# Master Thesis on JUNO

Status Report 01

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## **Outline**

### Introduction

My background Purpose of the Master Thesis

# Atmospheric muon background

Previous work

## **Atmospherical Neutrinos**

Simulations
Type of interactions
Charge Spectrum
First Hit Time

## **Next steps**

# Introduction: My background

- ▶ Bachelor and Master in Physics at ULB
- Mostly particle and mathematical physics courses
- Research internships at IIHE about JUNO (2023 and 2025)





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## **Introduction:** Purpose of the Master Thesis

**Title:** Atmospheric neutrino signal - Analysis of the first data from the JUNO experiment (provisional)

### 1. First Part

- Overview of the JUNO experiment
- Simulation data analysis for atmospheric neutrinos

### 2. Second Part

- JUNO first data analysis in its final configuration (scintillator)
- ▶ Identifying and characterizing atmospheric neutrino events
- Selection cuts and reconstruction algorithms optimized and validated with simulations

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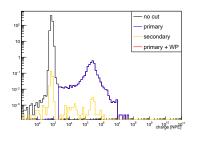
# Atmospheric muon background : Previous work

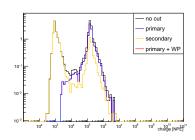
#### For water in the detector:

- 1. Offline Data Simulations
  - Monte Carlo simulation for cosmic ray muons
- 2. Real Data with water
  - Data taken in early 2025
  - ▶ TQ level (COTI reconstruction at software level)

# Atmospheric muon background : Previous work

### Atmospheric muon background :





- ► Atmospheric muon **charge spectrum** for water data in JUNO
- Comparison between MC simulations and TQ data

# **Atmospherical Neutrinos: Simulations**

#### Data set

- ▶ GENIE neutrino MC simulations for atmospheric neutrinos (TQ level)
- Energy distribution up to 20 GeV
- $ightharpoonup 5 imes 10^5$  events simulated  $(
  u_e + 
  u_\mu)$
- Propagation of secondary particles in liquid scintillator with GEANT4

#### Event selection

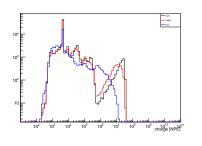
 Based on atmospheric muon background study (parameters to be optimized)

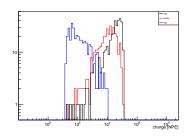
# **Atmospherical Neutrinos: Type of interactions**

## Different types of neutrino detected:

- ▶ Charged Current (CC) :  $\nu_e$  and  $\nu_\mu$ 
  - W<sup>±</sup> exchanges with matter
  - Produces a charged lepton :  $\nu_{\ell} + N \rightarrow \ell^{-} + X$
  - Lepton flavour is preserved
- ▶ Neutral Current (NC) :  $\nu_{\mathsf{x}}$ 
  - Z<sup>0</sup> exchanges with matter
  - ▶ Neutrino remains in final state:  $\nu_{\ell} + N \rightarrow \nu_{\ell} + X$
  - No charged lepton produced, harder to detect

# **Atmospherical Neutrinos: Charge Spectrum**





- Charge spectrum for atmospheric neutrino events before (left) and after (right) selection cuts (secondary particles and background noise removed)
- NC events lower in charge :  $\nu_e+$  atom  $\rightarrow \nu_e+$  atom\* (energy stays in the  $\nu_e$ )

Atmospherical Neutrinos 7/10

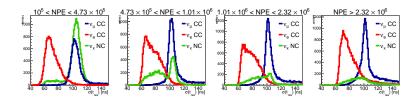
# **Atmospherical Neutrinos: First Hit Time**

- ► High precision measurement of the **photon arrival time** on the PMTs
- Gaussian smearing applied on the true MC hit time (3" PMTs)
- Residual time, t<sub>res</sub>, to be aligned with a realistic DAQ window:

$$t_{\text{res}}^{i} = t_{\text{hit}}^{i} - \left(\frac{n.R_{V}^{i}}{c}\right)$$

- **FHT**  $\equiv \sigma(t_{\text{res}})$  defined by the RMS of  $t_{\text{res}}$
- ▶ Discrimination variable for  $\nu_e$ ,  $\nu_\mu$  and  $\nu_x$  interactions

# **Atmospherical Neutrinos: First Hit Time**



Graph taken from *Atmospheric neutrino spectrum reconstruction with JUNO*, Oct. 2019 (ROOT stopped running on my computer...)

- ► **FHT distribution** for CC and NC interactions
- Separated by deposited charge (NPE)
- Variable used for event selection in order to discriminate neutrino types

## Next steps

- Optimization of selection cuts (including FHT)
- ► **Neuronal Network** implementation ??
- State of the art (JUNO experiment and atmospheric neutrino analyses/parameters)

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