PeV gamma ray searches with IceCube

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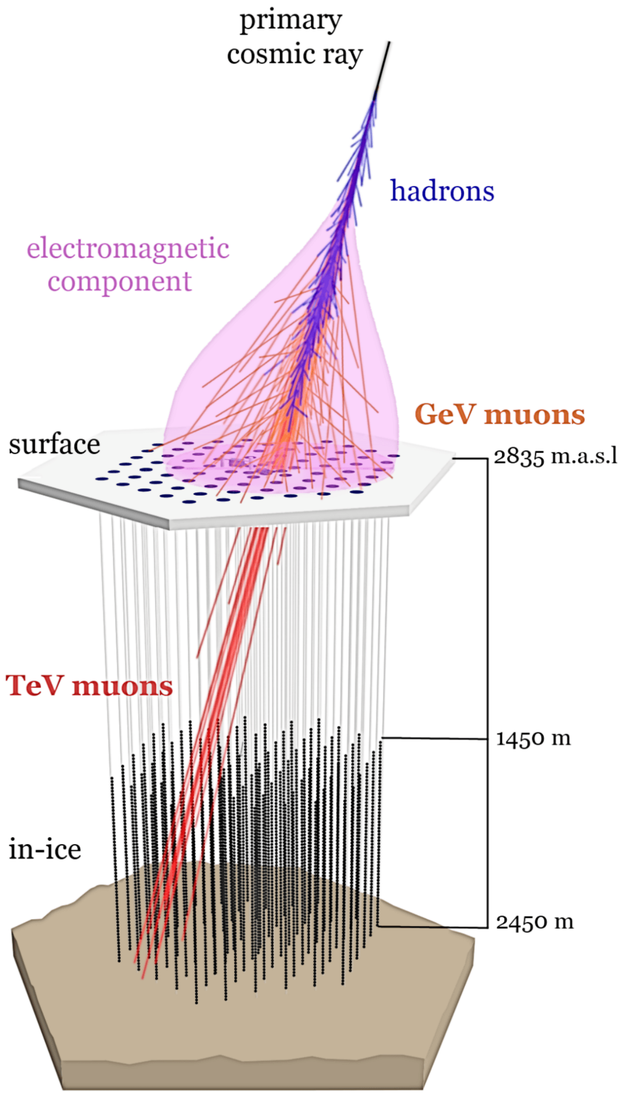
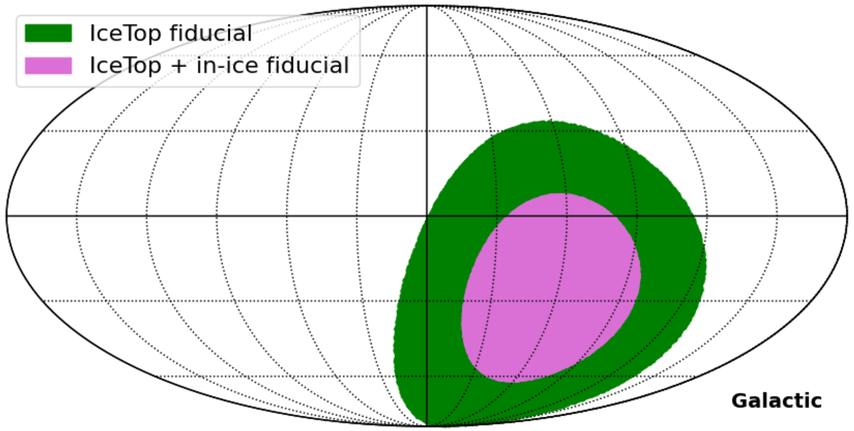
Detection strategy



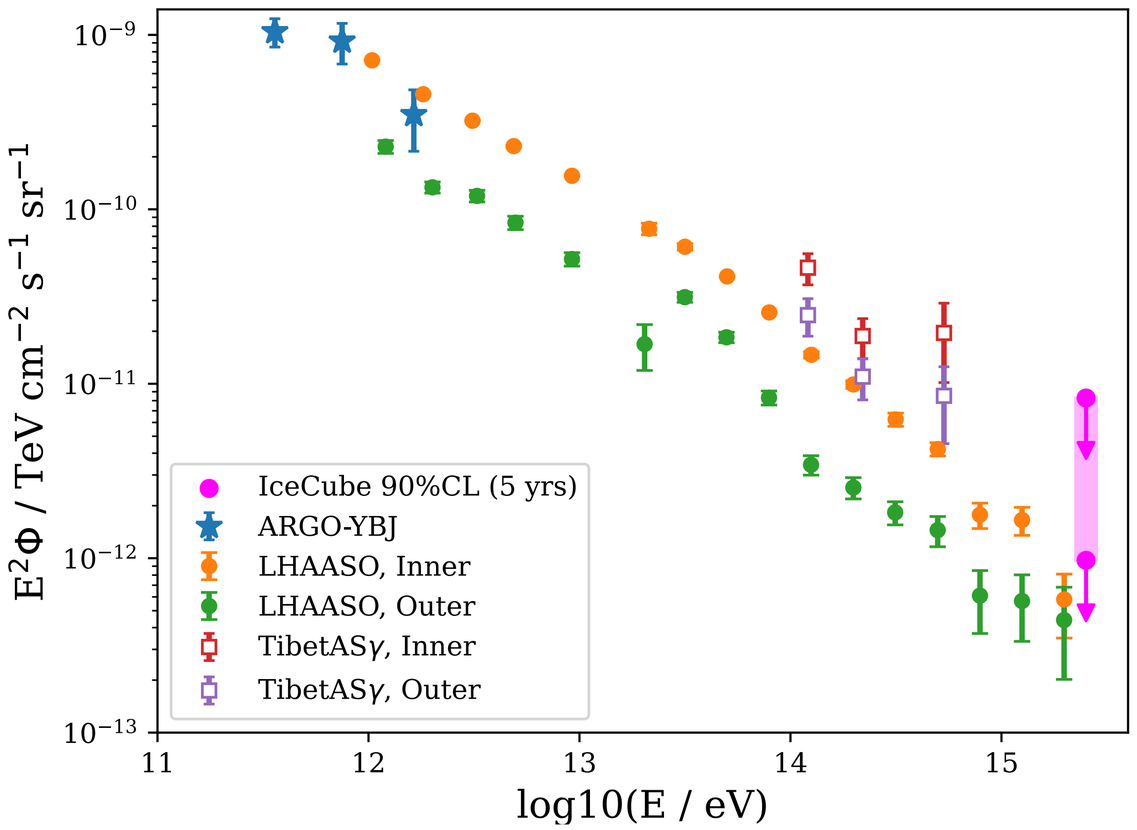


Use IceTop surface detector to reconstruct shower direction and energy

Use in-ice detector as veto. This reduces the field of view.

Gamma / hadron separation based on the muon content of the shower

IceCube searches for PeV gamma-rays



https://arxiv.org/abs/1908.09918

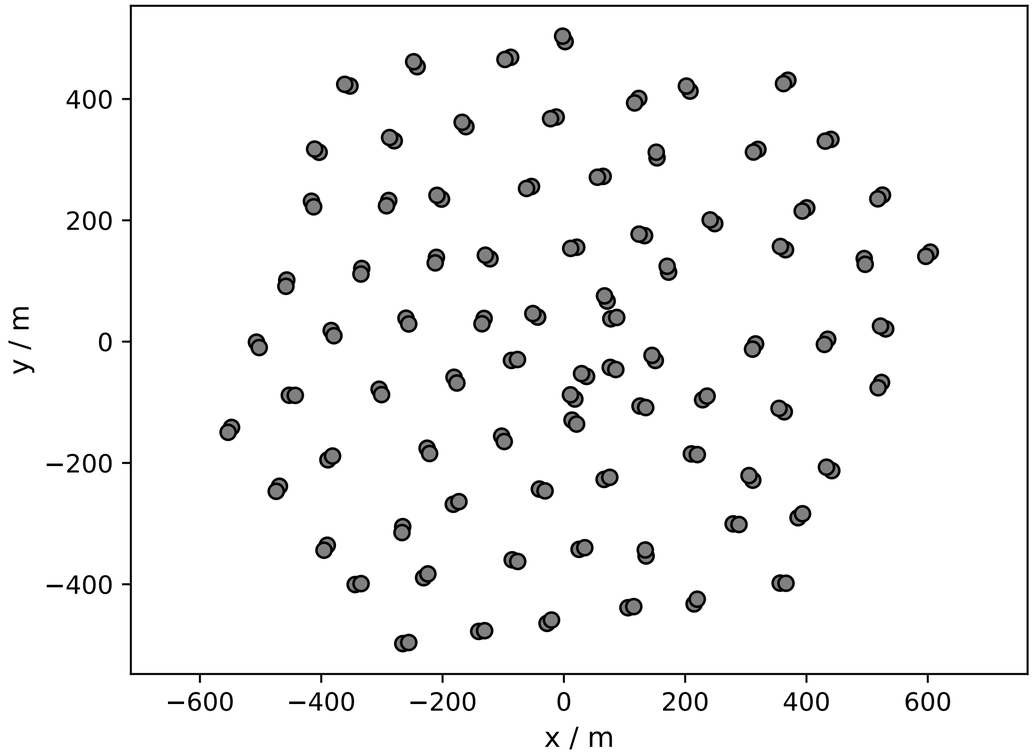
- Unique view of the Southern Sky
- Publication in 2020, using 5 years of IceCube data from 2011 to 2015
- No significant point-like source, upper limit on the diffuse gamma-ray flux at 2 PeV

Go from upper limit to measurement:

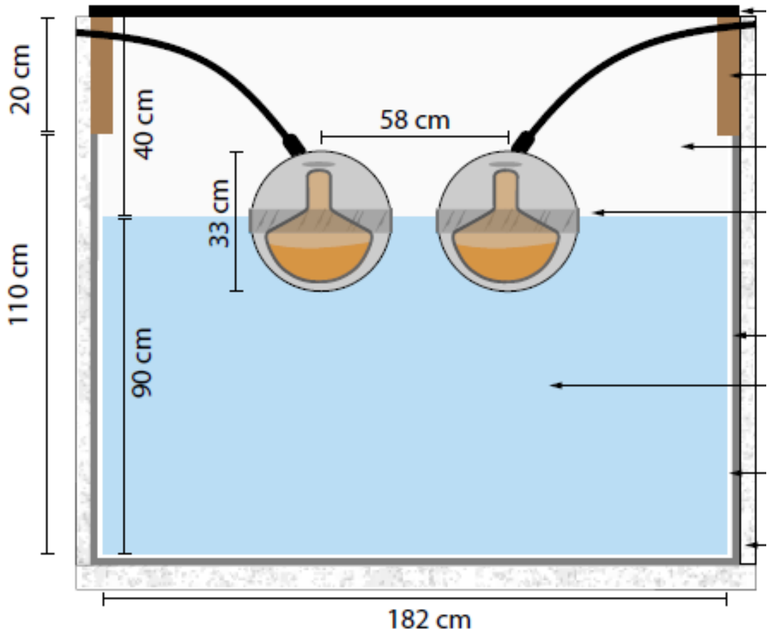
- 10+ years of data (snow accumulation to be taken into account)
- Better gamma/hadron separation
- Updated hadronic interaction model Sibyll2.3d

IceTop: the air shower detector

1 km^2 array at 2835 m a.s.l.



81 stations, with two ice Cherenkov tanks each, equipped with two DOMs per tank



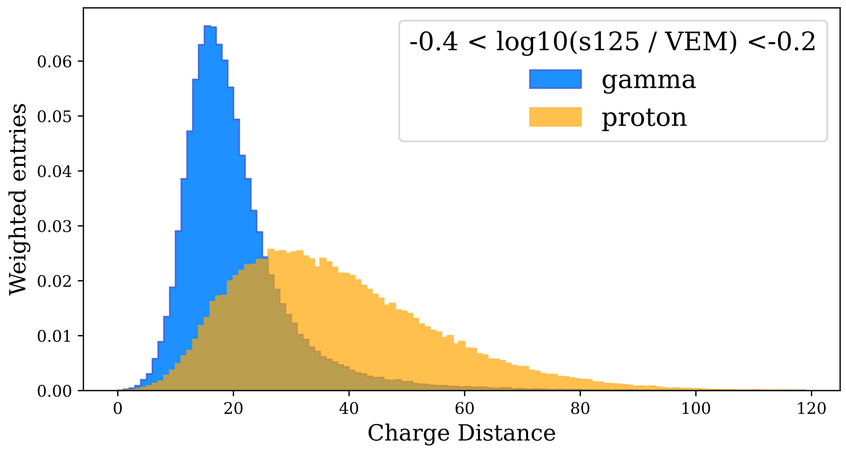
HLC: two tanks of a station have at least one DOM over threshold

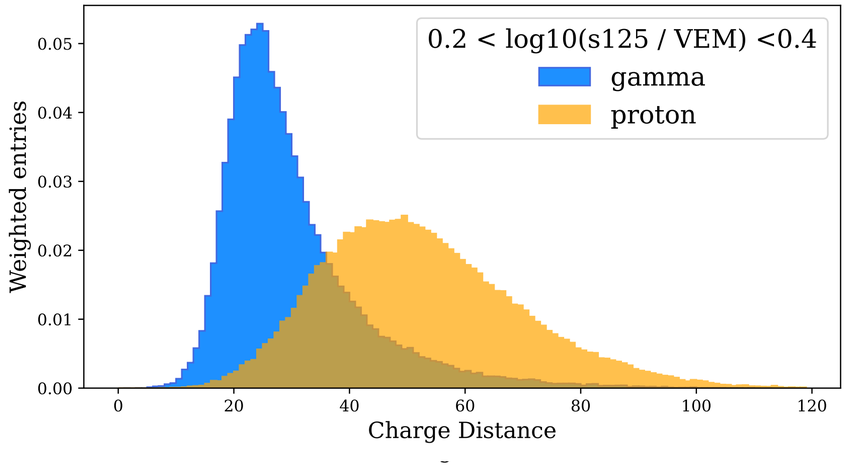
SLC: only one tank in the station has a DOM over threshold

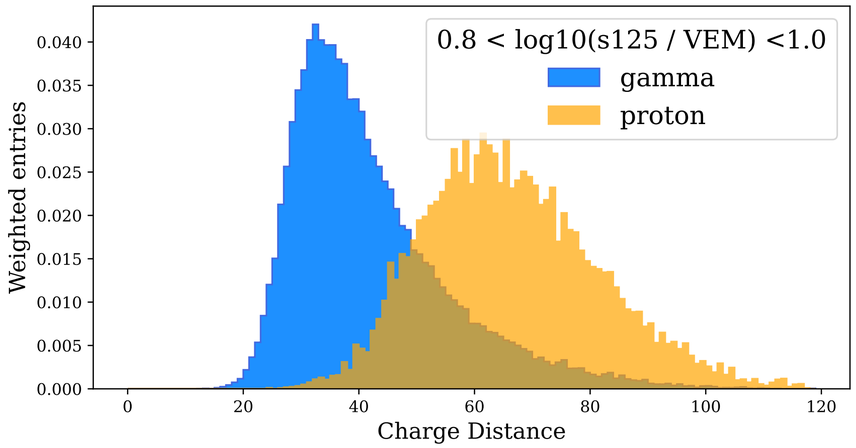
Gamma / hadron separation with IceTop

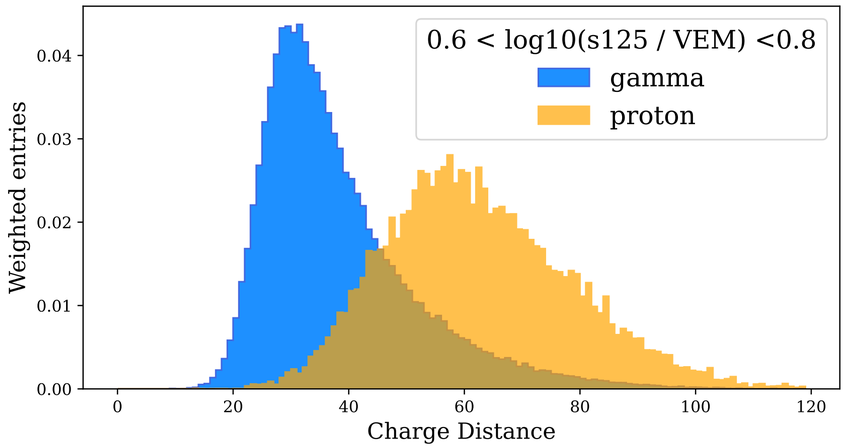
Proton MC 2012
Gamma MC 2012

$$\text{Charge-Distance} = \log_{10}\left[\sum_i \left(\frac{q_i^{\text{SLC}}}{0.1\text{VEM}}\right)^{\frac{R_i}{10\text{ m}}}\right]$$

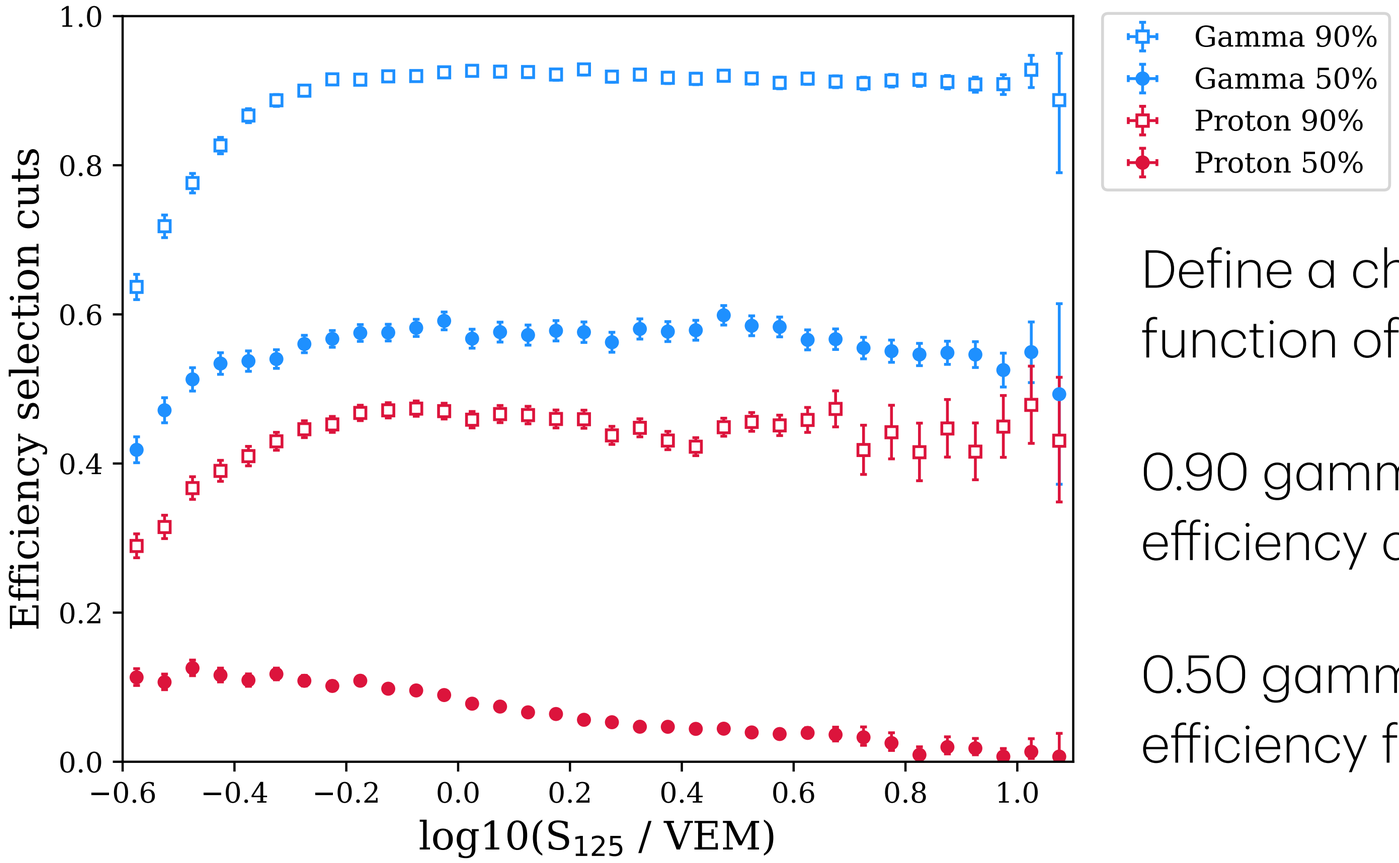








Gamma / hadron separation with IceTop

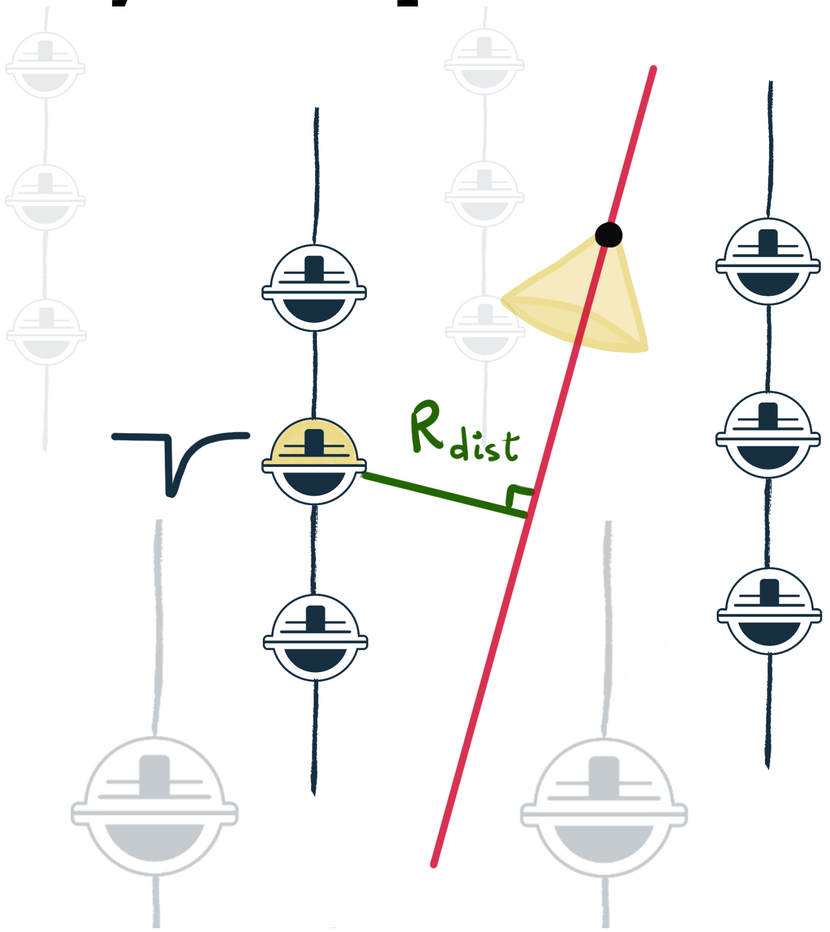


Define a charge-distance cut as a function of S_{125}

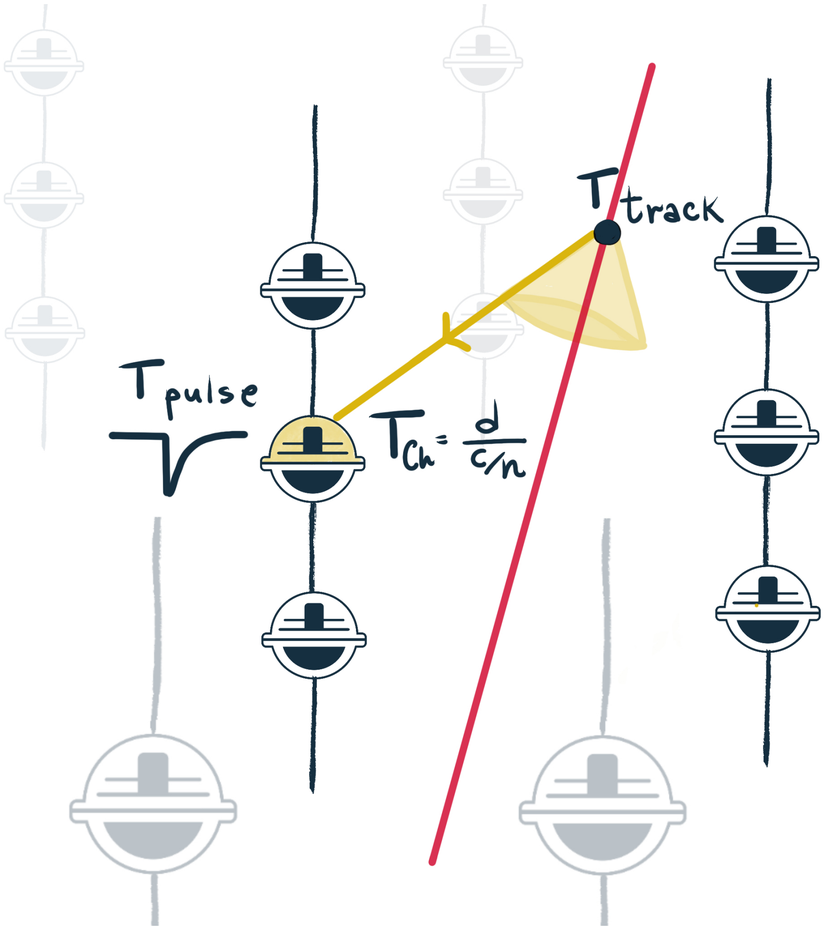
0.90 gamma quantile: proton efficiency at ~40%

0.50 gamma quantile: proton efficiency from 10^{-1} to 10^{-2}

Identify in-ice pulses correlated with air-shower muons



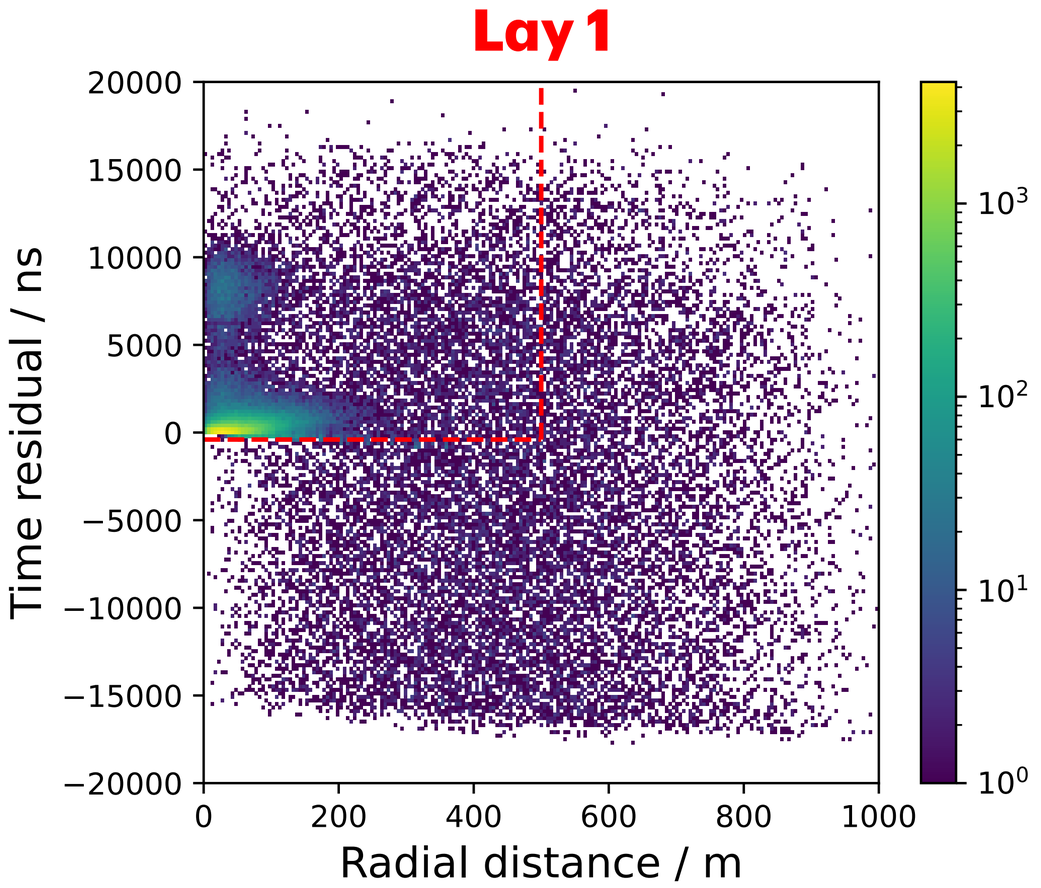
Radial distance



Time residual :

$$T_{pulse} - (T_{track} + T_{Ch})$$

Gamma / hadron separation with the in-ice array



Proton MC 2012
log10(E_{MC} /eV) = [15.1, 15.2]

Use in-ice array as a veto

Define phase space of DOM pulses correlated with the high energy muons

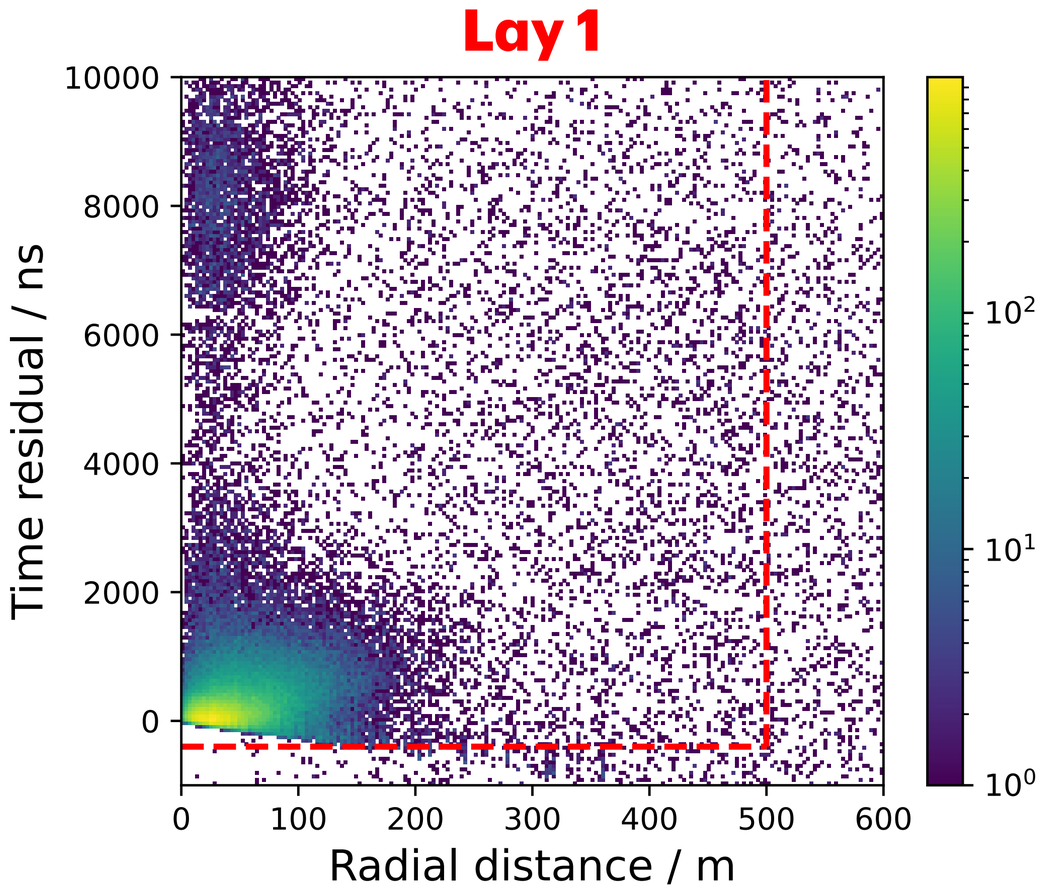
$$N_{lay} = 1,2,3,4$$

$$R < 300 \text{ m}$$

$$\Delta T > -400 \text{ ns}$$

If there is any charge deposit in the veto region, reject the event

Gamma / hadron separation with the in-ice array



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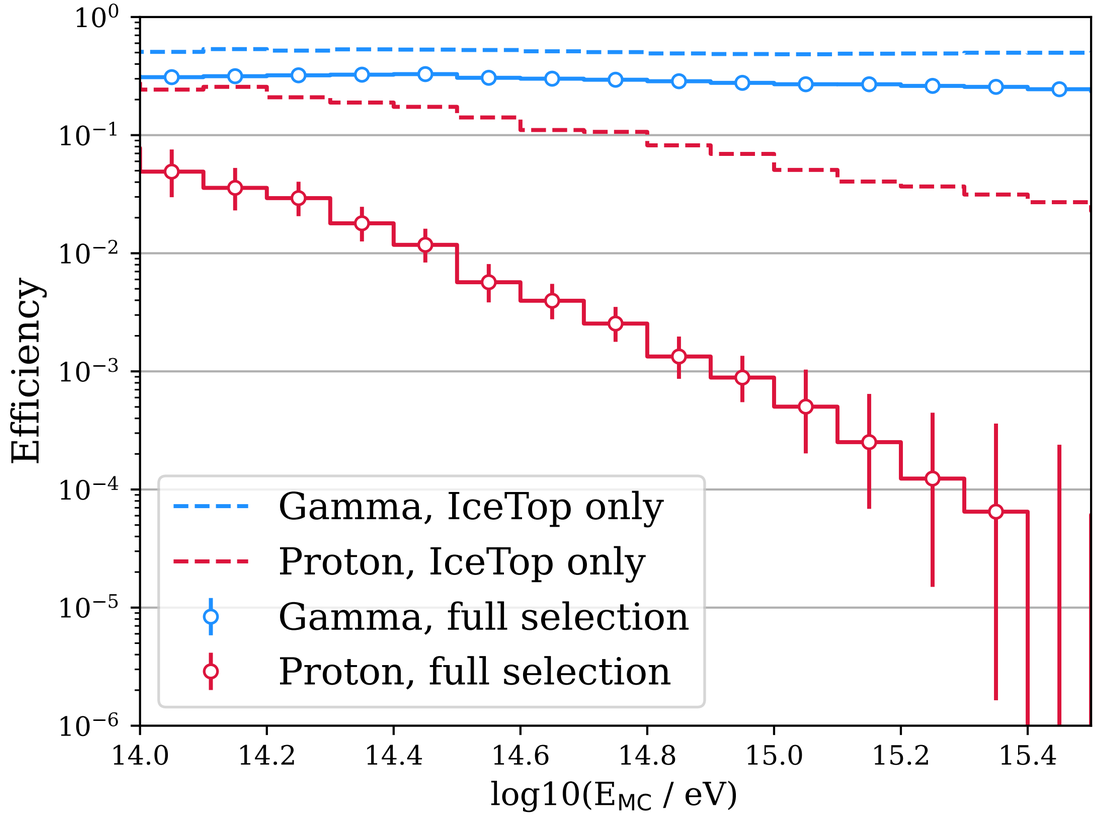
Gamma / hadron total separation



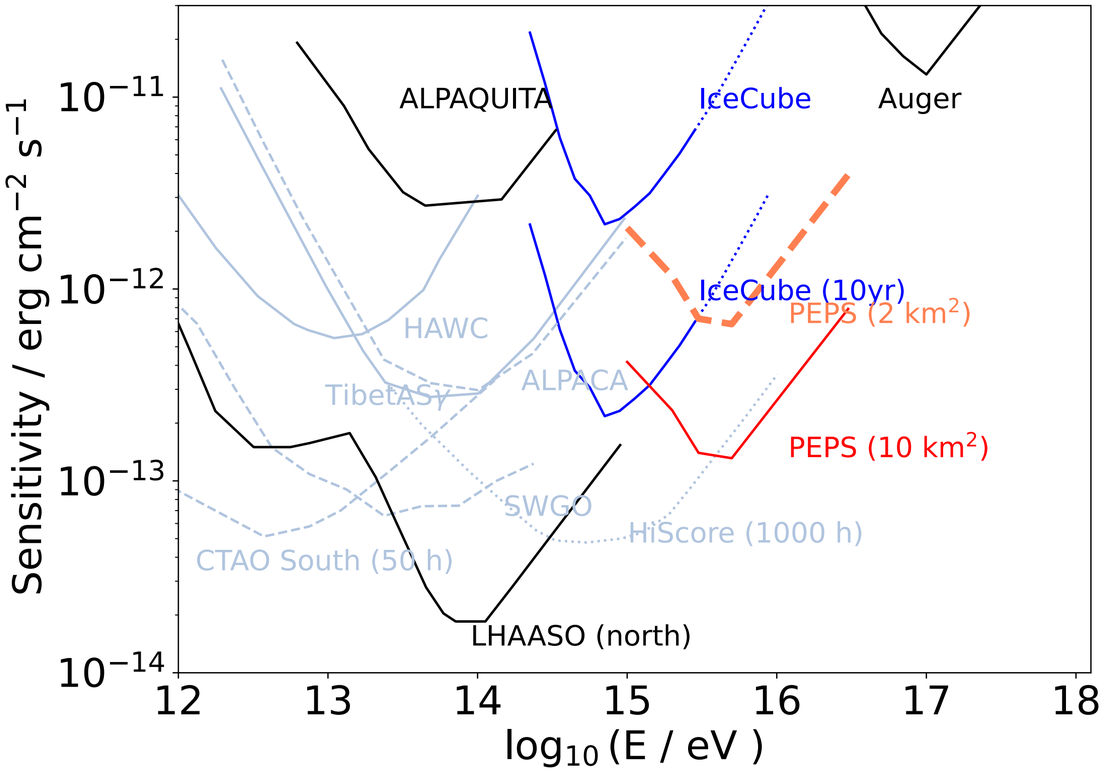
Proton rejection: charge-distance + in-ice veto

With the in-ice veto request, selection is complete

Achieved proton rejection is better than 10^3 above PeV

Conclusions

Preliminary IceCube 1yr sensitivity to point like source



- Unique view of the Southern Sky in the PeV regime
- Promising perspective with full 10+ years dataset
- Improved gamma / hadron separation

Outlook:

- Compute sensitivity to diffuse flux
- Produce simulations for different years