

# The Askaryan Radio Array

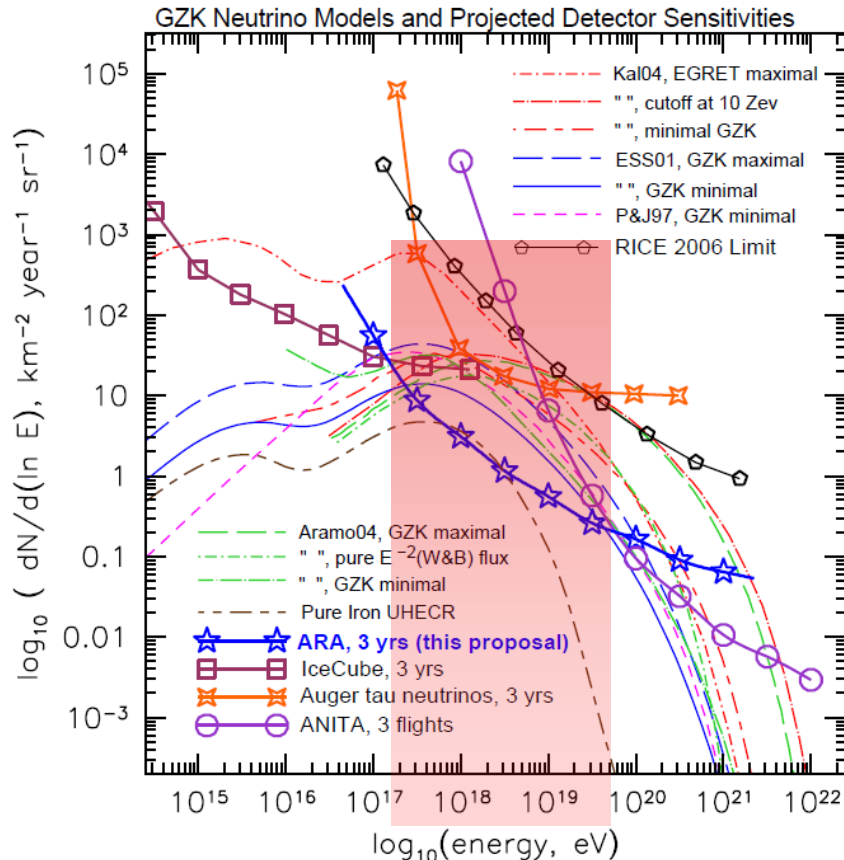
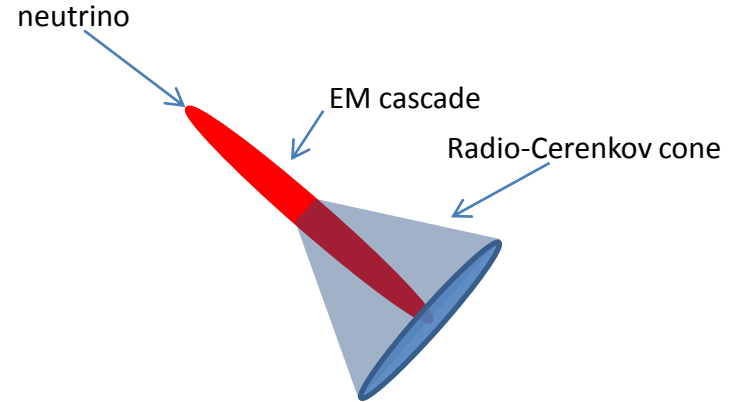
Kael Hanson, Thomas Meures, Yifan Yang

# Outline

- ARA
  - Motivation/goal
  - Detector geometry
  - Event reconstruction
  - The test bed
  - Power supplies
- Current IHE involvement
  - Digital data transmission system

# The motivation/goal

- Expecting  $<1$  GZK neutrino interaction per  $\text{km}^3$  per year
  - **Big detector needed**



## Want to detect neutrinos based on:

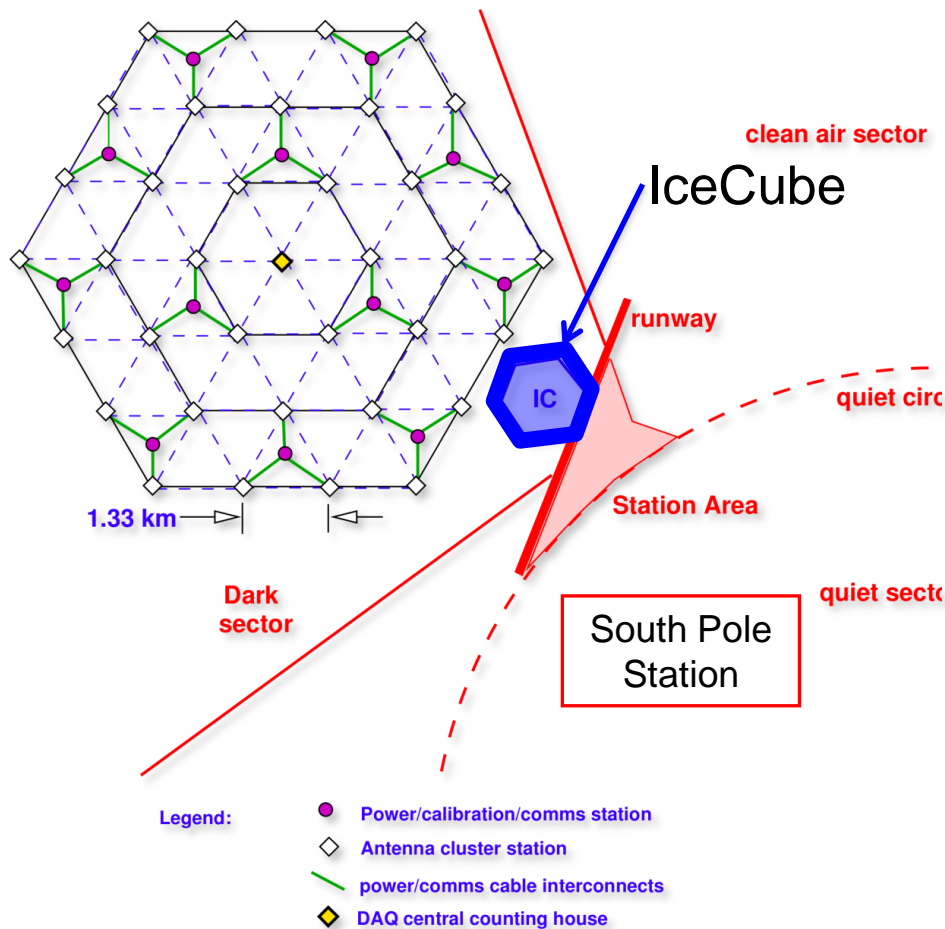
- Emission of coherent radio waves from neutrino-induced EM-cascades (predicted by Askar'yan, 1962)
- Verified at SLAC in 2000

## Why this detection method:

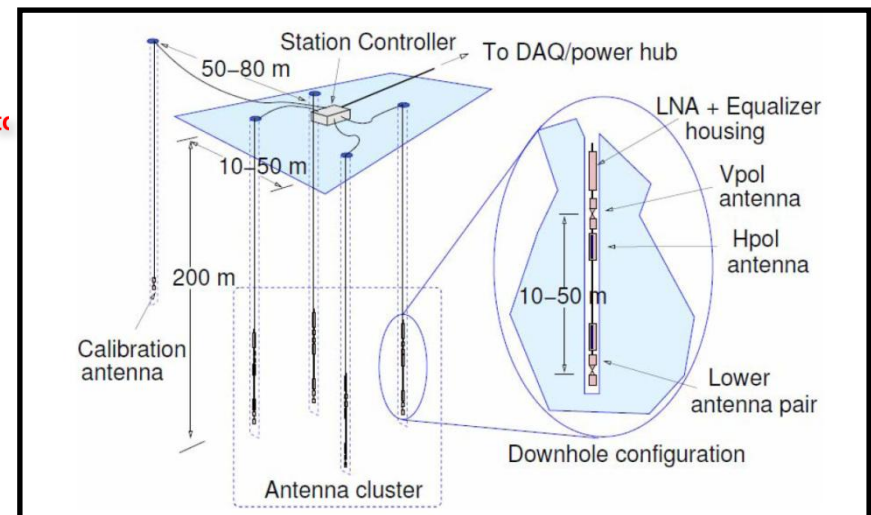
- Radio waves have high attenuation length in certain media (ex.:  $\sim 800\text{m}$  in ice)
- Radio antennas are cheap sensors compared to optical modules

# Detector geometry

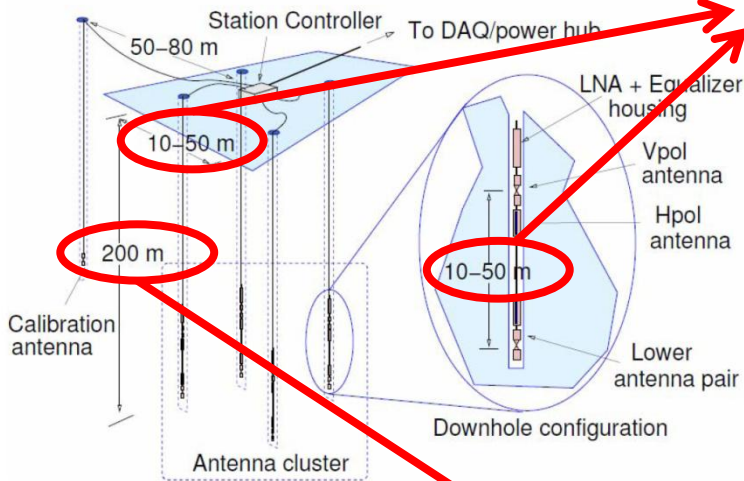
## Askaryan Radio Array



- 37 stations
- 80km<sup>2</sup> coverage
- **One station:**
  - 16 + 2 antennas  
(8 h-pol., 8 v-pol., 2 cal. pulsers)



# Detector geometry

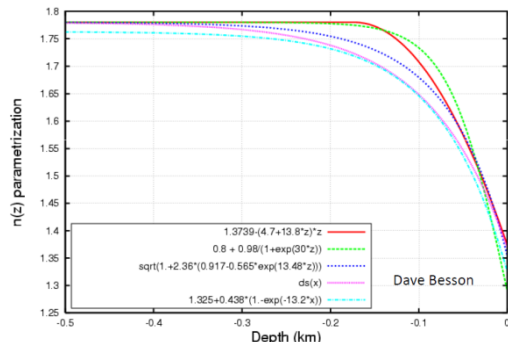


**20 m** - compromise between:

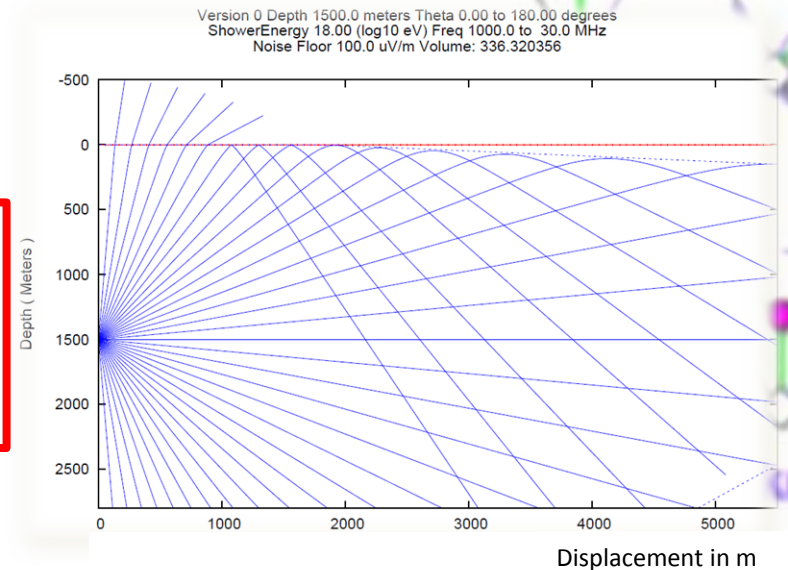
- Best angular resolution
- Maximum acceptable trigger rate

**1.3 km** station spacing:

- Attenuation length for radio waves in ice ~800m



- Ray tracing:  
due to  $n(z)$  dependence radio signal is bent down in the firn
- **200m depth**, accounting for drilling costs

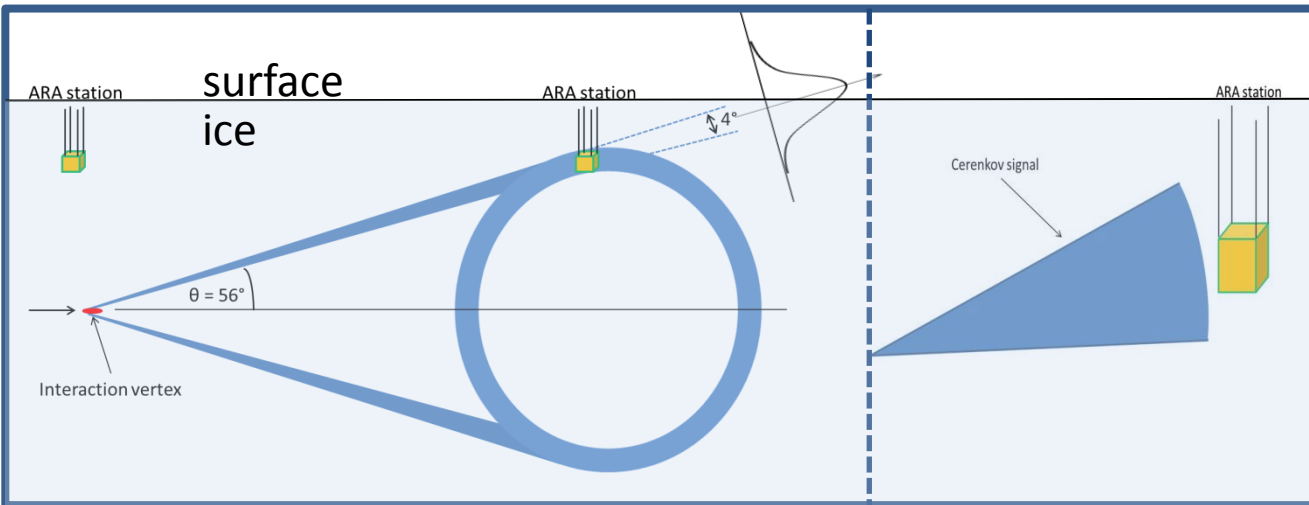
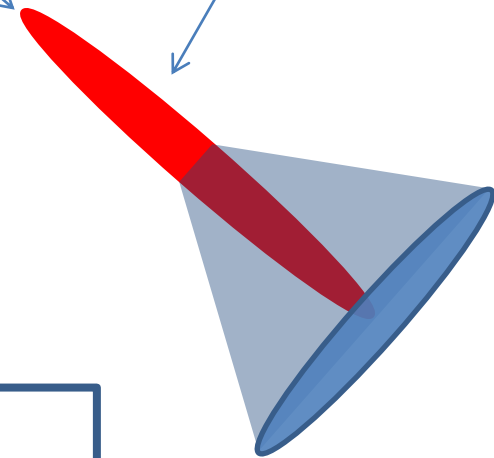


# Event reconstruction

- **Fraunhofer diffraction of the emitted signal**
  - Signal is Gaussian distributed over angle  $\Delta\theta$  around opening angle of the cone

neutrino

EM cascade



Reconstruction via grid search:

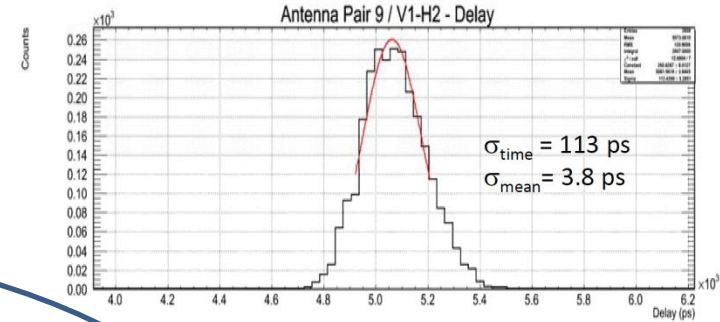
$$\chi^2 = \sum_i \frac{[\Delta t_i^{\text{obs}} - \Delta t_i^{\text{hyp}}(x_{\text{sh}}^{\text{hyp}})]^2}{\sigma_t^2}$$

$\Delta t$  = time difference between antenna triggers

# The test bed

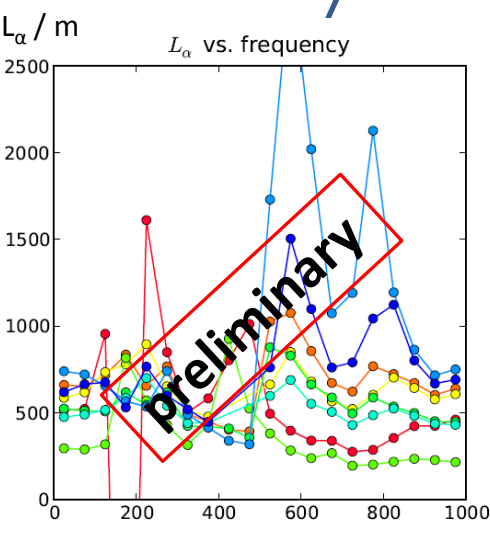
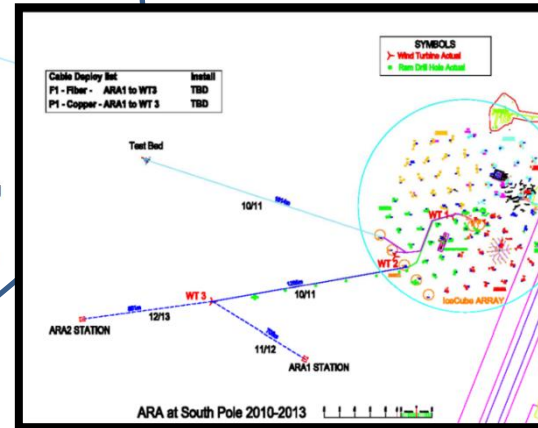
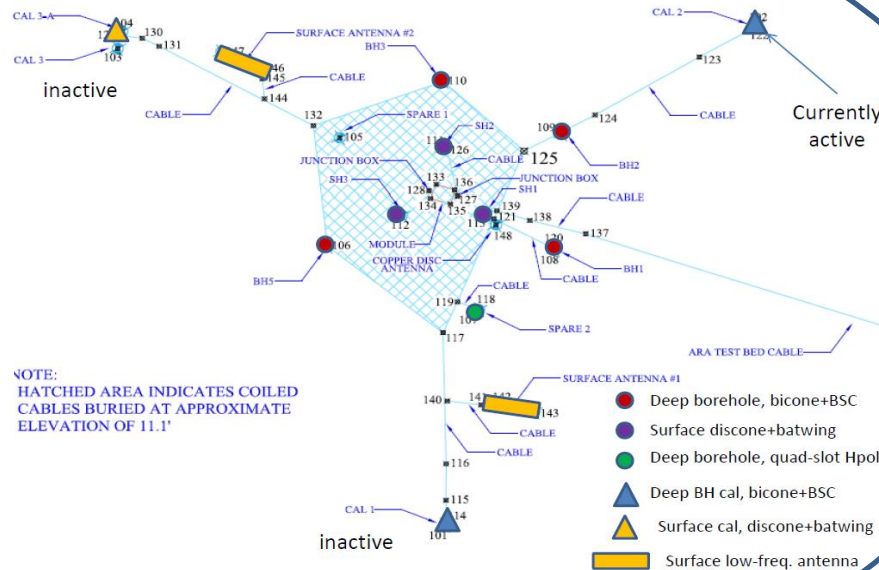
## Different goals:

- Study background noise
- Study reconstruction possibilities
- Measure frequency dependent attenuation in the ice



## ARA TEST BED ASBUILT

SURVEYED ON 12/31/10 & 01/05/11



# Power supplies

Move to "green" energy



Photos by Mark Dierckxsens

Estimated consumption: 300W / cluster (3 stations)  
 Future plan: Use only regenerative energies

### Why is it important:

- To be autonomous
- Technically difficult to distribute all the power from the south pole station

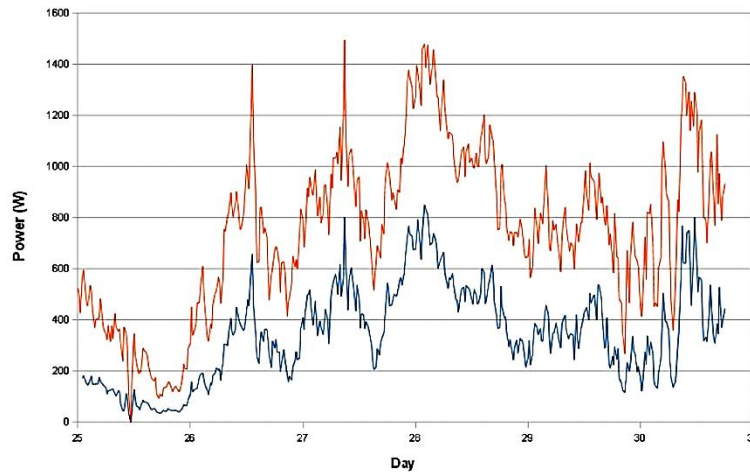
### Status:

- 3 wind turbines + batteries deployed for test measurements

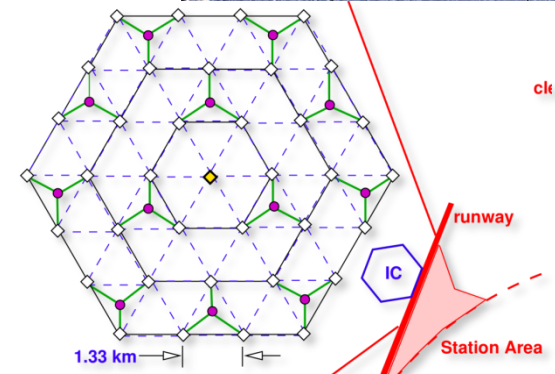


Photo by Ken Ratzlaff, KU

Hummer Turbine Power for January 25-30, 2011



— Predicted Power  
 — Measured Power





# IIHE work



# The motivation

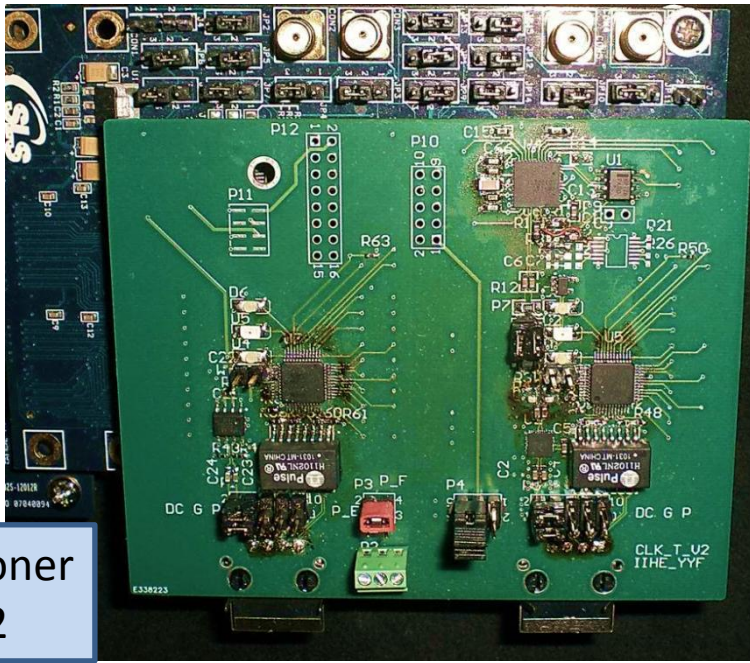
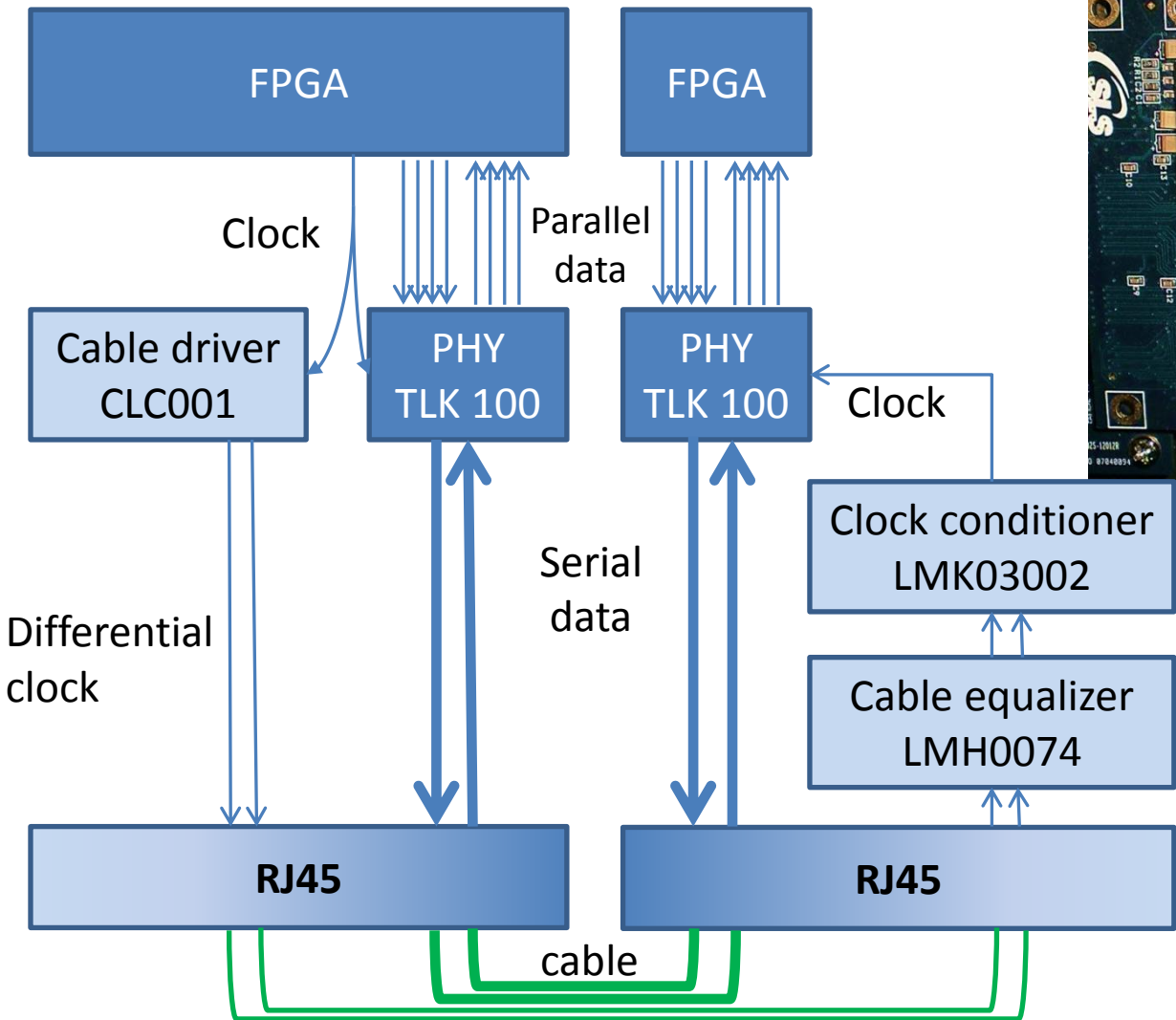
## Collaboration status in August 2010:

- analog data transmission via copper cable
- **But there are disadvantages:**
  - EMI problems
  - Bulky cables
  - Problems with timing calibration (high jitter)
  - Too much power needed for transmission

## Proposed alternatives:

- Digital over optic fiber -> **rejected at that time because of high costs and fragility**
  - Digital over copper:
    - Only small cable needed
    - Low power consumption possible
- } **Developed in Brussels**

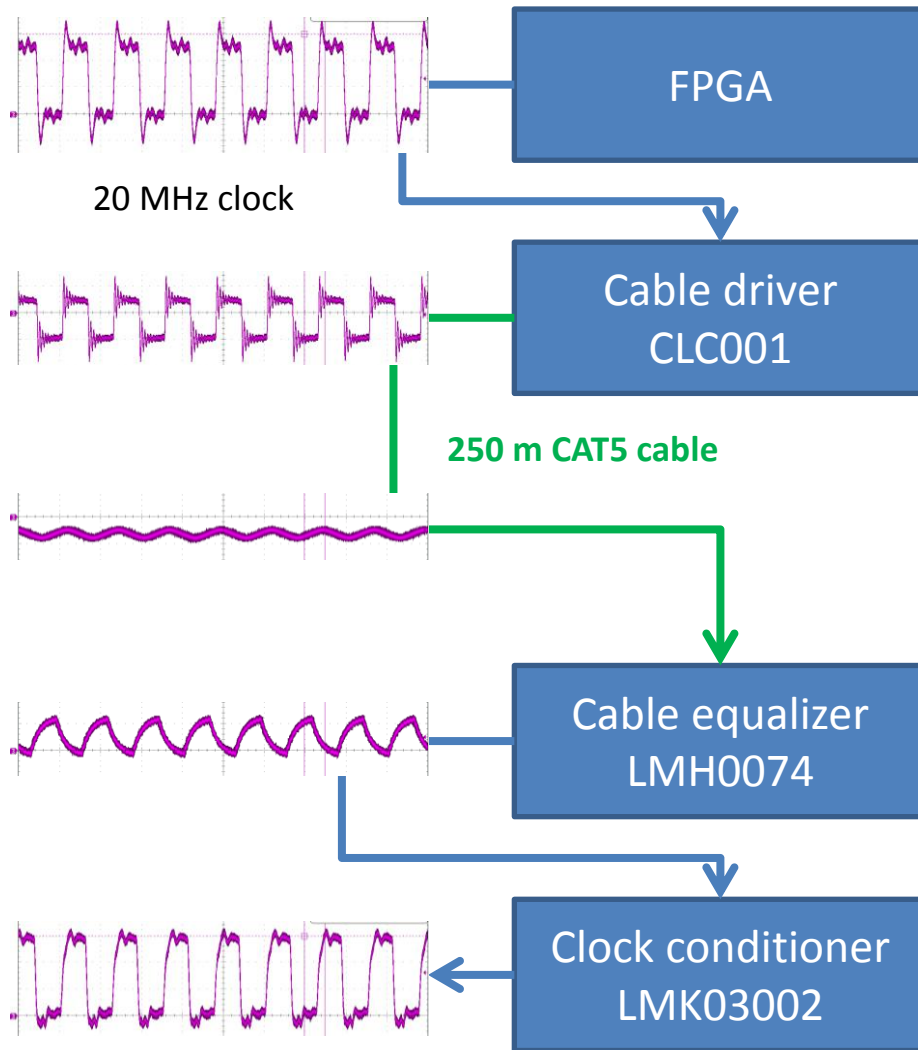
# The setup



- TLK100**
- 10/100 Mbits/s Ethernet phy, 25MHz input clock
  - **In current design:** driven by 20MHz clock
    - 8/80 Mbits/s
  - Power consumption: 200 mW

# The setup

## Clock transfer system



### Cable driver:

- Amplifying the incoming clock to transfer it
  - 400 ps rise time, 25 ps output jitter
  - 1 Vpp output
  - Power consumption: 520 mW

### Cable equalizer:

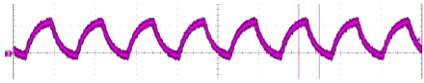
- equalization
- DC restoration
  - 750 mVpp output
  - Power consumption: 255 mW

### Clock conditioner:

- Loop filtering
- Jitter cleaning
- Clock distribution
  - 200 fs output jitter
  - Power consumption: 578 mW

# The setup

## Clock conditioner

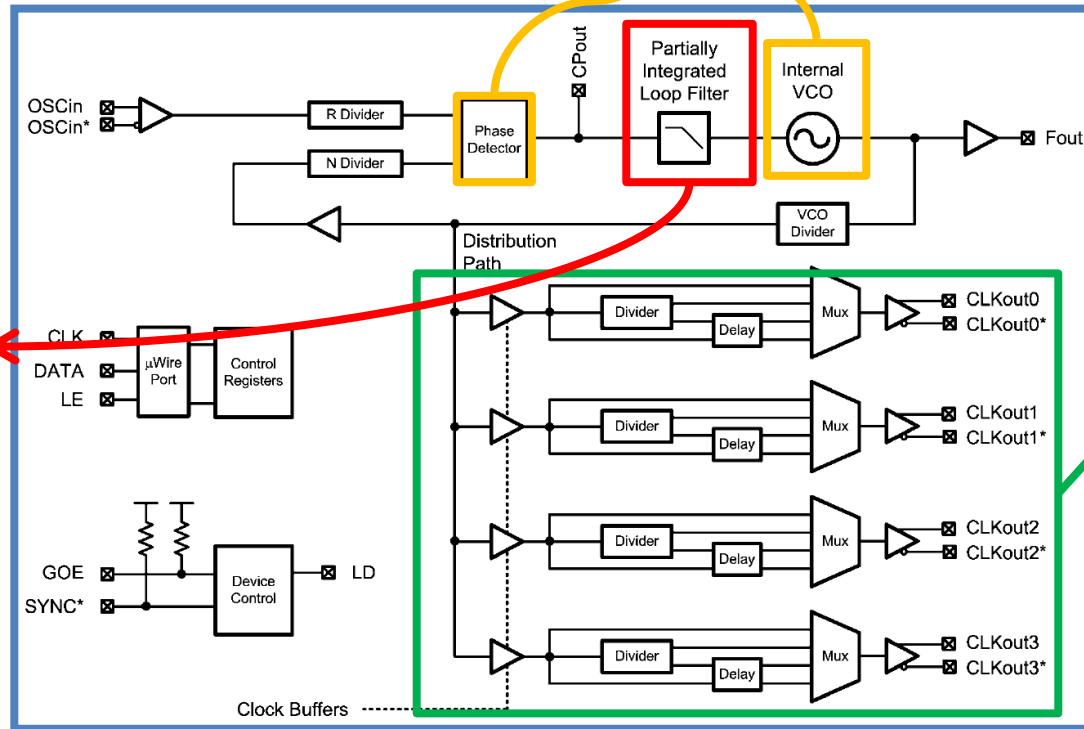


Makes clock detection easier for VCO

### Voltage Controlled Oscillator + phase detector

- VCO generates new clock
- Clock frequency is adjusted until phase detector doesn't find a difference to OSCin

"Fresh" clock with correct frequency supplied to PLL

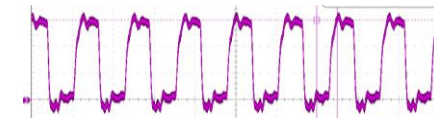


### Loop filter:

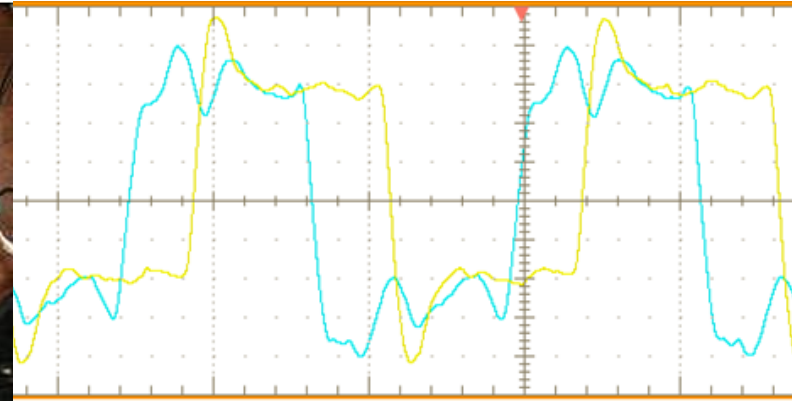
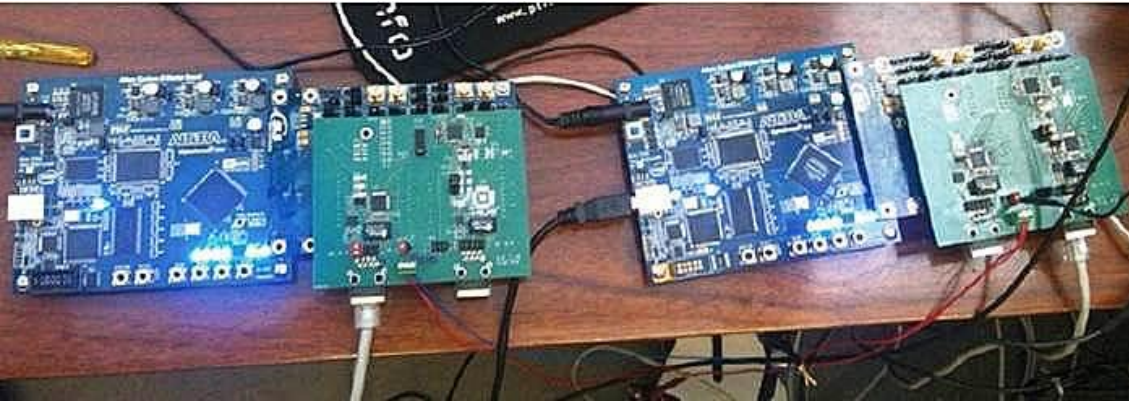
- 4 pole low pass, to remove high frequency noise

### PLL:

- Clock distribution
- Frequency modulation
- Phase adjustment



# Test results



Precision of period }  
 shift between clocks →

Description	Mean	Std Dev	Number of samples
Period1, Ch1	50ns	16ps	239988
Period2, Ch2	50ns	27ps	239976
TIE1, Ch1	0s	26ps	240000
TIE2, Ch2	0s	50ps	239988
Skew1, Ch1, Ch2	-8ns	52ps	239988
IP-failure rate (down to -40°C)	1.2E-8/32bits		4.3E11 (72 hours)
Power consumption	~2.3W		

# Future plans

## Presented system was not accepted by the ARA collaboration:

- Trigger rate raised from  $\sim 4\text{MHz}$  to  $\sim 40\text{MHz}$
- Copper prizes raised (well shielded cable is very expensive)
- Power consumption got less important
- RF over fiber is the current choice
- **Digital system might still be interesting for other experiments**



## Future plan:

### Digital over optic fiber:

- One fiber cable for both directions carrying data and clock
- Speed:  $\sim 2\text{GBits/s}$
- Clock recovered from the data (CDR – clock data recovery)
- Power consumption  $< 700\text{mW}$  (for one system)





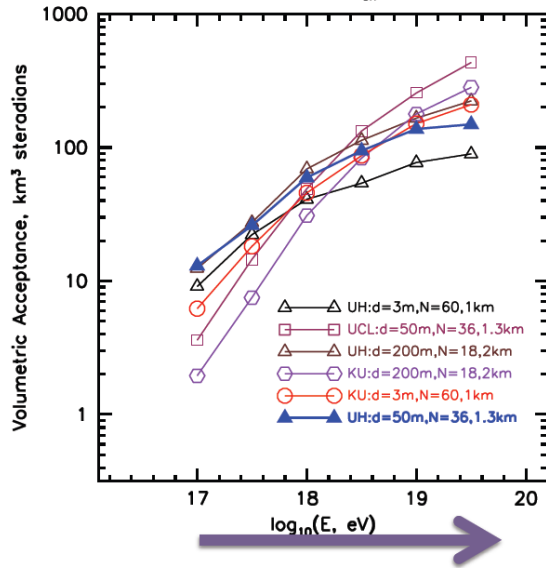
# Summary

- ARA:
  - ARA detector will be able to efficiently detect GZK neutrinos
  - Test bed and power supplies are deployed to investigate possibilities
  - Future: Two ARA-stations will be ready this summer, to be deployed in the coming two years
- IIHE group
  - Currently involved in Wireless communication, GPS, rubidium clock
  - Digital data transmission system was assembled and successfully tested at the IIHE
- Future plans:
  - Digital data transmission via optic fiber
  - Firmware design for ARA electronics

# Backup

### Detectable shower energy

Iceray Comparative  $V_{\text{eff}} \Omega$  Study



In Peter's code :

$$E(\omega, R, \theta) = E(\omega, R, \theta_C) e^{-\ln 2 \left[ \frac{\theta - \theta_C}{\Delta\theta} \right]^2}$$

$$\Delta\theta = 2.7^\circ \left( \frac{500\text{MHz}}{\nu} \right) \left( \frac{E_{\text{LPM}}}{0.14E_0 + E_{\text{LPM}}} \right)^{0.3}$$

$$E_0 = 10^{17.5} \text{eV} \sim 10^{19.5} \text{eV}$$

where  $E_{\text{LPM}} = 2 \text{PeV}$

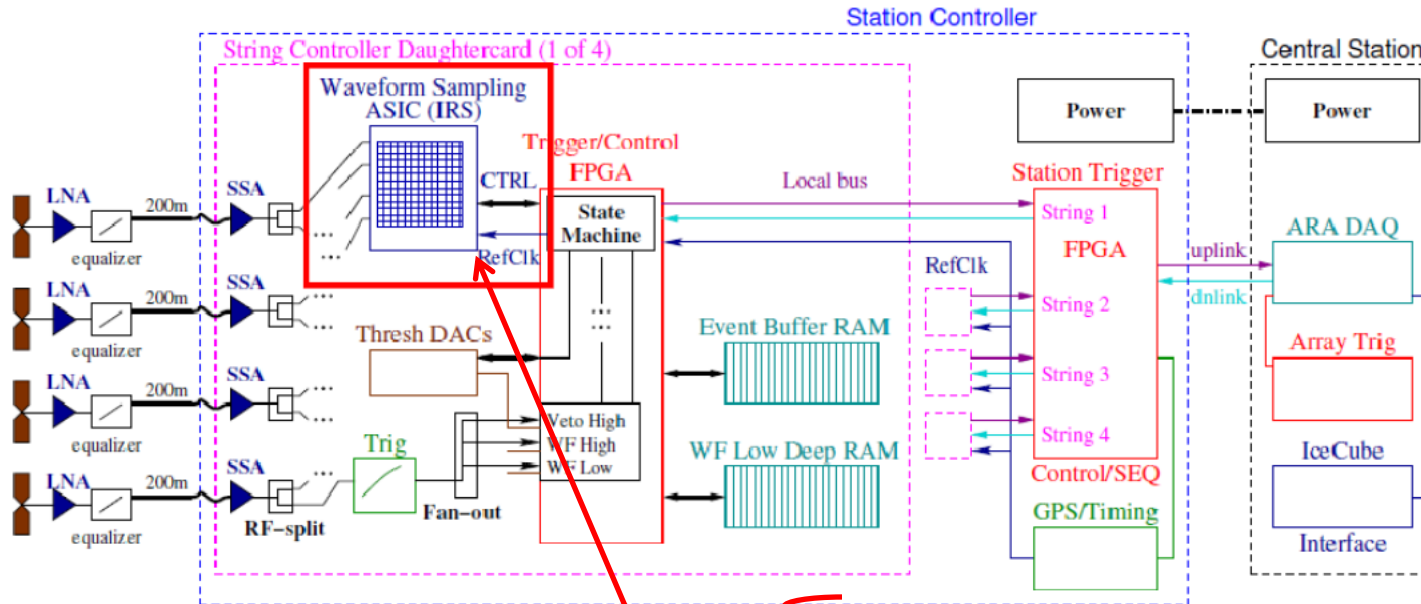
**Cherenkov Radio Pulses from EeV Neutrino Interactions:  
The LPM Effect, J. Alvarez-Muniz and E. Zas (1997)**

Hsin-Yi Tu

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# Station electronics



2 trigger levels:

- Coincidence: 5 of 16 antennas
- Pattern match:  
Can it be an RF signal?
- **Trigger rate: 10 – 50Hz**

32768	Samples/channel
8	channels
9	Bits resolution (12 bits logging)
Up to 4.5	GS/s
16	$\mu$ s trigger latency (at 2 GS/s)
120	$\mu$ s read-out time (512 samples)
1	kHz maximum trigger rate
250	ns blocked data -> <b>No dead time!</b>

# Outline

- The setup
  - Data transfer system
  - Clock transfer system
- Operation, tests & results
  - Clock precision
  - Power consumption
  - IP-failure rate
  - Behaviour for low temperatures