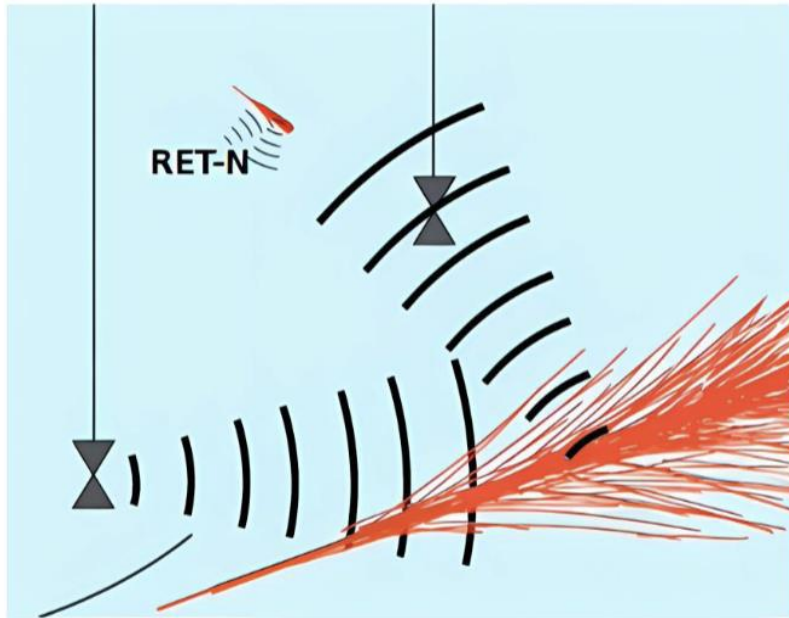


Proof of concept of the detector and method!!

HIGH-ENERGY COSMIC MESSENGERS – PATHFINDERS

NEUTRINOS – THE RADAR ECHO TELESCOPE (RET) – VUB/ULB-IIHE

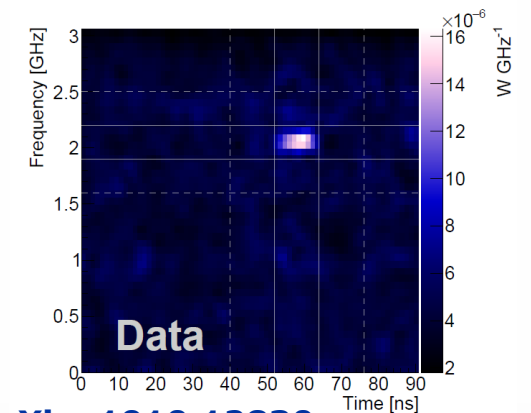
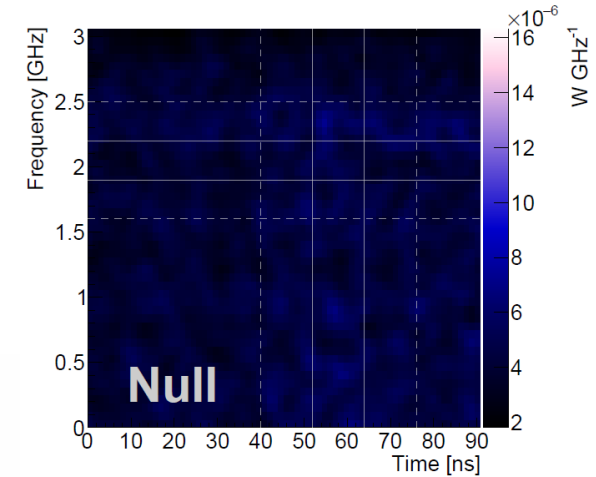
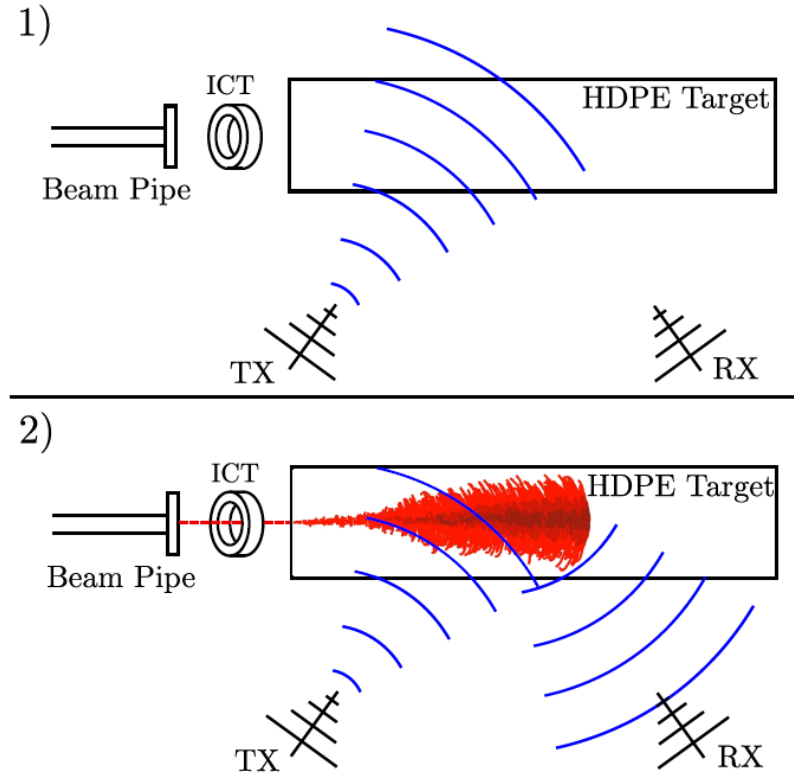
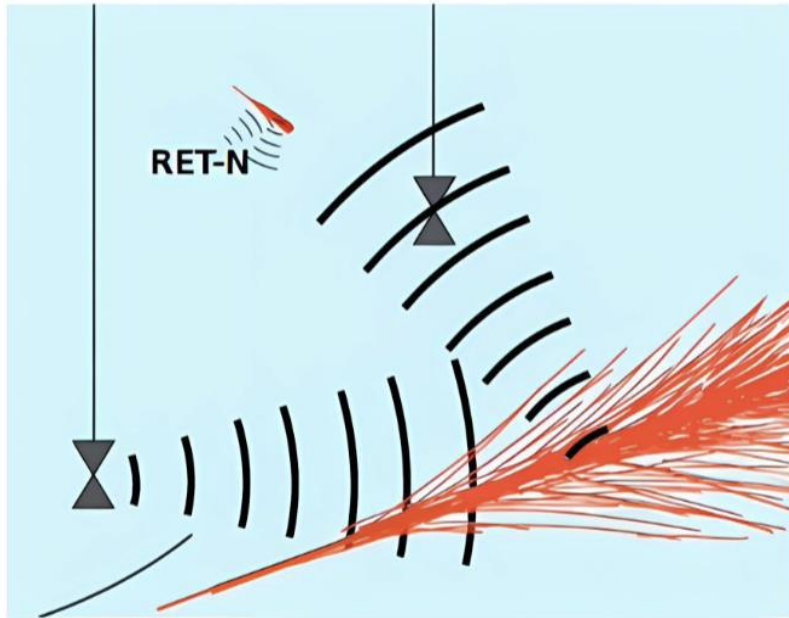
RADAR ECHO TELESCOPE



HIGH-ENERGY COSMIC MESSENGERS – PATHFINDERS

NEUTRINOS – THE RADAR ECHO TELESCOPE (RET) – VUB/ULB-IIHE

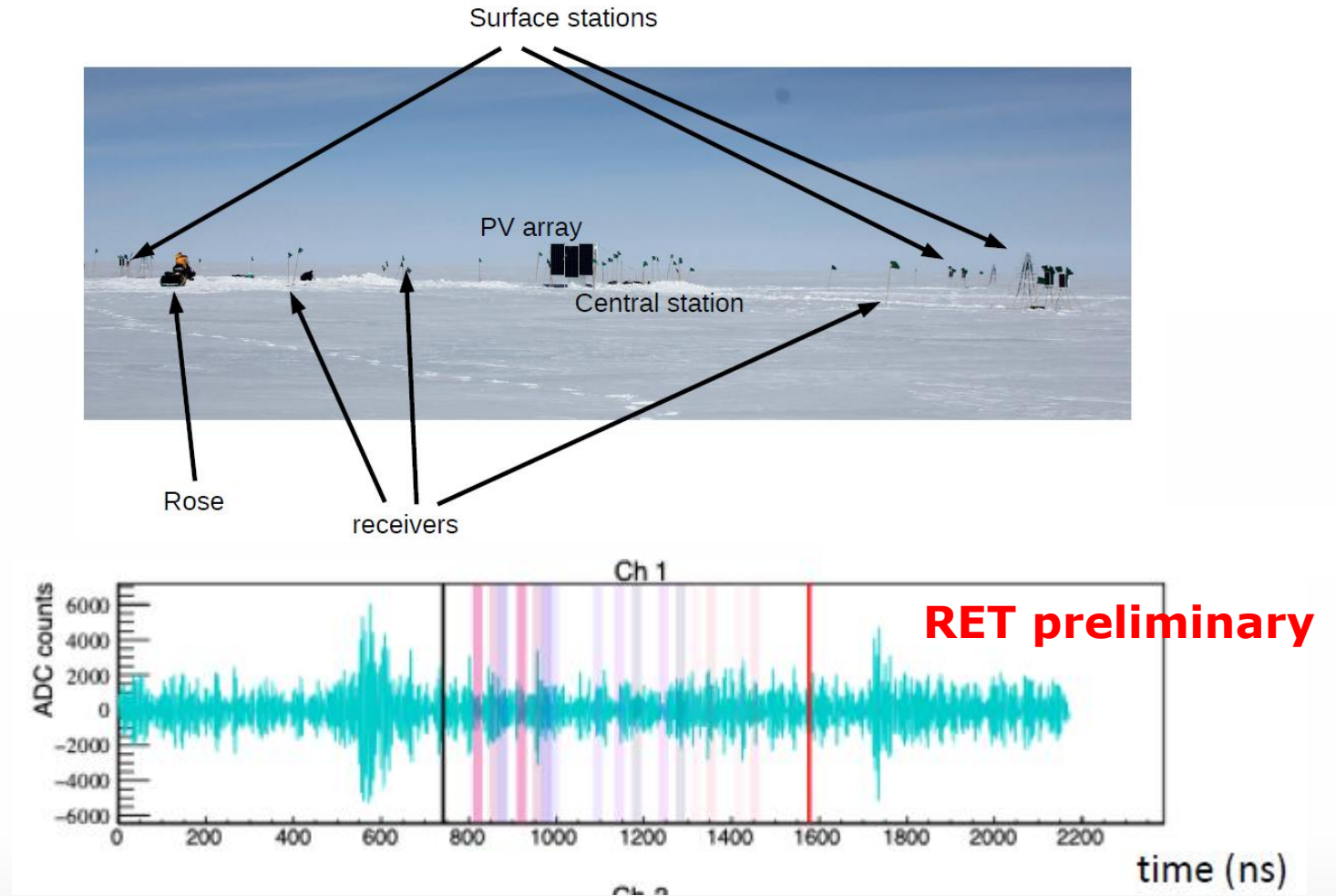
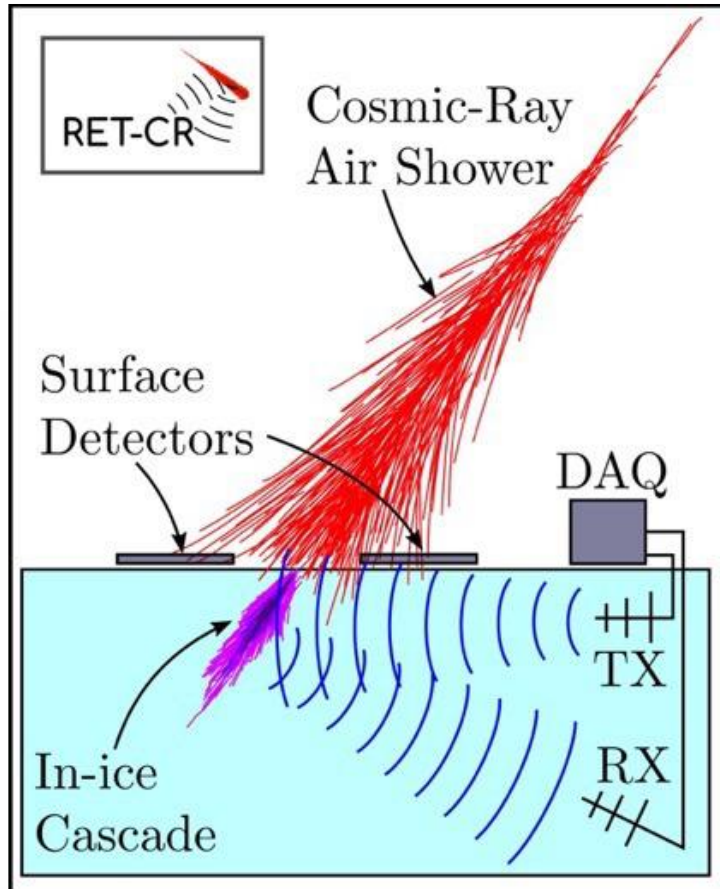
RADAR ECHO TELESCOPE



Phys. Rev. Lett. 124, 091101 (2020) ; arXiv: 1910.12830
Details on the method: Phys. Rev. D 100, 072003 (2019) ; arXiv:1810.09914

HIGH-ENERGY COSMIC MESSENGERS – PATHFINDERS

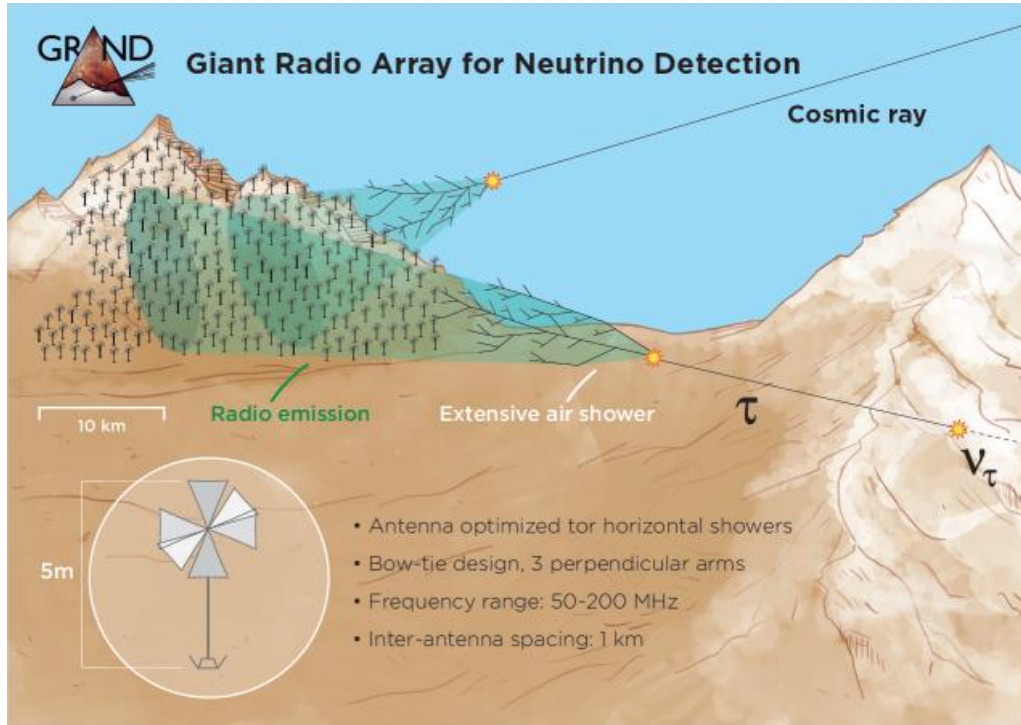
NEUTRINOS – THE RADAR ECHO TELESCOPE (RET) – VUB/ULB-IIHE



HIGH-ENERGY COSMIC MESSENGERS – PATHFINDERS

NEUTRINOS – GRAND – VUB-IIHE

Project PI: K. Kotera (10% VUB-IIHE)



✱ A staged approach with self-standing pathfinders

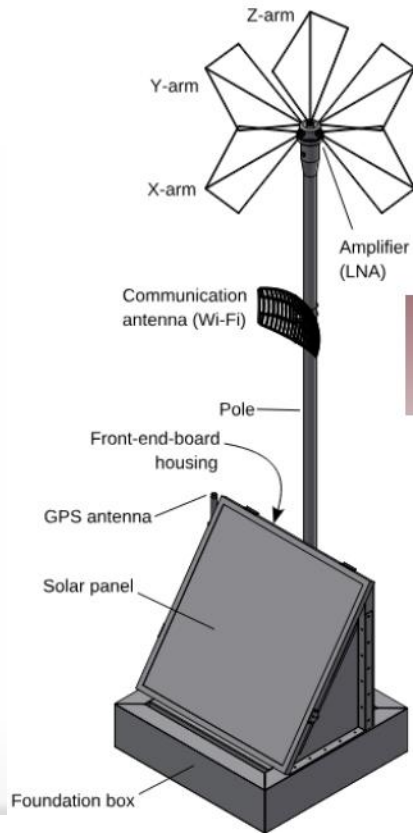
	GRANDProtos	GRAND10k	GRAND200k
	2023	2028	203x
Goals	Pure and efficient autonomous radio detection of very inclined air-showers Cosmic rays $10^{16.5-18}$ eV • Galactic/extragalactic transition • Muon puzzle • Radio transients	1st GRAND sub-array • discovery of EeV neutrinos for optimistic fluxes • radio transients (FRBs!)	sensitive all-sky detector Neutrino astronomy!
Setup	• GRANDProto300: 300 antennas over 200 km² • GRAND@Auger: 10 antennas for cross-calibration • GRAND@Nançay: 4 antennas for trigger testing	• 2 detectors of 5-10k antennas each in each hemisphere: GRAND-North (China) and GRAND-South (Argentina?) *Hybrid design possible (phased/standalone)	• 200,000 antennas over 200,000 km² • 20 sub-arrays of 10k antennas • on different continents
Budget	2 M€ 100 antennas produced Funded by China + ANR-DFG NUTRIG (France- Germany) + Radboud University	13 M€ 1500€/unit	300M€ in total 500€/unit to be divided between participating countries

HIGH-ENERGY COSMIC MESSENGERS – PATHFINDERS

NEUTRINOS – GRAND



GRANDProto300 & other prototypes: experimental setup



Deployment of 13 detection units (DUs)
in Gansu (China), 65 in 2025 and 289 later

Deployed Feb 2023



Deployment of 10 DUs on the Auger site in
Malargüe, Argentina (cross-calibration)

Deployed Aug 2023



Deployment of 4 DUs in Nançay radio
observatory (France) for trigger test (LPNHE)

Deployed Oct 2022



50-200MHz
analog filtering,
Electronics:
500MSPS sampling
FPGA+CPU
Bullet WiFi data
transfert

HIGH-ENERGY COSMIC MESSENGERS

BELGIAN LEADERSHIP AND RESPONSIBILITIES

Pierre Auger observatory:

- Publication Committee Chair Pierre Auger collaboration. (Ioana Maris, ULB-IIHE)
- Task leader: Auger as test environment. (Ioana Maris, ULB-IIHE)

GCOS:

- Core group member (Ioana Maris, ULB-IIHE)

PEPS:

- Project PI (Ioana Maris, ULB-IIHE)

LOFAR/SKA Cosmic-Ray detection:

- Project PI (S. Buitink, VUB-IIHE)

RNO-G:

- Project PI (N. van Eijndhoven, VUB-IIHE)
- Deep CR search working group lead (S. Toscano, ULB-IIHE)

RET:

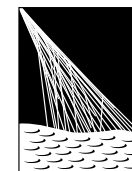
- Project PI (K.D. de Vries, VUB-IIHE)

GRAND:

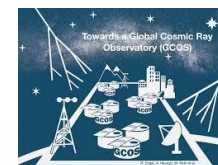
- Project PI (K. Kotera, 10% VUB-IIHE)

APPEC:

- FWO/FNRS representative (K.D. de Vries, VUB-IIHE / G. Bruno, UCLouvain)
- SAC member (I. Maris, VUB-IIHE)



PIERRE
AUGER
OBSERVATORY



LOFAR

SKAO



RNO-G
Radio Neutrino Observatory - Greenland

RADAR ECHO TELESCOPE

