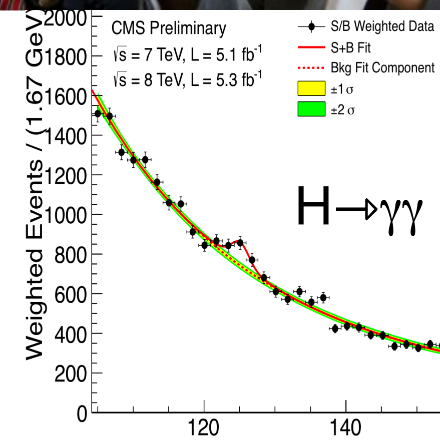
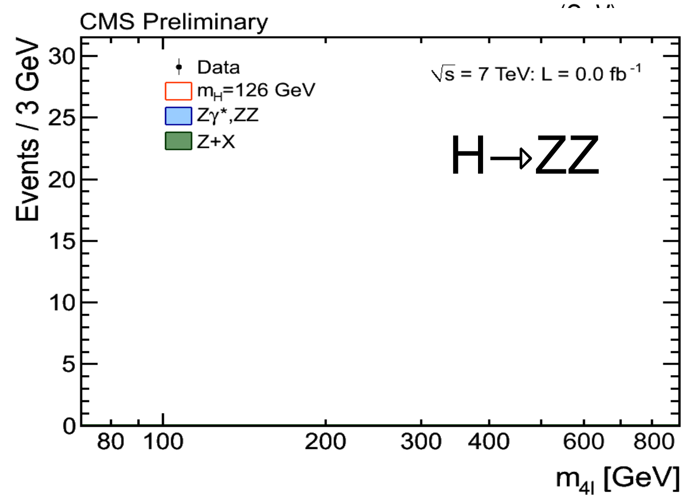
CMS:
Readiness for new energy frontier
Planning the long term future

Outline

- Following the new boson discovery the success of Standard Model of Nature is paradoxically highlighting its limitations.
- There are fundamental questions which are not yet answered by the Standard Model.
- With the LHC restarting at 13 TeV and opening a new energy domain we expect to be able to address some of these fundamental questions.
- I will show how we have prepared this exciting run and how we are preparing for the long term future



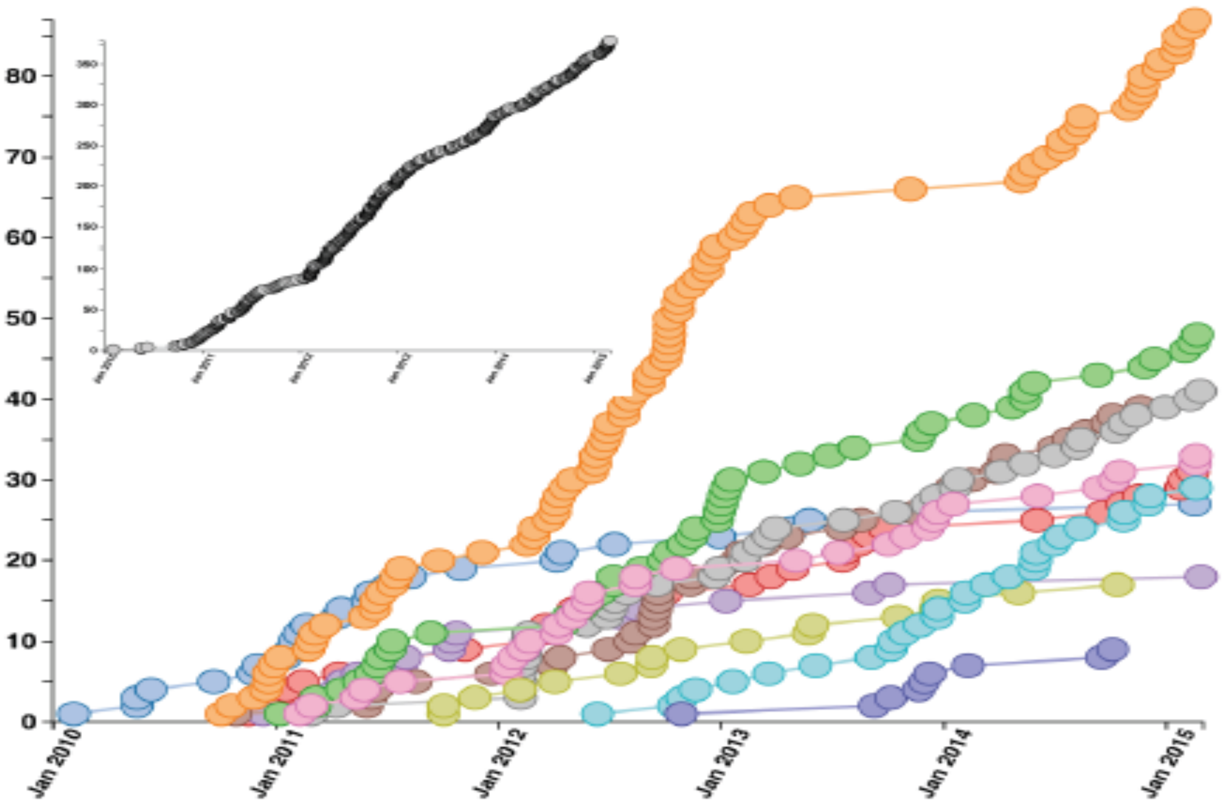


A rich legacy from Run1

Show all | Total | QCD | Exotica Searches | Supersymmetry | B Physics | Electroweak

Top Physics | Heavy

> 400 papers submitted!



Pub rate steady, ~2.5/week

403 papers submitted:

+23 Cosmic ray based

+32 ready for CWR or later

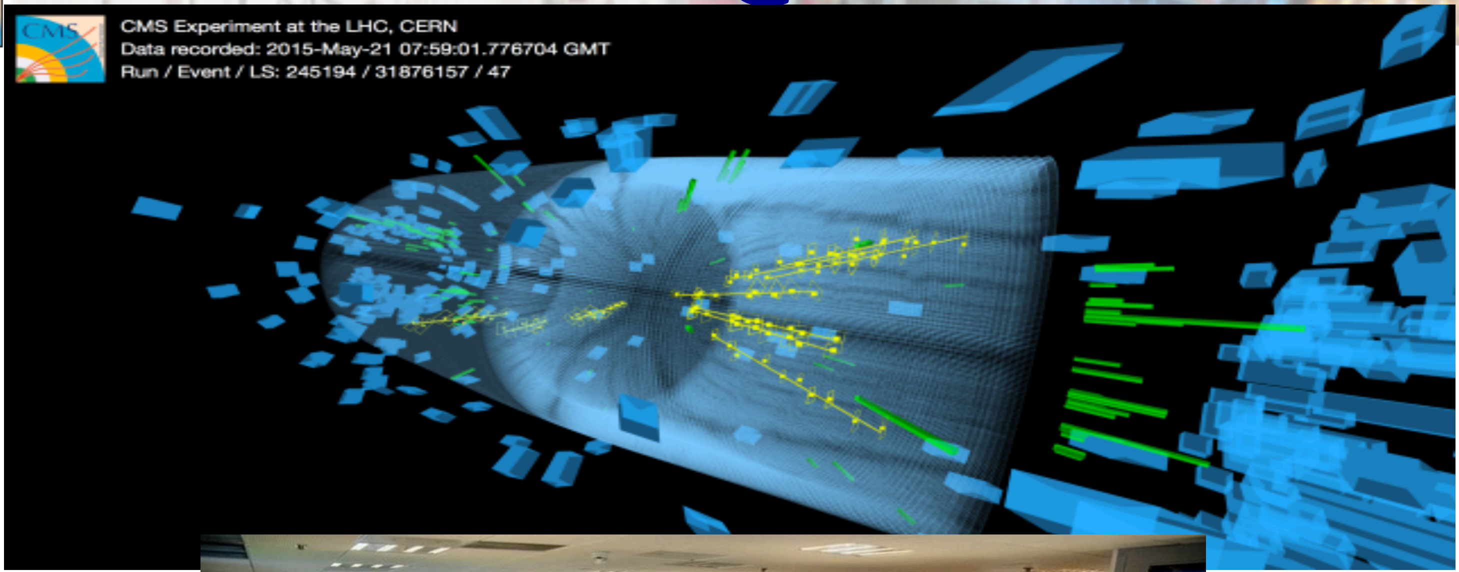
+17 PubDraft

In review process (128):

84 GoingToPreApp or higher

44 in PAG review

First collisions @ 13 TeV



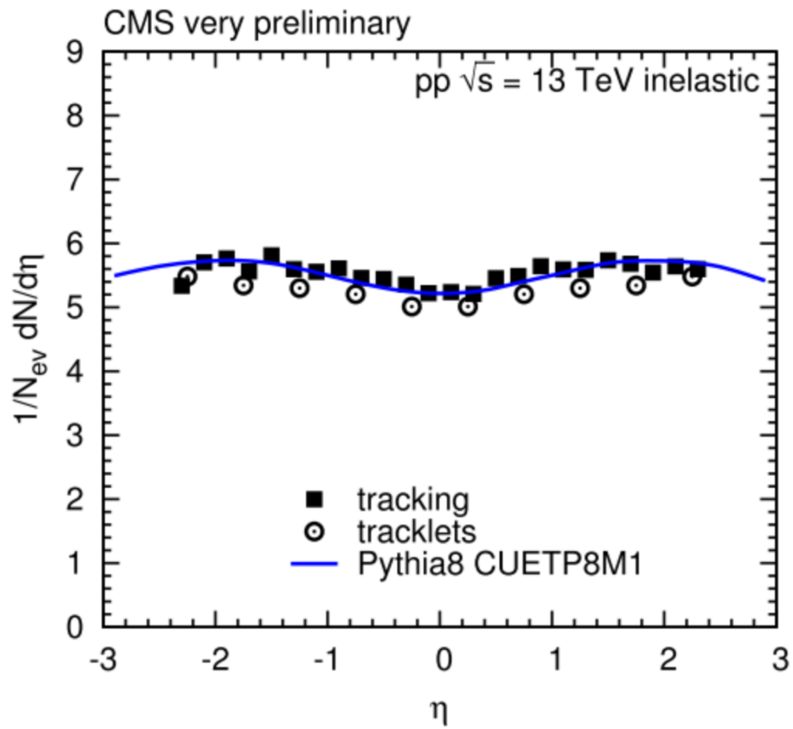
Readiness for LHC Run2

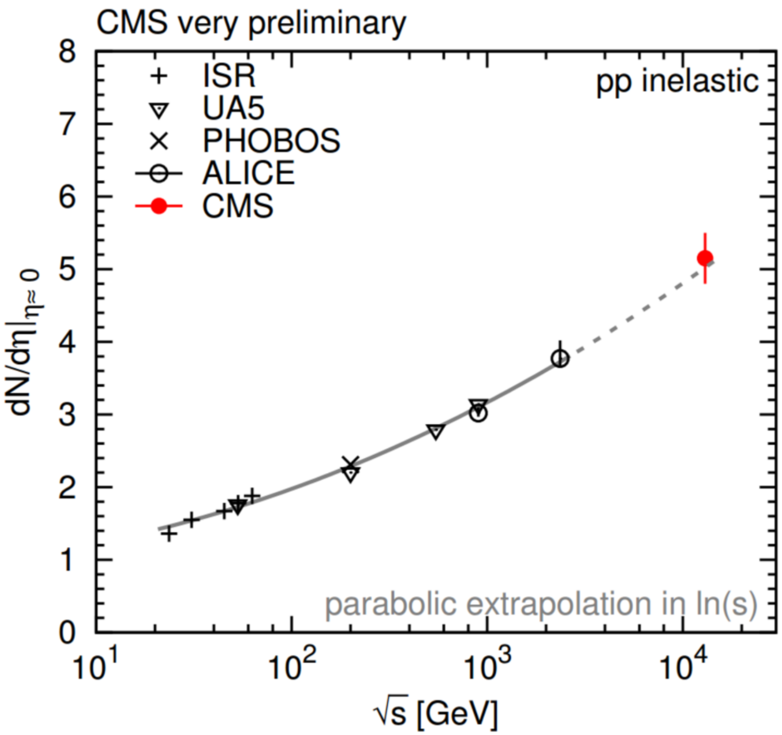
- CMS has undergone major changes during the Long Shutdown, amongst other:
  - Added two new endcap planes of muon chambers
  - DAQ: changed the event building fabric and the Timing and control system
  - Replaced the readout devices of the Outer Hadron and the Forward Hadron calorimeter
  - Validated the readout at -20 degrees for the tracker and pixel
  - Implemented the first stage of the Upgraded level 1 calo trigger
- The setup has been re-commissioned using cosmic ray in progressive steps since July 2014: running 24/7 since early February
- During the preparation for the re-start we realized that the cold box feeding liquid He to our magnet had been polluted by the oil from the compressors: swift action has allowed recovering of the system: we are being re-commissioning the magnet in view of the full intensity run of LHC.
- The data delivered during the first two weeks of LHC re-commissioning with beams at low luminosity have been collected with B=0
- The goal for the 2015 campaign is to collect ≤ 10 fb^{-1} of luminosity of proton collisions (a period fo Heavy ion collisions is foreseen for the last month)

...and we are ready

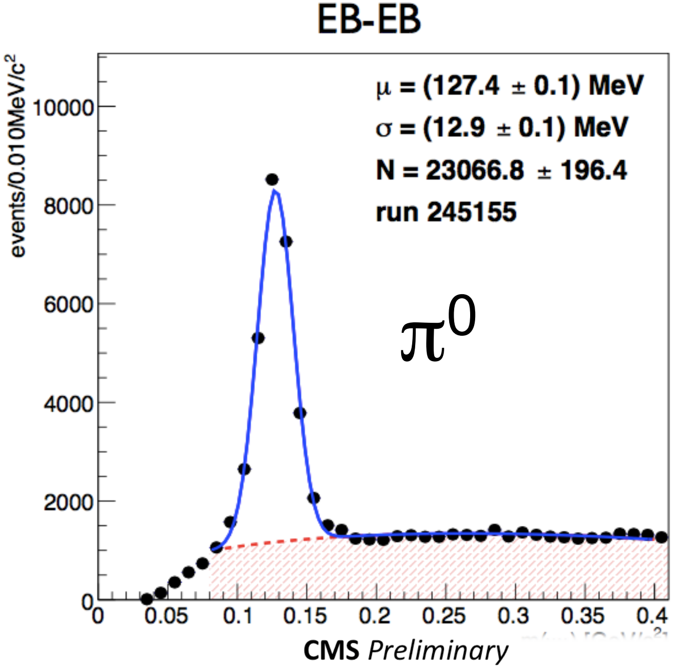
- We were ready on day 1 to do physics ...and we did despite our magnet issue

A demonstration of physics readiness of the whole system :
Preliminary dN/dη results @ 13 TeV (with B=0): obtained during first 'tuning' collisions: a demonstration of readiness, the data collected with B=0 will be completing the measurement to be done at full field.

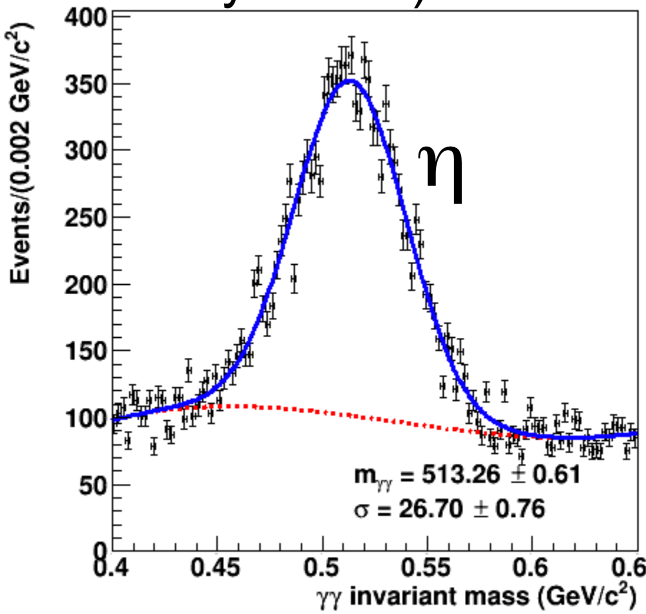


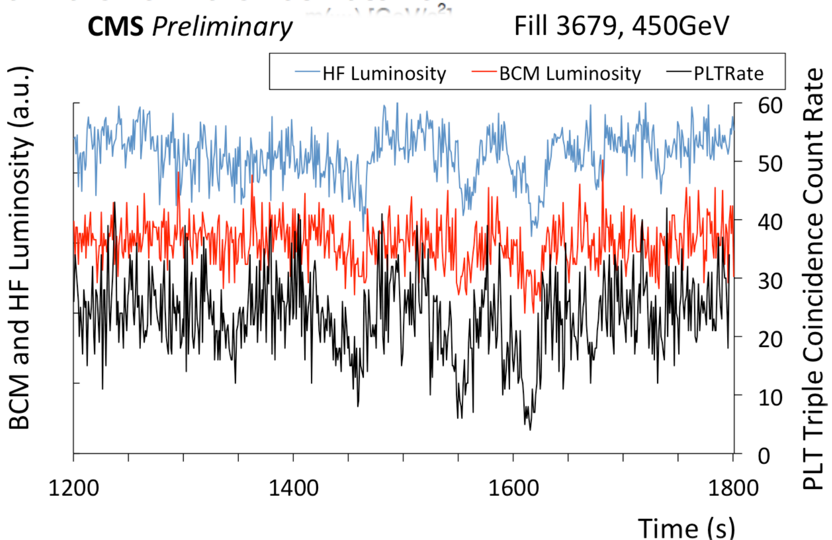


First data: example of performance



Online plot of ECAL calibration stream (central value affected by B field)





3 luminosity monitors being commissioned

The Physics objectives:

The next years are for frontier exploration

The past (pre July 4th,2012)

Unified gauge theories of weak and electromagnetic interactions provide an attractive framework for the interpretation of weak-interaction phenomena.[1] Such theories are universal in the prediction that existing data explore only the low-energy tail of a spectrum of yet-to-be-discovered particles. The most familiar of the hypothetical particles are the massive vector bosons $W^{\pm}$ and Z^0 associated with the observed charged and neutral weak currents. Somewhat more obscure are the massive scalar Higgs bosons which are connected with the spontaneous breakdown of gauge symmetry.

Lee, Quigg, Thacker, PRD, 1 September 1977

Theoretical considerations[3] suggest that the Higgs-boson mass must exceed about 4 GeV/c^2, and we have recently derived a conditional upper bound[4]

$$M_H \leq M_c = (8\pi\sqrt{2}/3G_F)^{1/2} \simeq 1\ \mathrm{TeV}/c^2, \qquad (1.1)$$

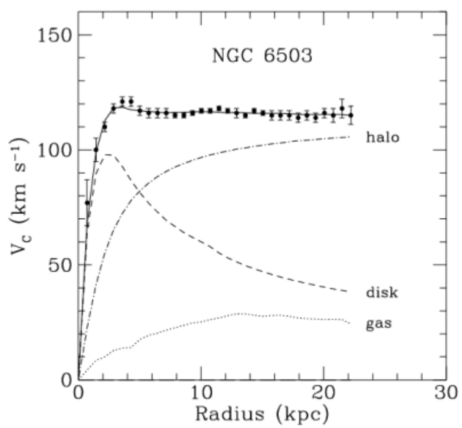
where G_F is the Fermi constant. The precise meaning of the upper bound is that if M_H exceeds the critical value M_c, weak interactions will become strong in the TeV energy regime in the sense that perturbation theory will cease to be a faithful representation of physics.

The beacon for the last 40 years

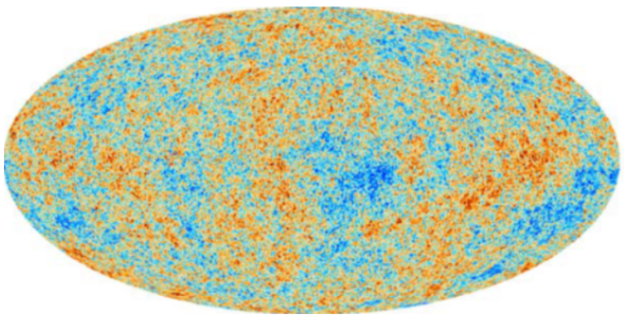
Today: answers missing

- Many evidences of physics beyond the Std Model

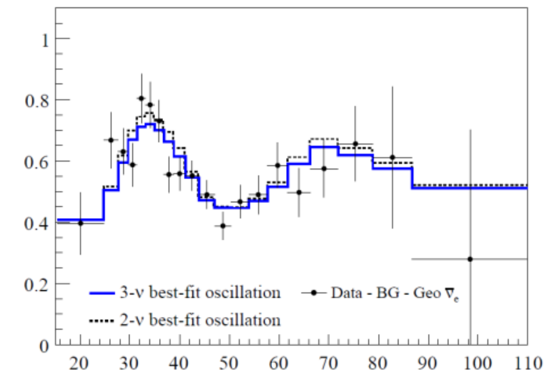
Dark Matter



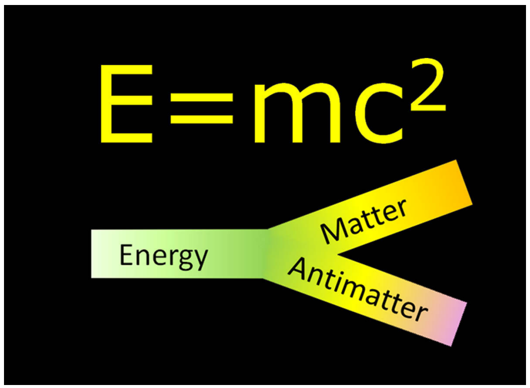
Inflation



Neutrino masses

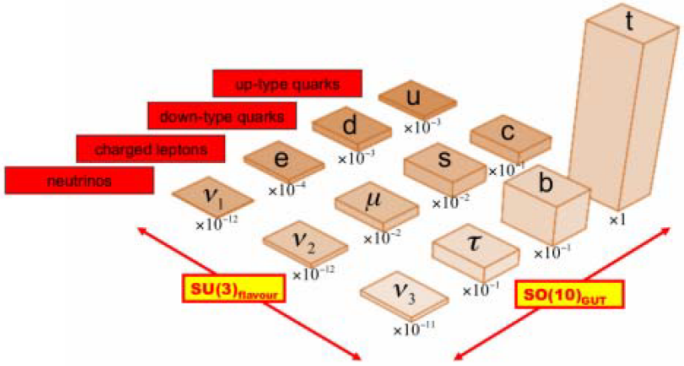


Matter-Antimatter asymmetry

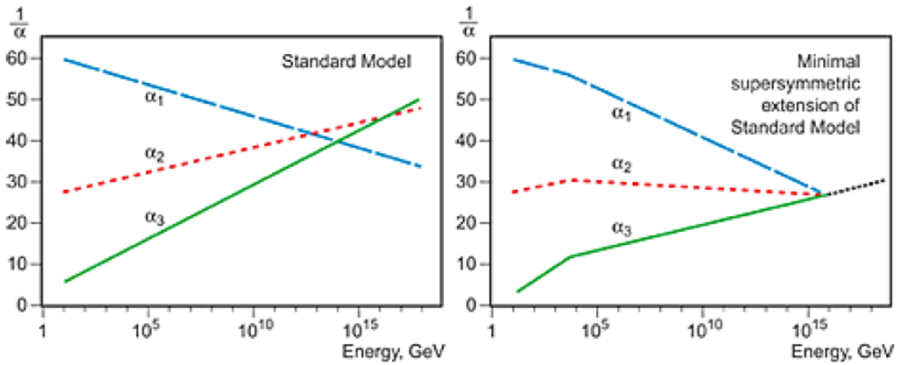


Today: SM cannot explain

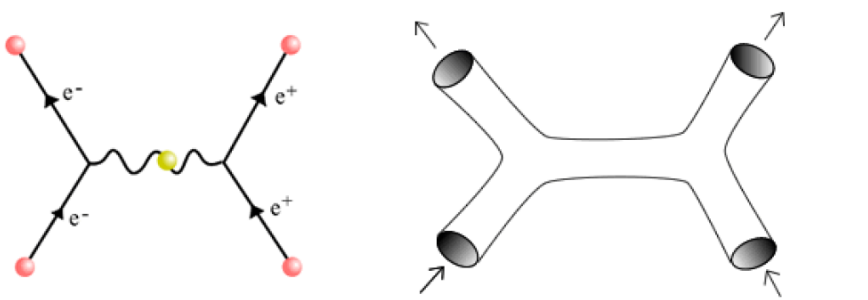
- Flavor families and mass Hierarchy



- Unification of forces



- Quantum gravity



Today: we lack a beacon

- The heuristic principle of 'Naturalness' DOES NOT WORK: the fact that LHC has not yet found any BSM evidence is EVIDENCE of fine tuning at sub permill level….

- … so the 'scale' of new physics (Λ_{NP})is not at E-weak scale … but it is hard to think that the unanswered questions and the internal difficulties of the SM do not imply some new interactions/state of matter

- Lacking a 'guiding light' it is now up to the experiments to trace the route !

Relations between theory and experiment (as seen by theorists)

A defendable picture when you have very tight predictions: e.g. BEH boson, rare decays rate

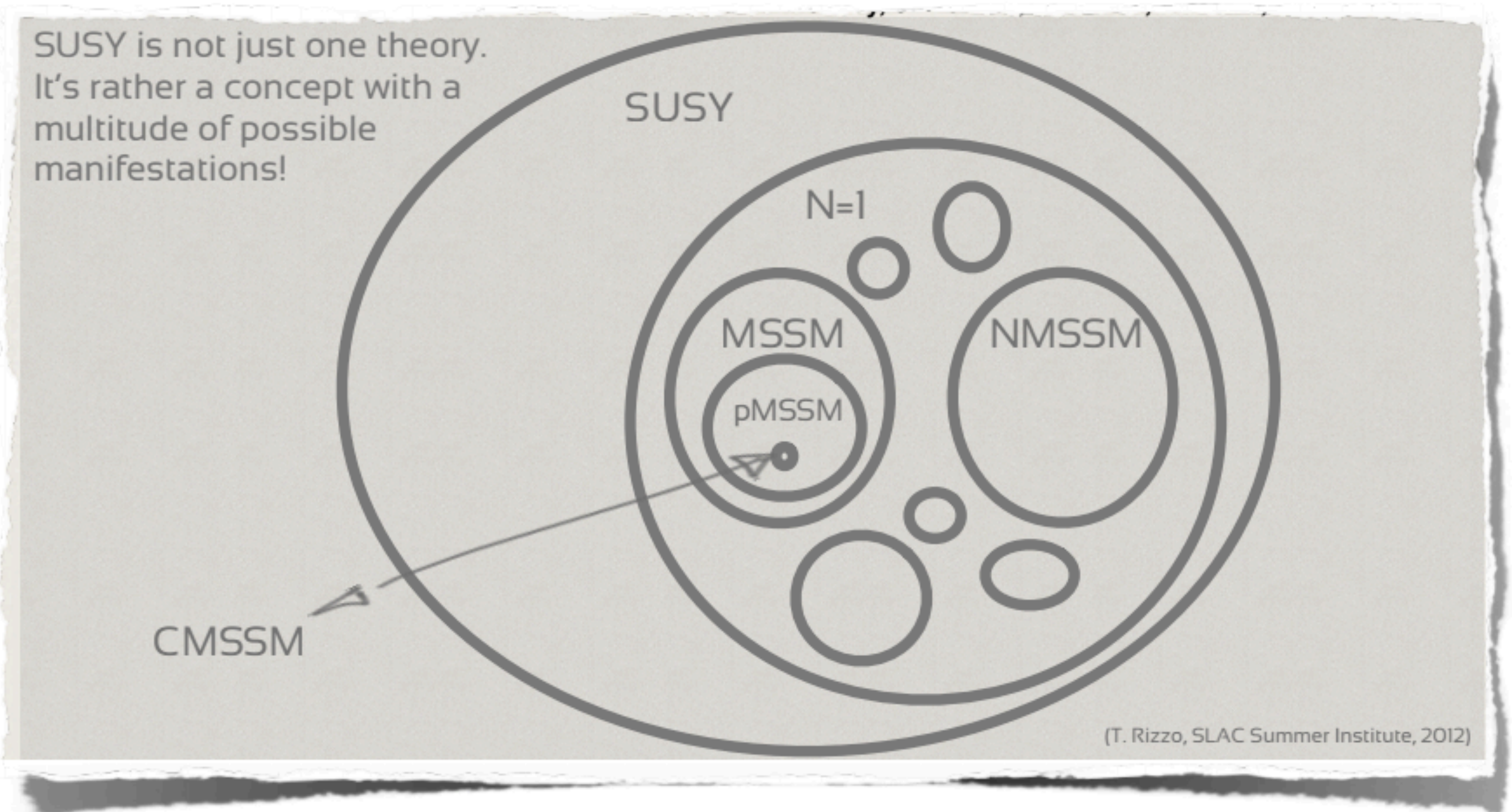
Courtesy of H. Yamamoto

..as seen by experimentalists

Beyond the Std Model theories

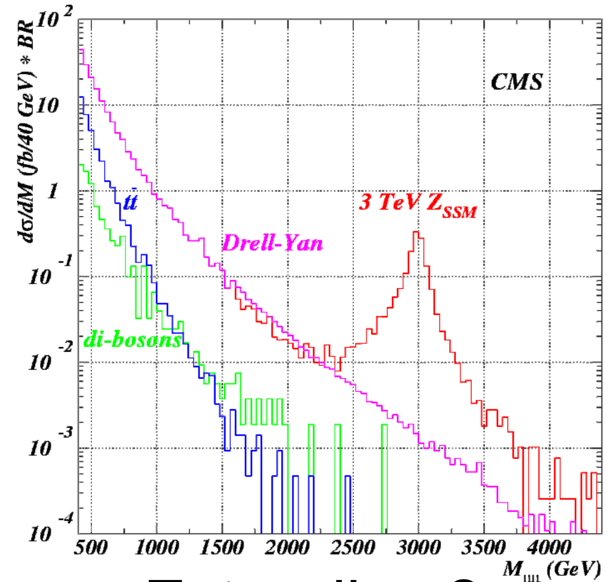
Is SUSY still worth seeking?

SUSY 'space' has been excluded only for the easiest part so far (CMSSM… still a lot to explore..but it might not be easy)



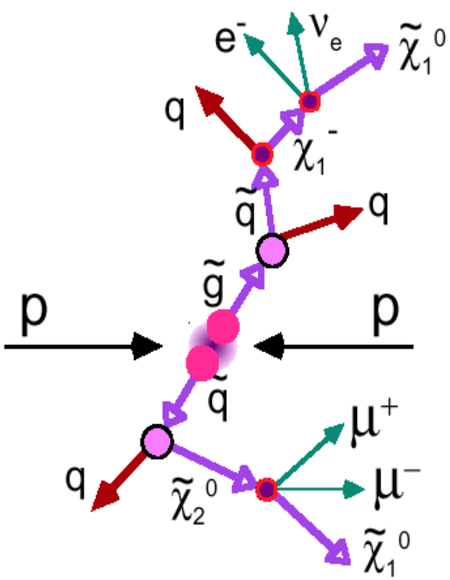
2015 and RUN2

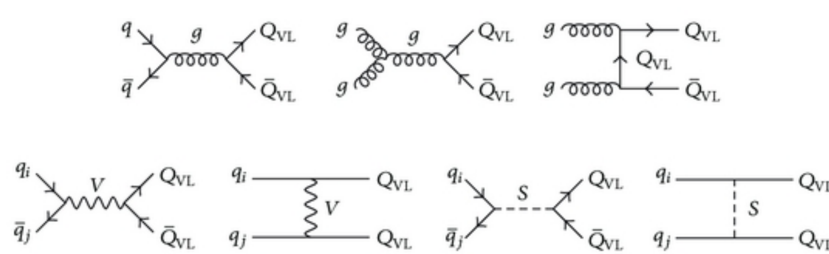
A lot at stake: we have to be ready!
We know that there must be more than the new scalar boson…and that in order to make things simple the new physics should be around the corner: we have to make sure that we do not miss the corner!



Extra dims?

Susy?





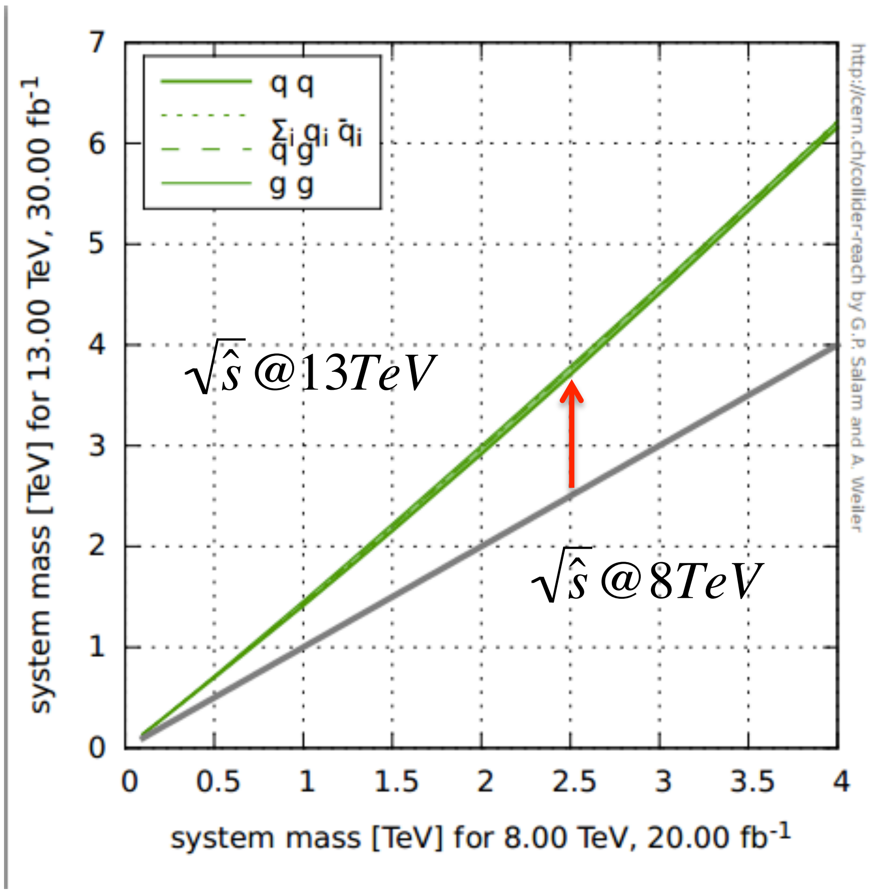
Something else?

Discoveries come early

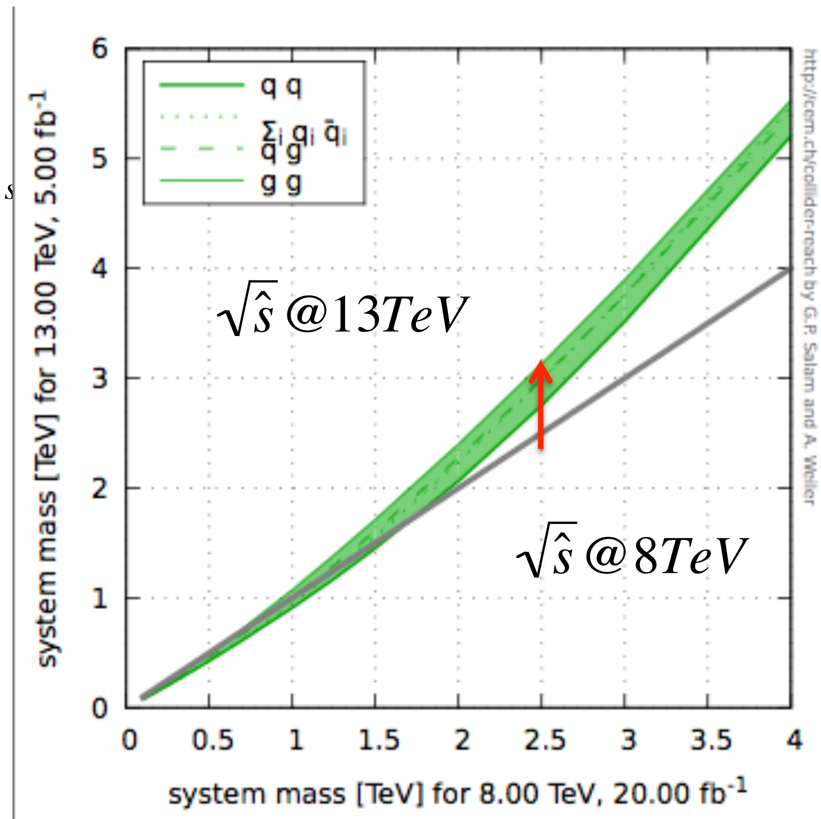
Courtesy of Gavin Salam

Discovery potential comparison 20fb^{-1} @8TeV vs X fb^{-1} @13 TeV

30 fb^{-1} @ 13 TeV



5 fb^{-1} @ 13 TeV



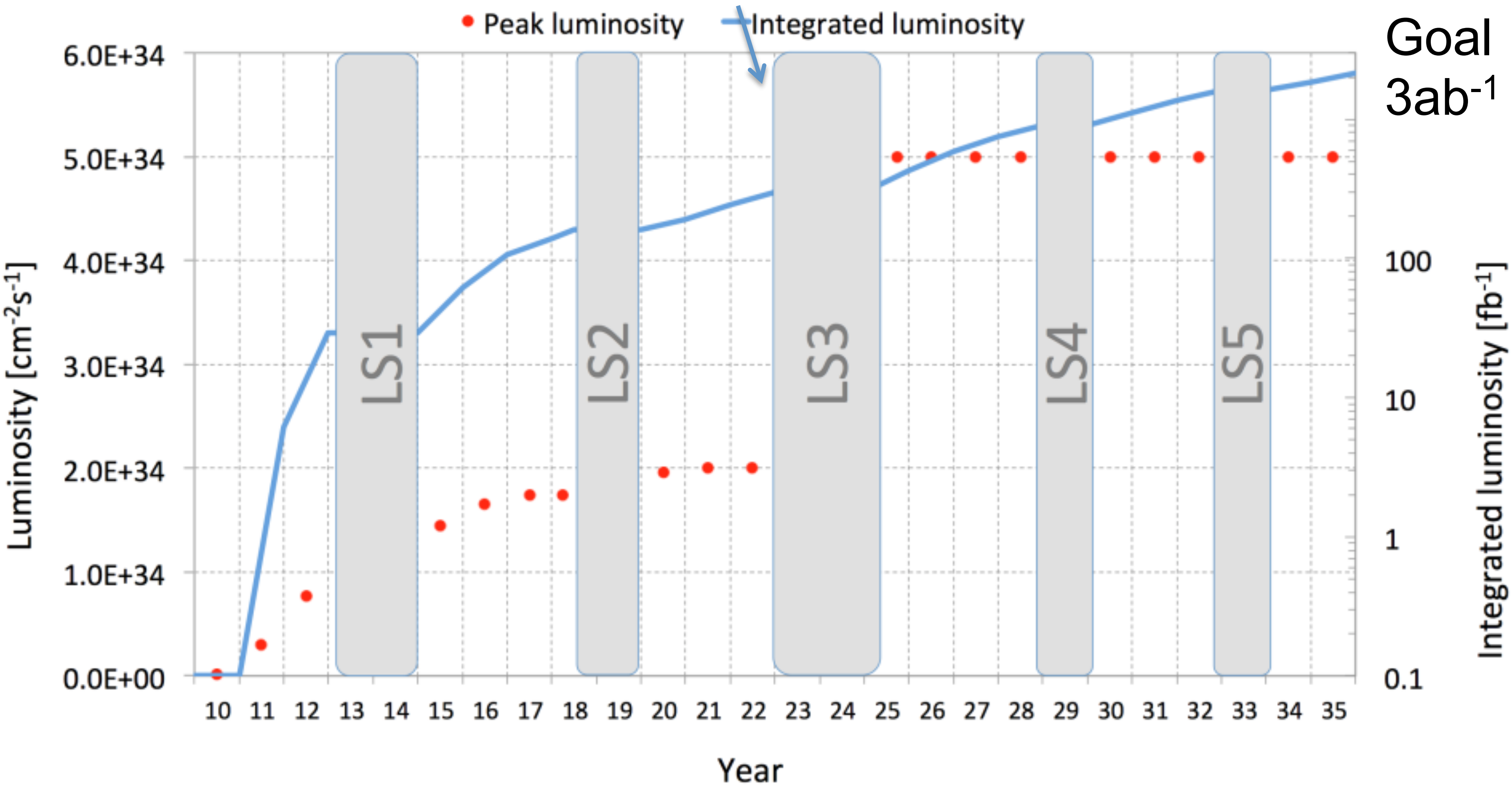
The long term future

- The CERN council has approved the extension of the LHC program beyond the initial phase (meant to deliver 300 fb^{-1} of integrated luminosity)
- The High Luminosity LHC program is now a project : a three years shutdown (LS3) will mark the start of new era with an accelerator able to to reach 10^{35} Hz/cm^2 luminosities (levelled to 5 10^{34}, possibly 7 10^{34} Hz/cm^2) and aiming to deliver in excess of 3000 fb^{-1}
- This will require major upgrades to the LHC detectors where several components will have reached their end of lifetime limits and other will have to be upgraded to exploit the higher instantaneous luminosities and radiation levels
- *US P5 :Recommendation 10*:
  Complete the LHC phase-1 upgrades and continue the strong collaboration in the LHC with the phase-2 (HL-LHC) upgrades of the accelerator and both general-purpose experiments (ATLAS and CMS).
  The LHC upgrades constitute our highest-priority near-term large project.
- **US and EUROPE are in synch regarding priorities !**

LHC goals

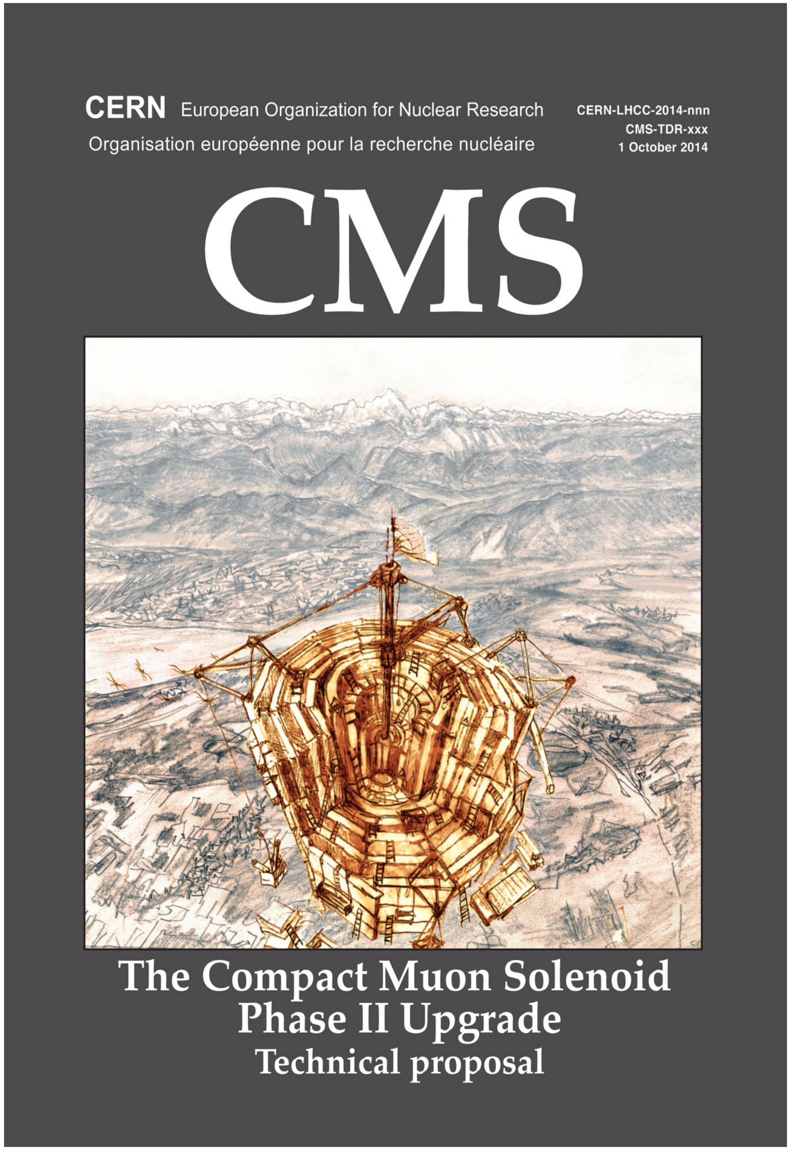
Goal 300fb^{-1}

Goal 3ab^{-1}



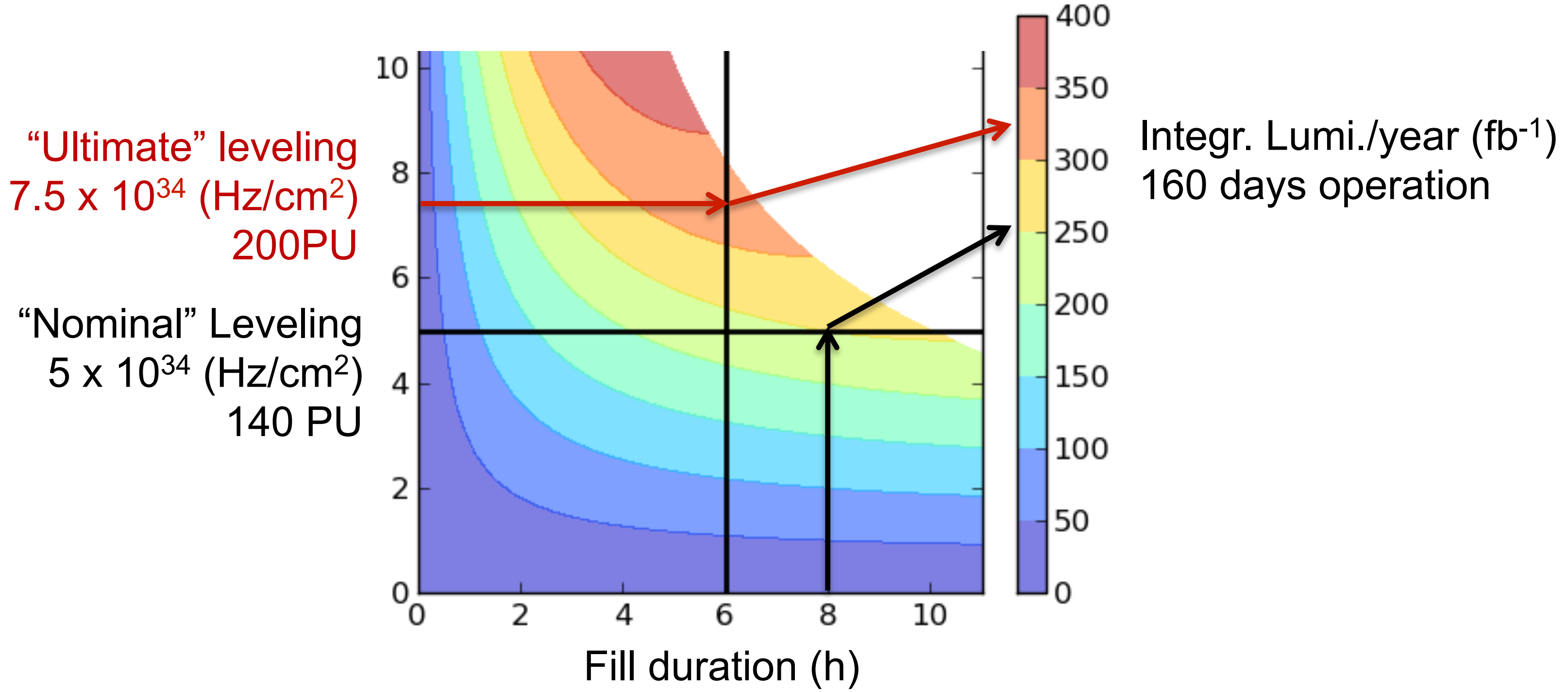
CMS Upgrades

We have just submitted our Technical Proposal for the High Luminosity LHC upgrade



Upgrade scope: the issue is Pileup

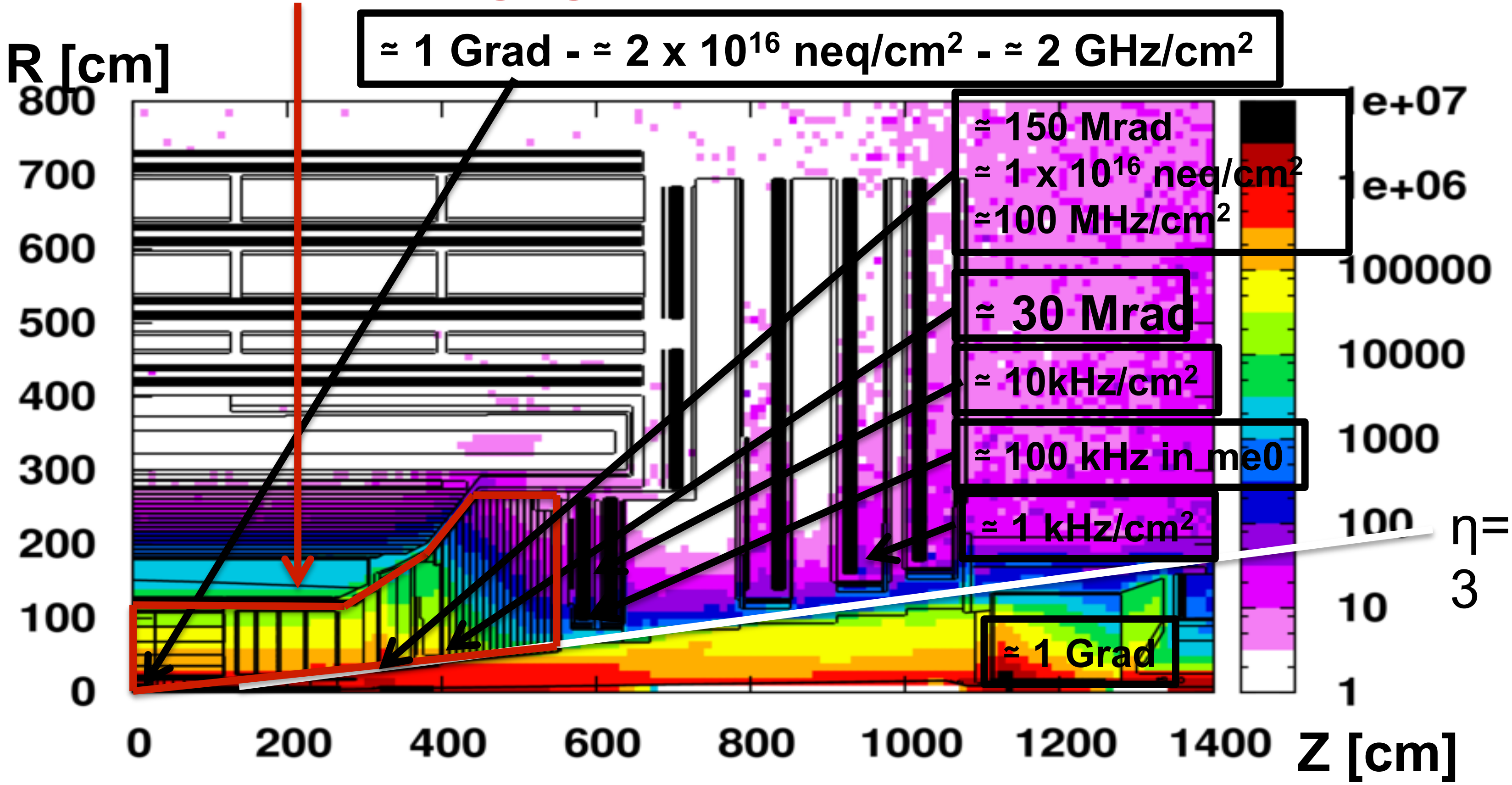
Pileup= number of concurrent pp intreactions when 2 bunches cross



CMS upgrades enable operation at 200 PileUp, maintain current performance at 140 PU, allowing moderate degradation up to 200 Pileup and radiation tolerance ≥ 3000 fb^{-1}

Radiation dose - neutron fluence - particle rates

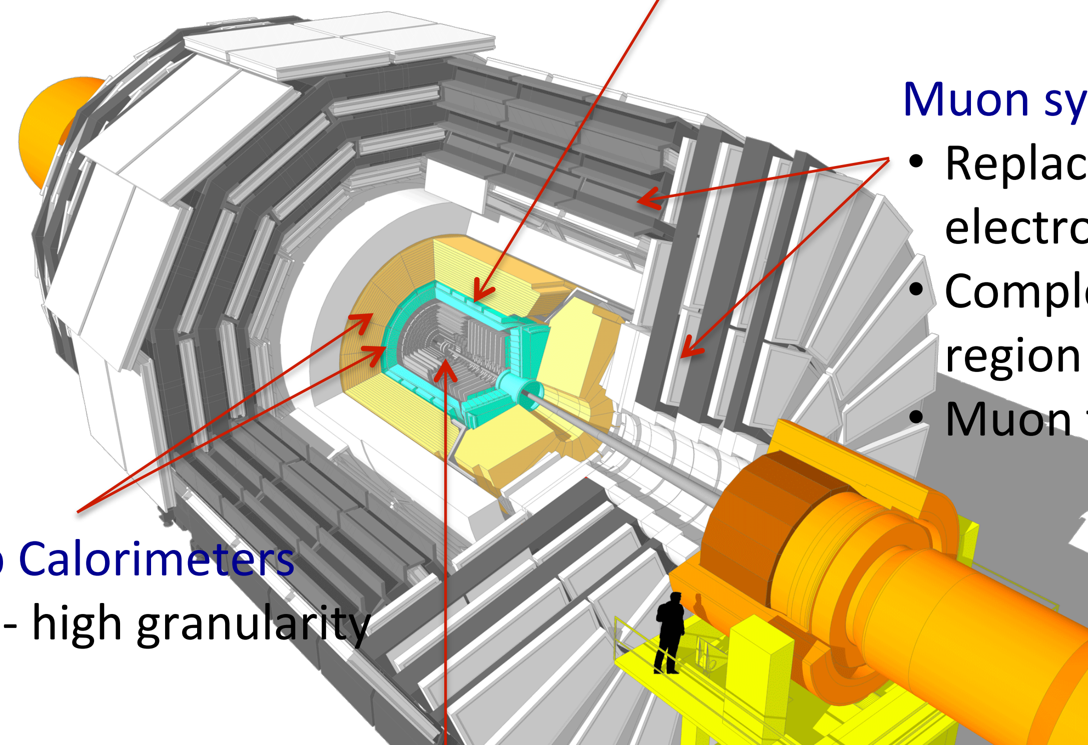
- 3000 fb^{-1} Dose map in [Gy] simulated with MARS and FLUKA
- Numbers in boxes indicate maximum doses - neutron equivalent fluence - particle rates (for 5 x 10^{34} Hz/cm^2) seen by the various detectors
- **These studies show that Tracker & End cap Calorimeters need replacement due to aging**



Summary of the CMS upgrades for Phase-II

Trigger/HLT/DAQ
- Track information at L1-Trigger
- L1-Trigger: 12.5 μs latency - output 750 kHz
- HLT output ≃7.5 kHz

Barrel EM calorimeter
- Replace FE/BE electronics
- Lower operating temperature (8∘)

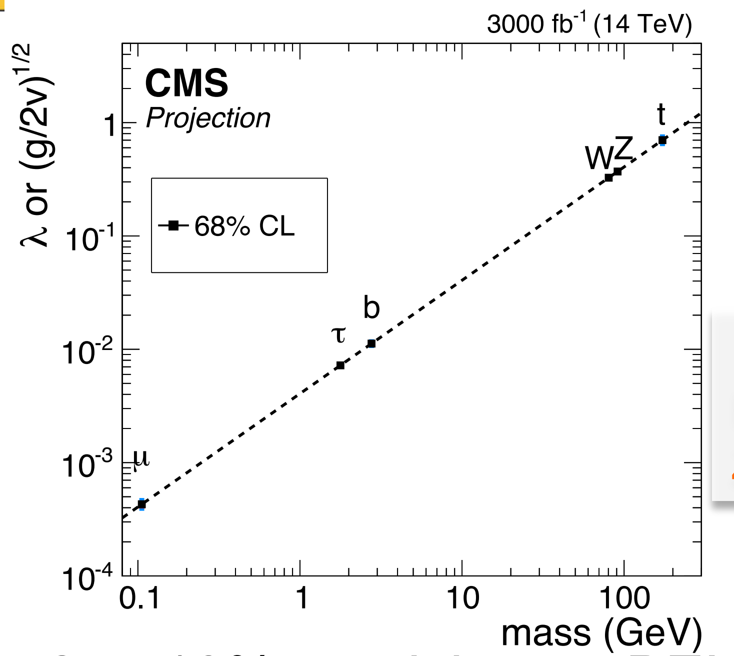


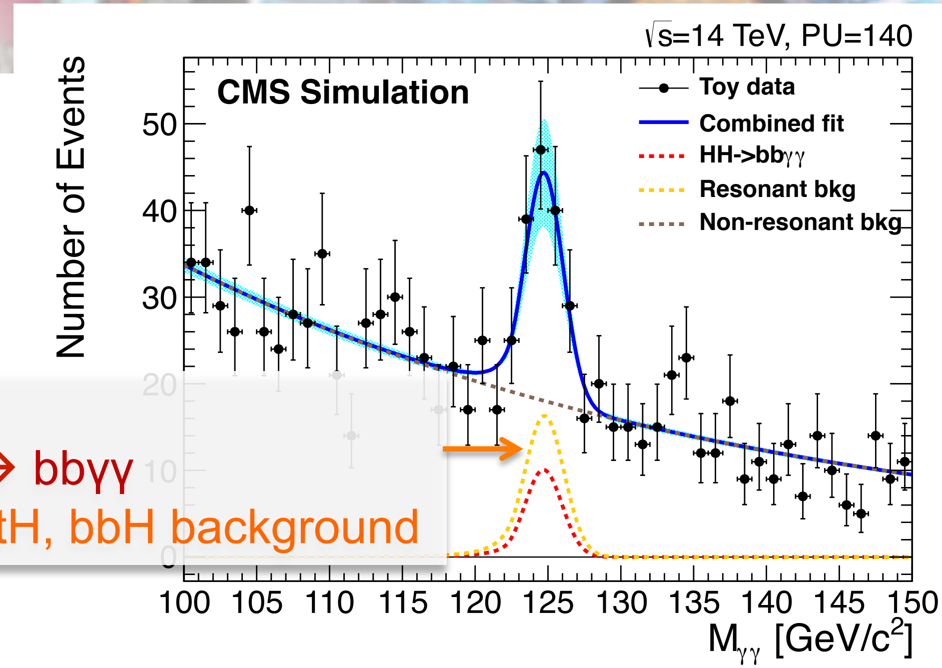
Muon systems
- Replace DT & CSC FE/BE electronics
- Complete RPC coverage in region $1.5 < \eta < 2.4$
- Muon tagging $2.4 < \eta < 3$

Replace Endcap Calorimeters
- Rad. tolerant - high granularity
- 3D capability

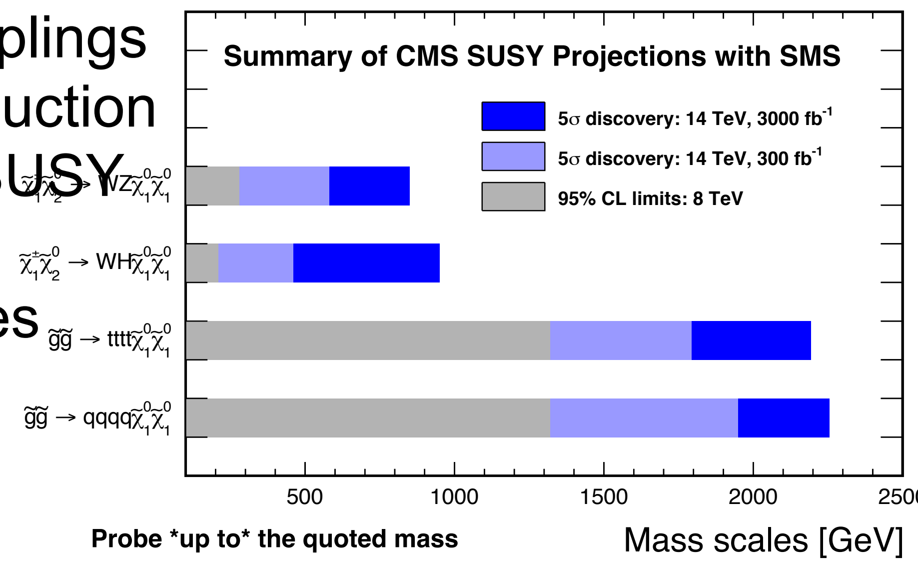
Replace Tracker
- Rad. tolerant - high granularity - significantly less material
- 40 MHz selective readout (Pt≥2 GeV) in Outer Tracker for L1-Trigger
- Extend coverage to $\eta = 3.8$

Examples of physics reach





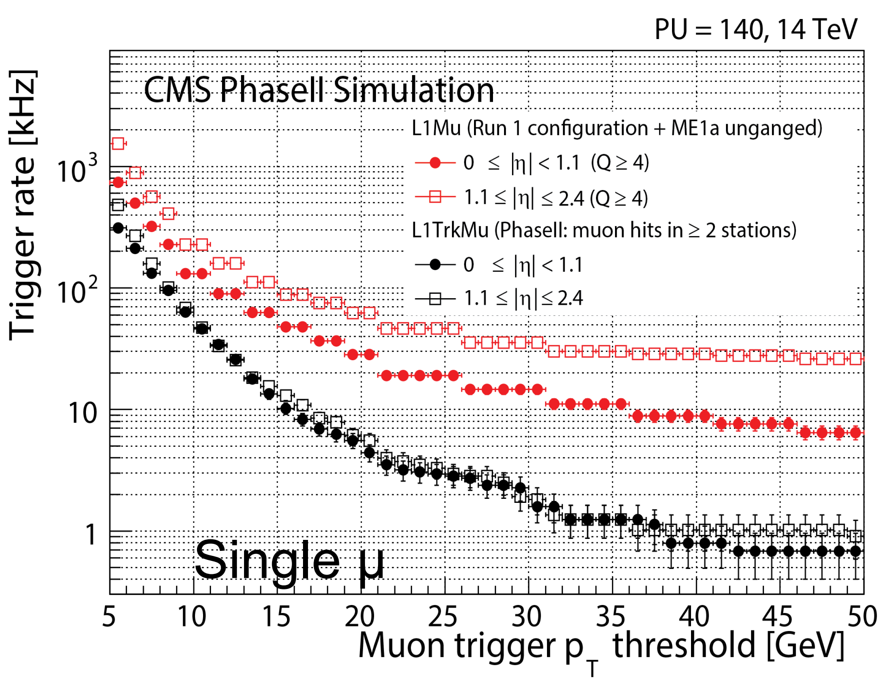
- 2 to 10% precision on BEH couplings
- Evidence of di-BEH boson production
- Access to small cross section SUSY processes
- Several other SM rare processes and BSM physics predictions
- And increased mass range

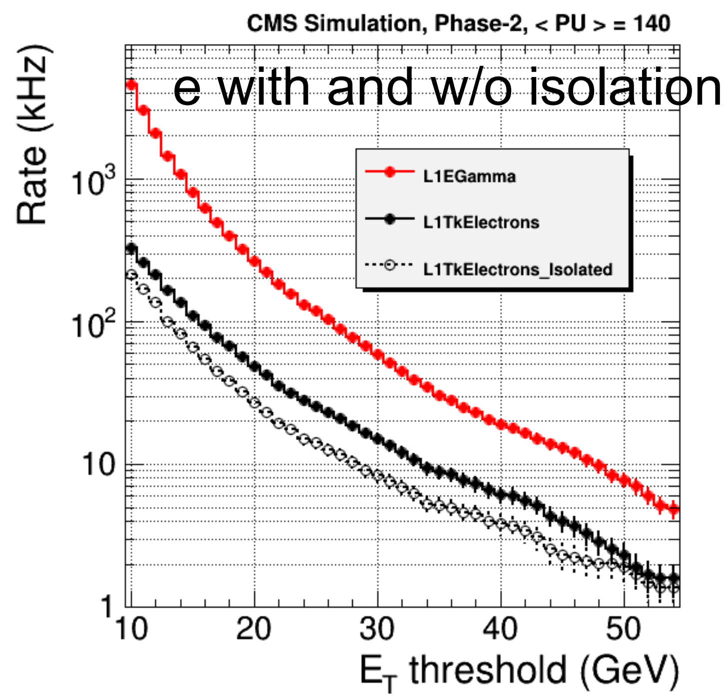


A major improvement: first level track trigger

Powerful scheme to control all inclusive trigger rates at first 40 MHz stage

- Single μ rate divided by 10
- Single e rate divided by 5(10) w/o (with) isolation
- γγ rate/5 from isolation
- τ efficiency x 2 at same rate
- Vertex ≃ 1 mm resolution → HT & MET rates divided by 10 to 100

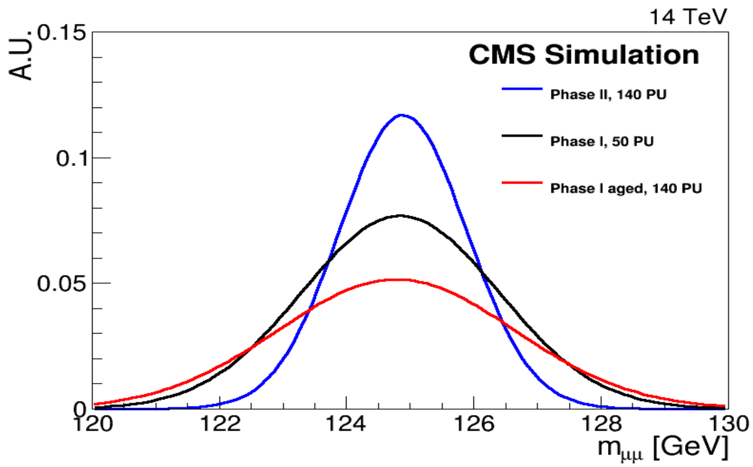


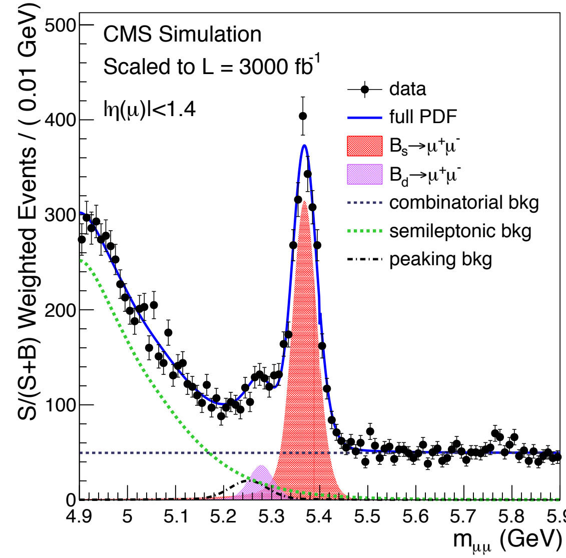


L1-Trigger studies with Phase-I menu thresholds including Track-Trigger:

- Requires ≃ 500/750 kHz rate at 140/200 PU (with 1.5 safety margin)

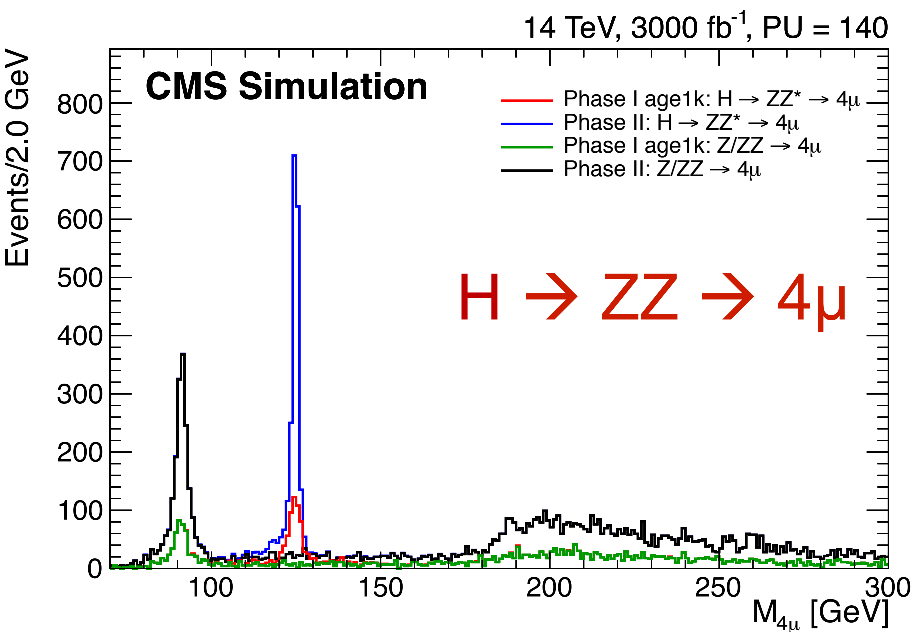
- Tracker momentum resolution

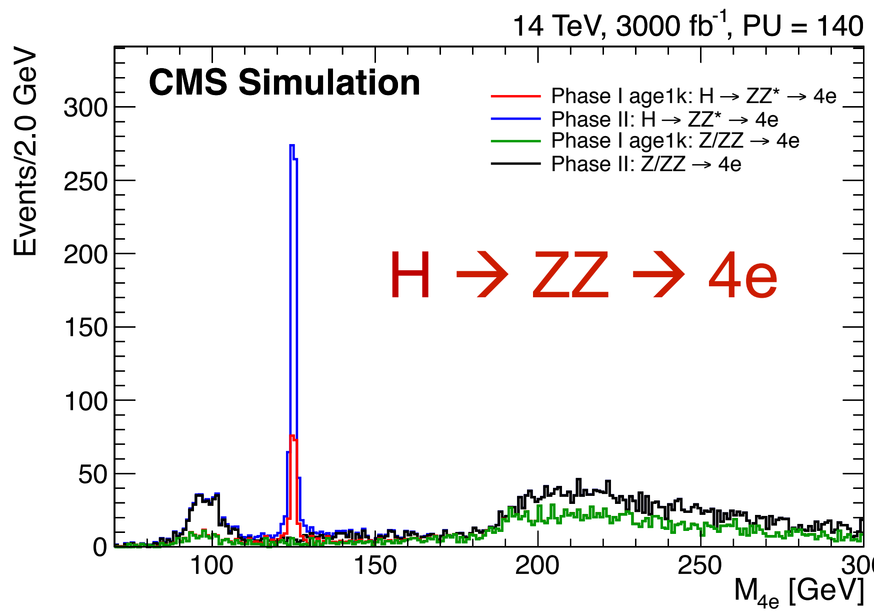




H→μμ - 20% efficiency & 40% mass resolution improvement

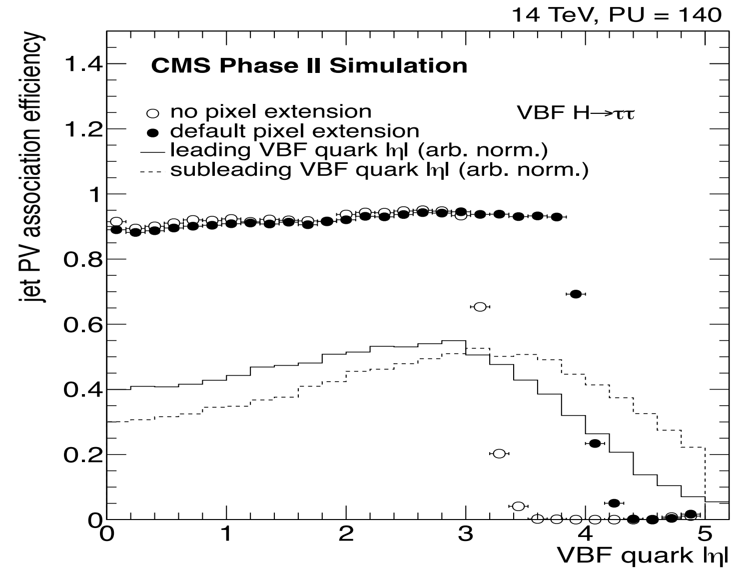
$B_d/B_s \to \mu\mu$ resolved two decay peaks measure enabled by Track-Trigger



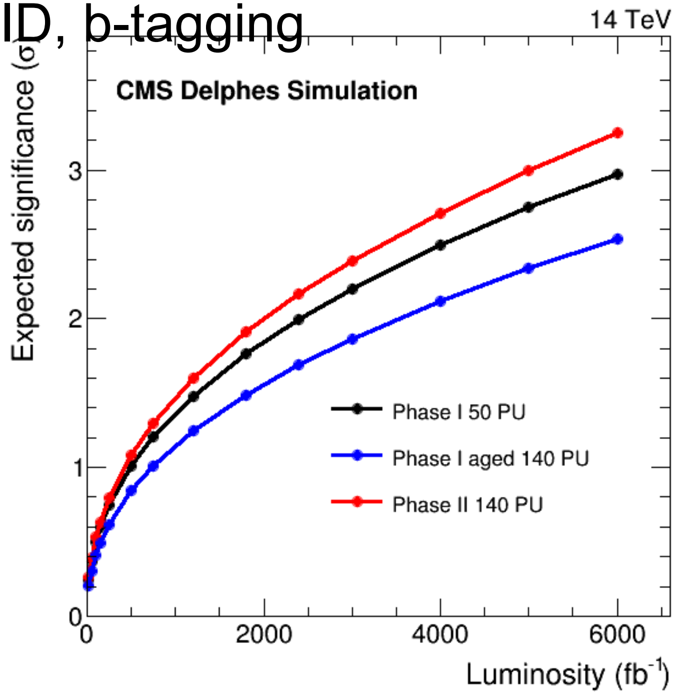


With aged Phase-I Tracker huge loss of efficiency for H→ ZZ → 4l
Upgrade restores efficiency and extensions increase the acceptance by 20%

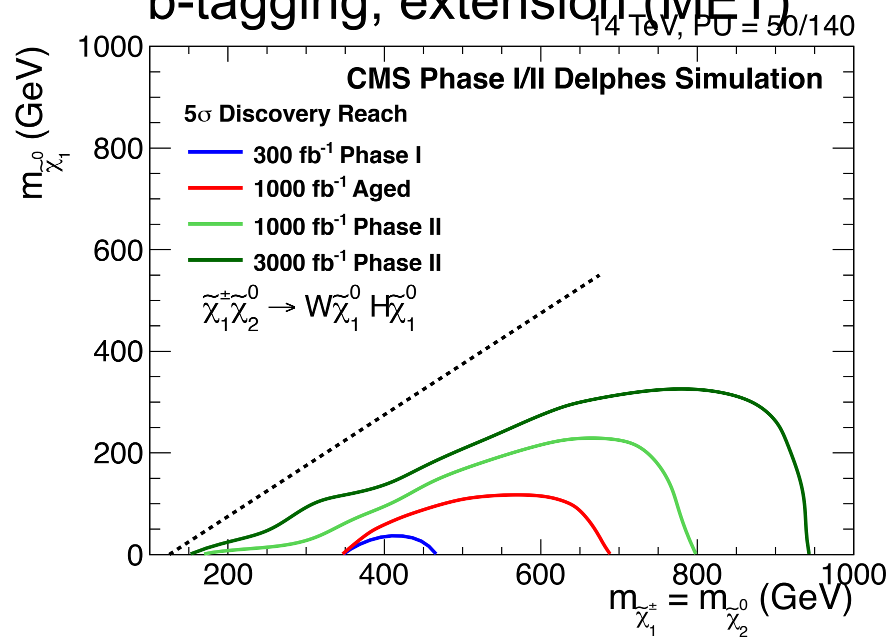
- VBF H → ττ - x 5.5 acceptance with Track-Trigger - 90% efficiency for Jet-ID with tracks and 15 % gain expected from improved mass resolution (MET)



- $W_L W_L$ scattering better significance with 140 PU Phase-II than 50 PU Phase-I - enabled by Tracker extension - EC Lepton-ID, b-tagging



- Neutralino mass range increase - enabled by Tracker b-tagging, extension (MET)



Discovery can come early …or late

Explored:

- 9 different experimental signatures.
- 5 different types of SUSY models.

Different types of SUSY models lead to different patterns of discoveries in different final states after different amounts of data

Exploring SUSY model space

Exploring experimental signature space

Analysis	Luminosity (fb^{-1})	Model				
		NM1	NM2	NM3	STC	STOC
all-hadronic (HT-MHT) search	300				$< 3\sigma$	$3-5\sigma$
	3000				$3-5\sigma$	$> 5\sigma$
all-hadronic (MT2) search	300	$3-5\sigma$	$> 5\sigma$	$> 5\sigma$		
	3000	$> 5\sigma$	$> 5\sigma$	$> 5\sigma$		
all-hadronic $\tilde{b}_1$ search	300	$< 3\sigma$	$< 3\sigma$	$< 3\sigma$	$3-5\sigma$	$< 3\sigma$
	3000	$< 3\sigma$	$< 3\sigma$	$< 3\sigma$	$> 5\sigma$	$< 3\sigma$
1-lepton $\tilde{t}_1$ search	300	$> 5\sigma$	$> 5\sigma$	$> 5\sigma$	$3-5\sigma$	$< 3\sigma$
	3000	$> 5\sigma$	$> 5\sigma$	$> 5\sigma$	$> 5\sigma$	$> 5\sigma$
monojet $\tilde{t}_1$ search	300					$3-5\sigma$
	3000					$3-5\sigma$
$m_{\ell^+\ell^-}$ kinematic edge	300	$< 3\sigma$	$< 3\sigma$	$< 3\sigma$		
	3000	$> 5\sigma$	$< 3\sigma$	$< 3\sigma$		
multilepton + b-tag search	300	$> 5\sigma$	$> 5\sigma$	$> 5\sigma$	$3-5\sigma$	
	3000	$> 5\sigma$	$> 5\sigma$	$> 5\sigma$	$> 5\sigma$	
multilepton search	300	$< 3\sigma$	$< 3\sigma$	$< 3\sigma$	$< 3\sigma$	
	3000	$3-5\sigma$	$3-5\sigma$	$< 3\sigma$	$3-5\sigma$	
ewkino WH search	300		$< 3\sigma$			
	3000		$3-5\sigma$			

$< 3\sigma$ | $3-5\sigma$ | $> 5\sigma$

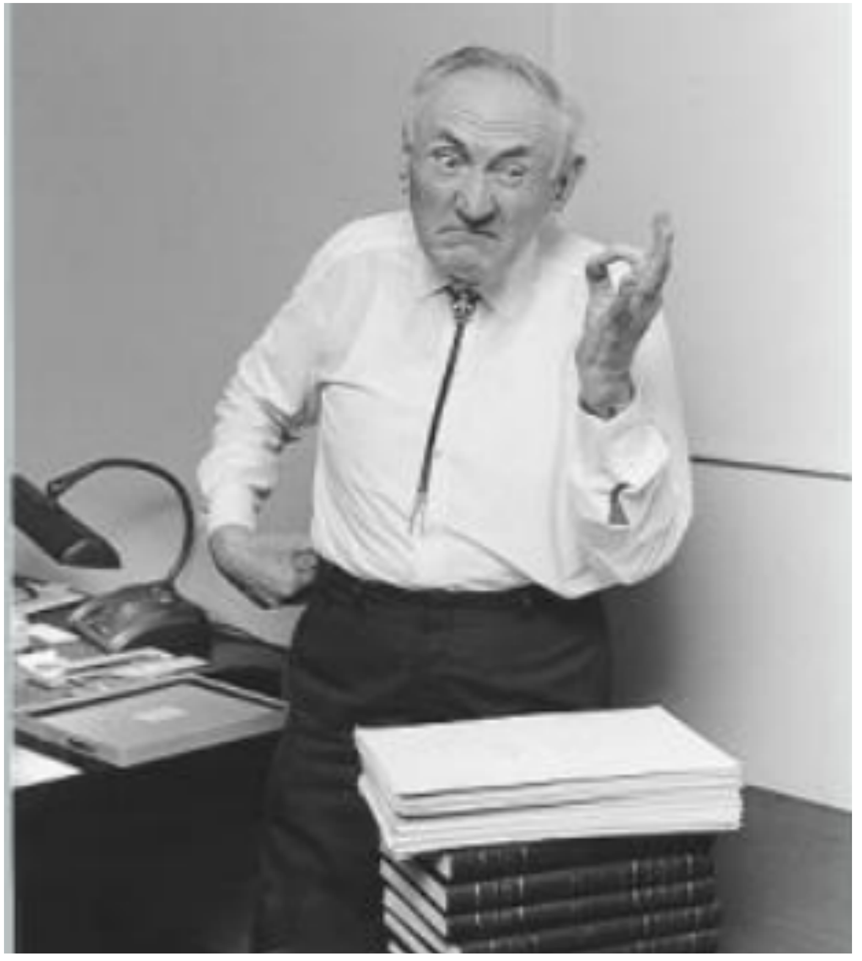
A primary Energy Frontier objective for the future: Dark matter

Fritz Zwicky : 1933

Galaxies in the Coma cluster were moving too rapidly.

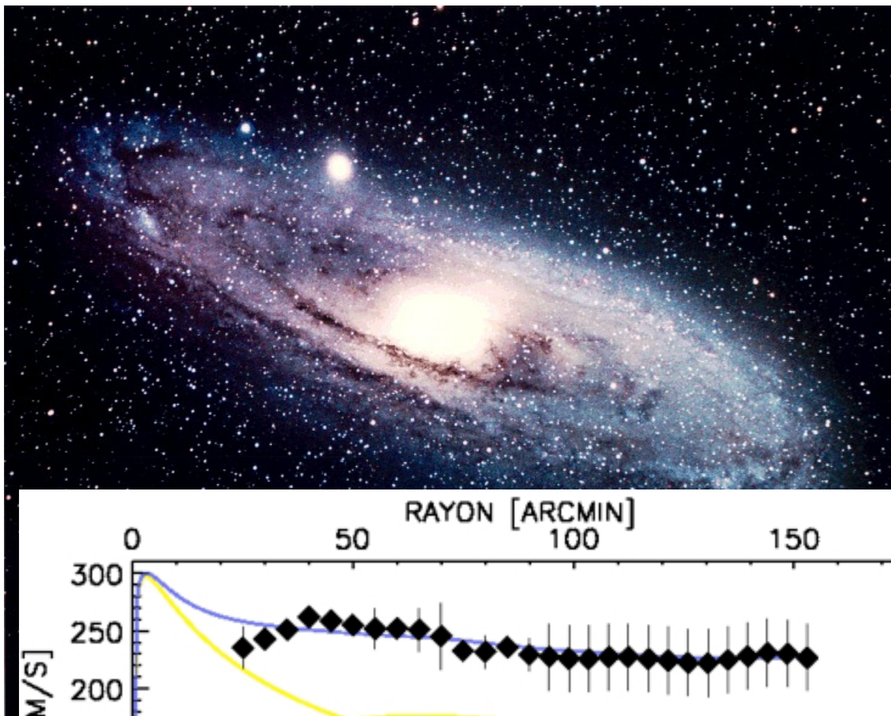
He proposed "Dunkle Materie" as the explanation.

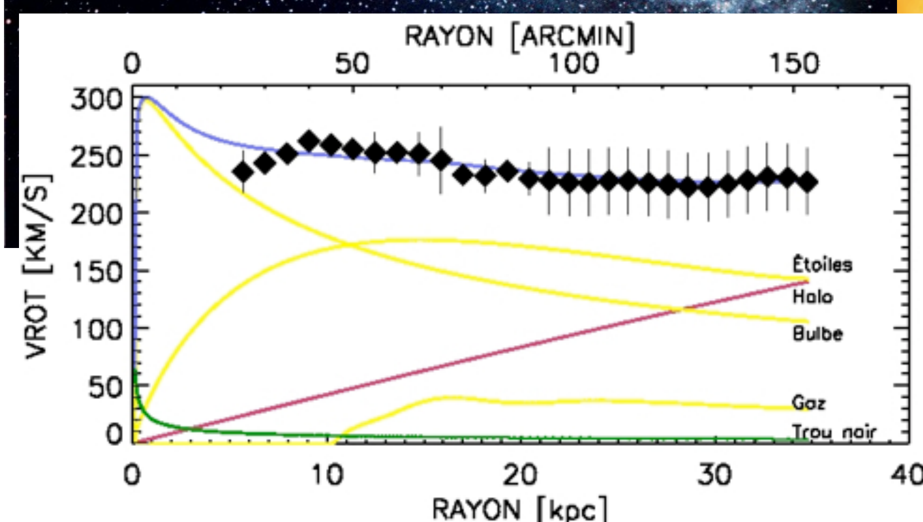
The beginning of the Dark Matter mistery



Vera Rubin (1970s)

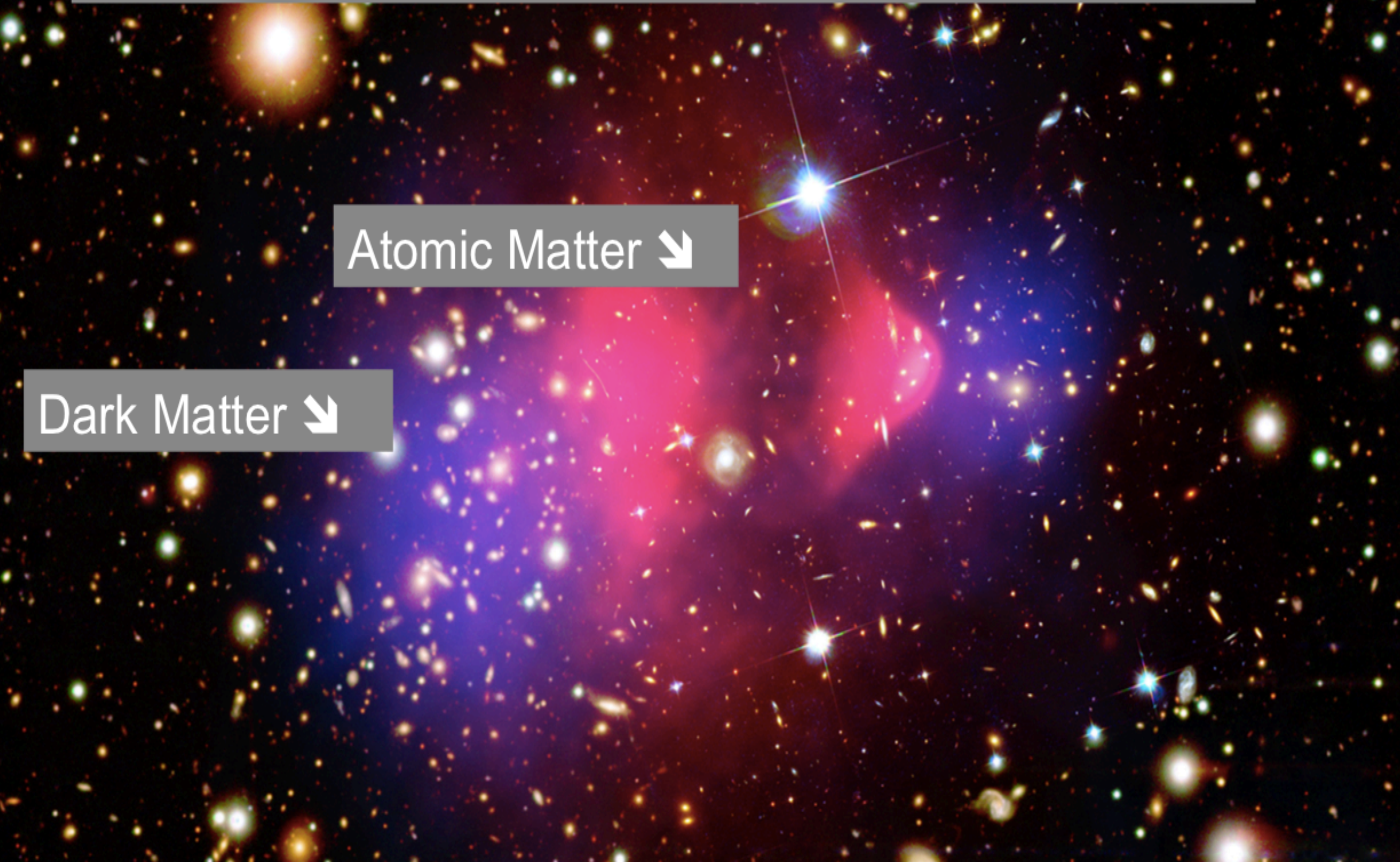
Rotation Curves of Galaxies are Flat





The Bullet Cluster:

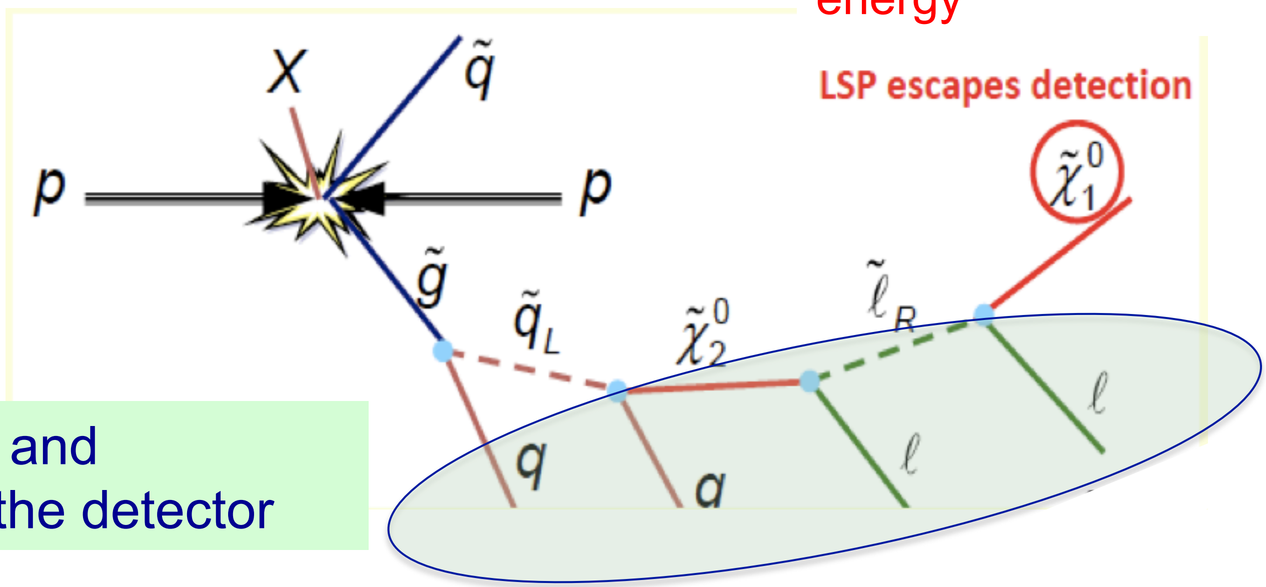
Two merging clusters: dark matter passes through while atoms get stuck



How could we ‘detect’ Dark Matter particles

- Direct way is by detecting anomalous Missing energy from final states

Dark Particle which carries ‘away’ invisible energy



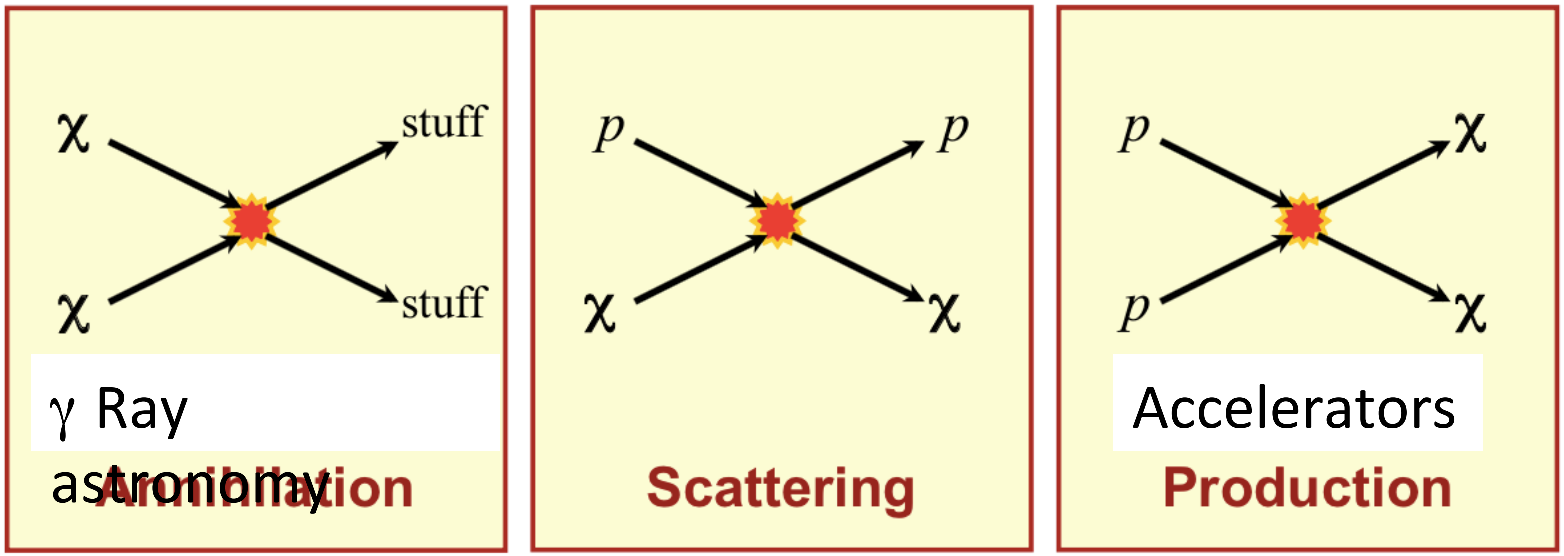
Particles visible and measurable in the detector

Will not be easy…
but we have built our detectors to be able to do it !

Dark matter: many ways to skin it

THREE PRONGED APPROACH TO WIMP DETECTION

Interactions with Standard Model particles



Indirect Detection:
Halo (cosmic-rays), capture in Sun (ν's)

Direct Detection:
Look for scattering events in detector

Accelerators:
LHC

Conclusion

- CMS is ready to continue the exciting journey into the mysteries of Nature
- We are all excited by exploring the new energy frontier
- We are preparing our long term future: CMS has been conceived as an Energy frontier experiment, but has demonstrated that with adequate detector a proton collider is also a precision measurement tool
- Stay posted for exciting times ahead !

Backup

Dark matter : reachable at LHC?

- **Weakly Interacting Massive Particles are the best motivated dark matter candidates**, Annihilation rate in the early universe determines the density today.
- The annihilation rate comes purely from particle physics and automatically gives the right answer for the relic density!

$$\Omega_X h^2 = \frac{3 \cdot 10^{-27}\, cm^3/s}{\langle \sigma v \rangle_{ann}}$$

- This is the mass fraction of WIMPs today, and gives the right answer (23%) if the dark matter is weakly interacting and the **WIMP mass in the range GeV – 10 TeV**